

A PROJECT REPORT ON

ALGORITHMS FOR PHASED ARRAY BEAM GENERATION

**PUJA VACATION PROJECT CUM TRAINING
AT GMRT, KHODAD , NCRA , TIFR**

BY

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CERTIFICATE

This is to certify that Miss Swagata Chakrabarti and Miss Somprova Nandy students of Institute of Radio Physics & Electronics, C.U. have carried out the project entitled “ALGORITHM FOR PHASED ARRAY BEAM GENERATION” satisfactorily under my guidance during the puja vacation 2006 and submitted the project report as per the requirement of GMRT.

**Project Guide.
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Our cordial thanks to all who have contributed directly or indirectly towards the completion of our assignment.

Somprova Nandy
Swagata Chakrabarti

ABSTRACT

We are to develop an algorithm that will generate phased array beams for any symmetric or unsymmetric antenna array configuration. The array factor derived theoretically is used to find the position of nulls and peaks of the power pattern.

The linear array is to be phased for broadside and end fire case and the Beam Width at First Null obtained experimentally from graph and derived from theory are compared. The match of the two will exhibit the proper working of the algorithm.

Next, the aspect of multiple beam generation over a single dimension or over a 2-D sky plane will be developed. The study is broadly done over the GMRT central square and the arm antenna.

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GMRT

A brief introduction to GMRT

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2. PHASING OF BEAMS OF A LINEAR ARRAY

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- I. Phasing of beams and calculation of BWFN for linear – X array.
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Subtopics:

- I. Phasing of GMRT central square at different angles

- and study of HPBW.
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- 1. List of variables
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REFERENCES

Introduction to GMRT

The GMRT consists of 30 45 m diameter parabolic dishes spread over 25 km near the village of Khodad, about 80 km north of Pune, India. The site coordinates are:

latitude = 19 deg 05 min 48 sec North,
longitude = 74 deg 03 min 00 sec East,
altitude = 588 m.

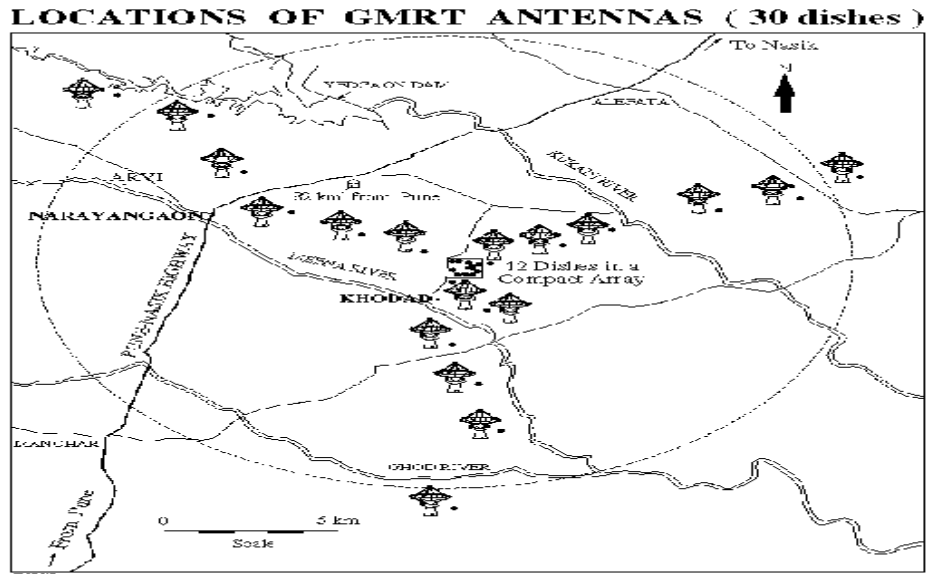
There are fourteen telescopes randomly arranged in the central square, with a further sixteen arranged in three arms of a "Y"-shaped array (similar to the [VLA](#)) giving an interferometric baseline of about 25 km. The GMRT is an [interferometer](#) which uses a technique known as [aperture synthesis](#) to make image .

Each antenna is 45 metres in diameter and, instead of a solid surface like many radio telescopes, the reflector is made of wire rope stretched between metal struts in a parabolic configuration. This works because of the long wavelengths (21 cm and longer) at which the telescope operates. Each antenna has four different receivers mounted at the focus. Each individual receiver assembly can rotate so that the user can select the [frequency](#) at which to observe.

The maximum baseline in the array gives the telescope an [angular resolution](#) (the smallest angular scale that can be distinguished) of about 1 arcsecond at the frequency of neutral [hydrogen](#) (1420 MHz).

Astronomers from all over the world regularly use this telescope to observe many different astronomical objects such as [galaxies](#), [pulsars](#) and [supernovae](#).

Why metre wavelengths : The metre wavelength part of the radio spectrum has been particularly chosen for study with GMRT because man-made radio interference is considerably lower in this part of the spectrum in India. Although there are many outstanding astrophysics problems which are best studied at metre wavelengths, there has, so far, been no large facility anywhere in the world to exploit this part of the spectrum for astrophysical research



GMRT Antenna Specifications

The large size of the parabolic dishes implies that GMRT will have over three times the collecting area of the Very Large Array (VLA) in New Mexico, USA which consists of 27 antennas of 25 m diameter and is presently the world's largest aperture synthesis telescope operating at centimetre wavelengths. At 327 MHz, GMRT will be about 8 times more sensitive than VLA because of the larger collecting area, higher efficiency of the antennas and a substantially wider usable bandwidth because of the low level of man-made radio interference in India.

Electronic Frontends and Backends : Apart from the novel low-cost design of the parabolic dishes, the instrument has state-of-the-art electronics systems developed indigenously and consisting of the following main sub units.

- Antenna feeds at six different frequency bands between 50 MHz and 1500 MHz, having good polarization characteristics as well as simultaneous multiband operation.
- Low-noise amplifiers, local oscillator synthesizers, mixers, IF amplifiers.
- Optical fibres linking the entire array with the CEB. These are used both for the telemetry signals and local oscillator phase reference communication between the CEB and each antenna base.
- A digital 2,30,000-channel FX-type correlator providing upto 128 spectral channels and covering a maximum bandwidth of 32 MHz

CHAPTER 1

INTRODUCTION TO LINEAR ARRAY

This project involves the testing and modification of an algorithm for phased array beams. The platform OS is used is LINUX and program coded in C.

Phased Array: Group of antennas in which relative phases of respective signals feeding the antennas are varied in such a way that effective radiation pattern of array is reinforced in desired direction and suppressed in undesired direction.

The program given to us gmrt_arraypattern.c was expected to calculate the array factor for any given array.

For testing the code, we started with the simple linear array case.

For that let us derive the array factor for linear array of n elements of equal amplitude and spacing.

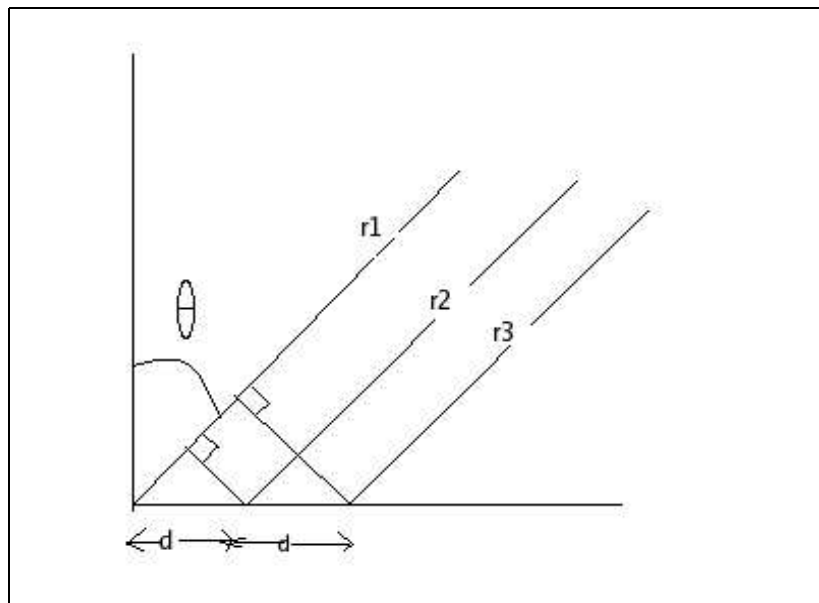


fig1.1

A uniform array is defined by uniformly spaced identical elements of equal magnitude

with linearly progressive phase from element to element.

$$\phi_1 = 0, \quad \phi_2 = \alpha, \quad \phi_3 = 2\alpha, \quad \phi_N = (N-1)\alpha$$

Inserting this linear phase progression into formula for general N element array

$$\text{Array Factor} = \text{AF} = 1 + \exp j(\alpha + kd \sin \theta) + \exp 2j(\alpha + kd \sin \theta) + \dots + \exp j(N-1)(\alpha + kd \sin \theta)$$

Since $r_1 = r$

$$r_2 = r - d \sin \theta$$

$$r_3 = r - 2d \sin \theta$$

$$\text{Therefore, AF} = \sum_{n=1}^N \exp j(N-1)\psi \text{-----(1)}$$

Now $\psi = f(\text{element spacing, phase shift, frequency, } \theta)$

If position of array is shifted so that the centre of the array is located at origin,

$$\text{AF} = [\sin(N\psi/2) / \sin(\psi/2)]$$



fig1.2

For main lobe, $\Psi = 0$

$$\text{AF (max)} = N$$

So the normalized array factor is

$$\text{AF} = [\sin(N\psi/2) / \sin(\psi/2)] / N \text{-----(2)}$$

Nulls are determined by 0 in the numerator of equation 3 while denominator $\neq 0$

Position of nulls are given by:

$$\theta_n = \sin^{-1} [(-\alpha \pm 2n\pi/N)/(kd)] \text{-----(3)}$$

If $\lambda = d$; and for broadside array i.e $\theta = 0$, $\Phi = 0$ where the array is in x axis.
In this case all elements are in phase and maxima occurs at $\theta = 0$ (when $\Psi = 0$)

$$0 = \alpha + kd \sin \theta$$

$$\implies \alpha = 0$$

$$\theta_1 = \sin^{-1} (1/N)$$

$$\theta_2 = \sin^{-1} (2/N)$$

$$\theta_3 = \sin^{-1} (3/N)$$

$$\theta_4 = \sin^{-1} (4/5)$$

If N= no. of elements=5,

$$\theta_5 = \sin^{-1} (5/N) = \pi/2$$

From experimental graph,the main lobe is repeated when $\theta = 90$.

Therefore, $\Psi = \alpha + k*d \sin \theta = kd \text{-----(4)}$

Therefore,when n=0;

$$\theta_0 = 0.$$

Therefore $\Psi = 0$.

But $\Psi = 0$. stands for main lobe . So when deriving for nulls,n =0,N,2N...are not used.

Lets test an array along X axis for **Broadside case**. To check whether the nulls occur at the value of θ calculated.

Here n=5 , d= $\lambda = 100$ m , $\theta = 0$ $\Phi = 0$

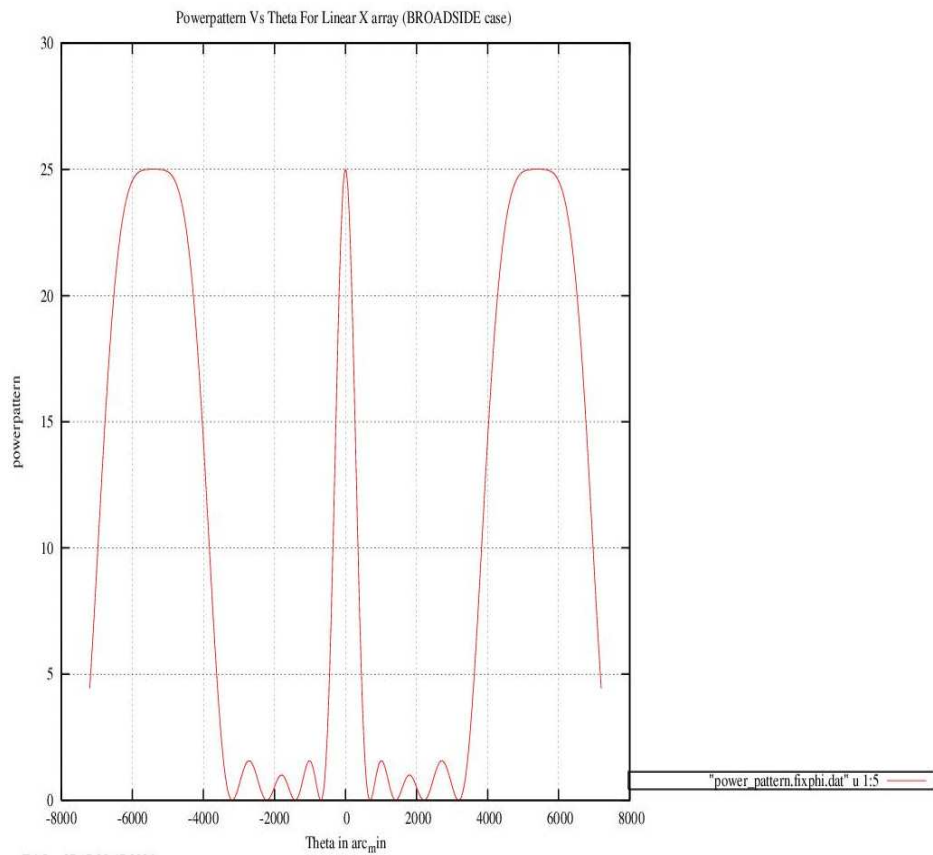


fig1.3

No of elements = N = 5

<i>BWFN</i>	<i>CALCULATED</i>	<i>THETA(deg)</i>	<i>MEASURED</i>
FIRST		23.1	23.27
SECOND		47.16	47.56
THIRD		73.74	73.42
FOURTH		106.26	106.46

For N = 5 ,

since in between two principal maxima there are 4 nulls (corresponding to n =

1,2,3,4) therefore we have three sidelobes.

So in general for N element array , there are (N-1) nulls and (N-2) sidelobes.

Peaks of array function are found by determining the zeros of the numerator term where the denominator term is simultaneously zero.

$$\text{i.e , } \sin(N \psi / 2) = 0 \text{ and } \sin(\psi / 2) = 0.$$

$$\text{When } \psi = (+ -) 2m \Pi = \alpha + kd \sin \theta_m$$

$$\theta_m = \sin^{-1}[\{ \lambda (- \alpha +- 2 m \Pi)\} / 2 \Pi d]$$

For Broadside $\alpha = 0$.

$$\text{Therefore } \theta_m = \sin^{-1}[\lambda m / d] \quad m = 0, 1, 2, \dots$$

$$\text{When } m=0 \quad \theta_m = \sin^{-1}[0] = 0$$

For broadside,

$$\psi = 0 \text{ so } m = 0 \text{ corresponds to main lobe.}$$

For $m=1$, $\lambda = d$

$$\theta_m = \sin^{-1}[1] = \Pi / 2$$

For broadside,

$$\psi = \alpha + kd \sin \theta = 0 + kd * 1 = kd \quad \text{-----(4)}$$

From (3) and (4) we get

When $n= 5$ it corresponds to same θ and ψ as $m = 1$

==> Next main lobe occurring at $m=1$ is also reflected for nulls when we take $n=$ multiples of no of elements of array.

Therefore $n \neq 0, N, 2N, \dots$

Lets test an array along X axis for **Endfire case**. To check whether the nulls occur at the value of θ calculated.

Here $N=5$, $d= \lambda = 100 \text{ m}$, $\theta = 90\text{deg}$ $\Phi = 0 \text{ deg}$.

Therefore $\alpha = -kd$.

Nulls:

$$\theta_n = \sin^{-1} [(-\alpha + 2n\pi/N)/(kd)]$$

Putting the values we get

$$\theta_1 = \sin^{-1} (-4/5)$$

$$\theta_2 = \sin^{-1} (3/5)$$

$$\theta_3 = \sin^{-1} (2/5)$$

$$\theta_4 = \sin^{-1} (1/5)$$

$$\theta_5 = \sin^{-1} (0/N)$$

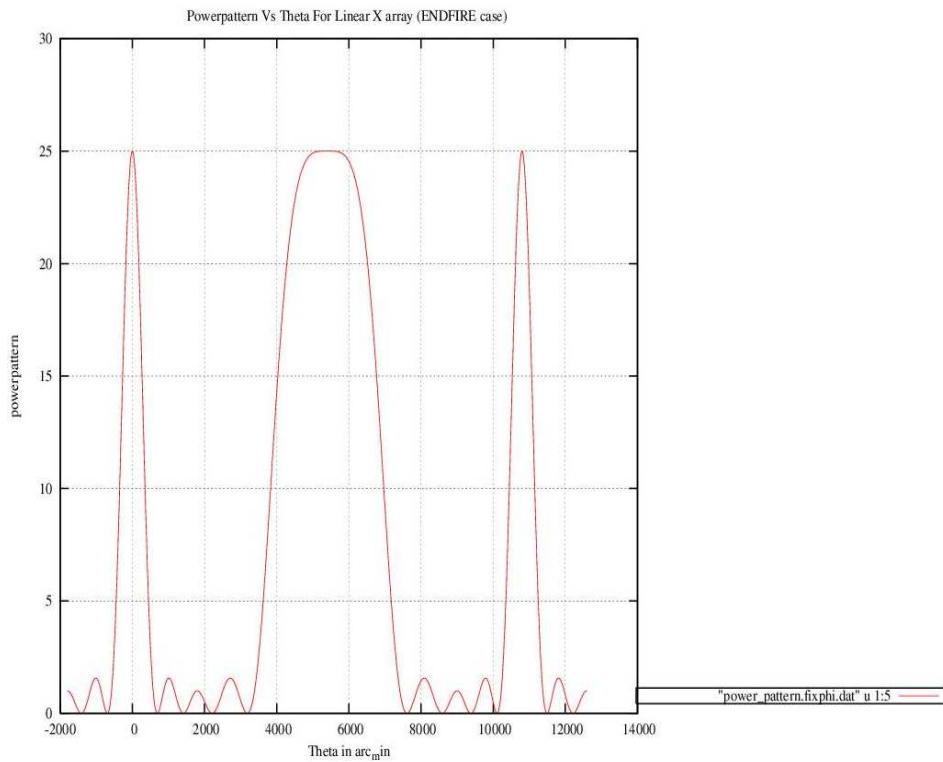


fig 1.4

Comparing the measured and the calculated value of BWFN we tabulate then as below:

No of elements = N = 5

<i>BWFN</i>	<i>CALCULATED</i>	<i>THETA(deg)</i>	<i>MEASURED</i>
FIRST		74	74
SECOND		106.28	106.39
THIRD		132.86	133.1
FOURTH		156.94	156.75

For Nulls $n \neq 0, N, 2N, \dots$

$$n = 1, 2, 3, 4$$

If $N=5$

For $n=0$;

$$\theta_0 = \sin^{-1}(-1) = -90 \text{ deg}$$

Therefore $\theta_0 = (180-90) = 90 \text{ deg}$.

$$\text{Now } \psi = \alpha + kd \sin \theta = -kd + kd \sin 90 = 0$$

==> Main lobe.

Between the main lobe and the grating lobe there are four nulls corresponding to $n = 1, 2, 3, 4$ Therefore there are three side lobes as obtained experimentally.

Peaks:

$$\theta_m = \sin^{-1} \left[\frac{\lambda (-\alpha \pm 2m \Pi)}{2 \Pi d} \right]$$

For $m=0$

$$\theta_0 = 90 \text{ deg}$$

Therefore $\psi = 0$

Hence it is the main lobe.

Therefore $n = 0$ (from null equation) and $m = 0$ (from peak equation) represent the same main lobe.

Lets test an array along X=Y line for broadside case.

Here $\theta = 0$, $\phi = 0$.

The formula for $BWFN = 2 \sin^{-1} [(-\alpha + 2n\pi/N)/(kd)]$

Here $\alpha = 0$.

gives, $= 16.26 \text{ deg (calculated)}$

But when it is measured from graph we get

$BWFN = 23.27 \text{ deg (measured)}$

As depicted in graph()

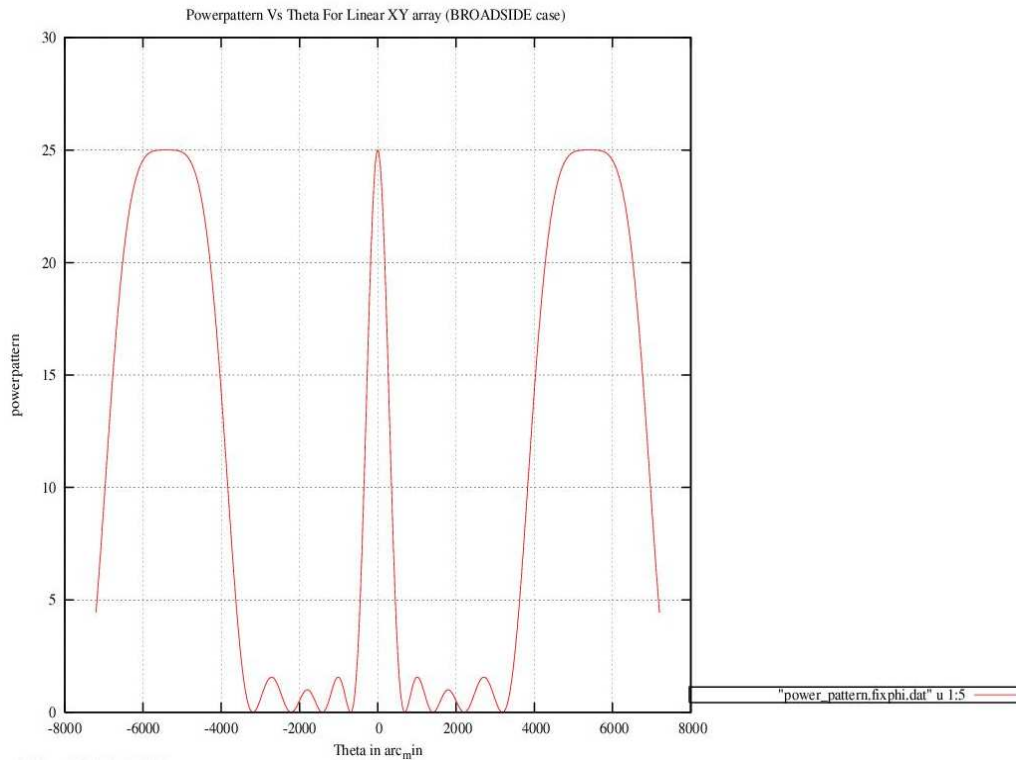


fig1.5

Hence there is a discrepancy of ratio
 = (23.27 deg (measured) / 16.26 deg (calculated))
 =1.4

This discrepancy of factor (sqrt2) is due to the fact that in deriving the formula for NULLS we assumed that there is no component due to azimuthal angle. In deriving the array factor we had inherently assumed that the plane where the field contribution due to the array points is to be found, must pass through the line containing the array points.

So to align the array points in the proper broadside for X-Y array, we must take

$$\theta = 0, \quad \Phi = 45\text{deg.}$$

Here BWFN(calculated)= 16.26 deg and (measured) = 16.37 deg.

As depicted in graph()

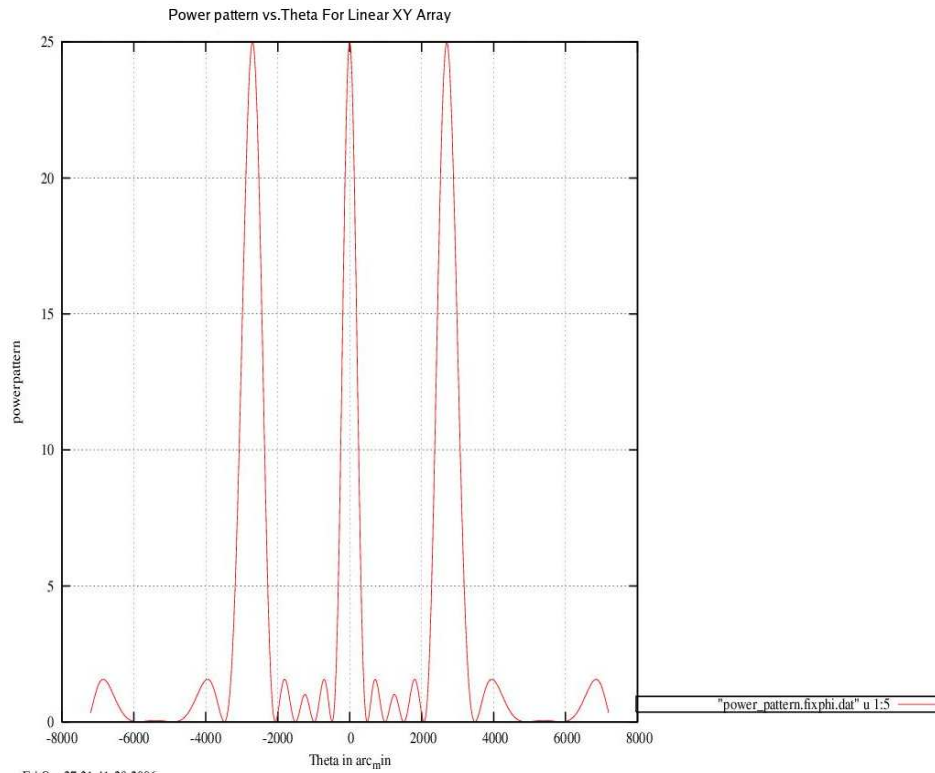


fig 1.6

If the number of array points contributing to the array factor is increased, the number of side lobes increases. Also if the array factor is not normalised then we note that the peak amplitude of the main lobe increases. It is proportional to (number of elements)². This is as shown in the graph()

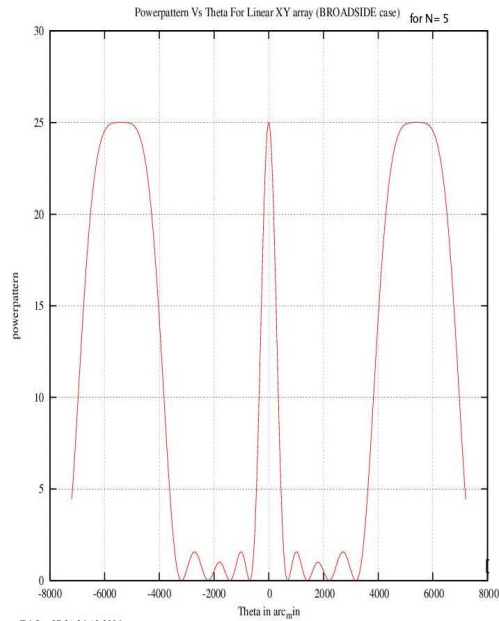


fig1.7

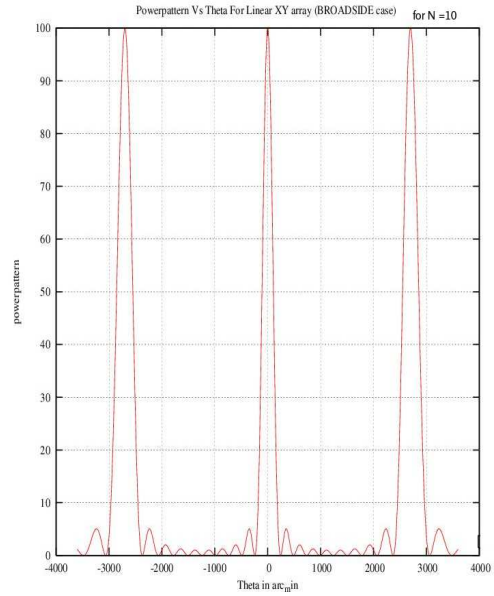


fig1.8

Observation:

For an array along X=Y line,

When $\Phi = -45\text{deg}$ i.e 135deg , Whatever be the value of θ , the power level remains constant. This is as depicted in the graph () below

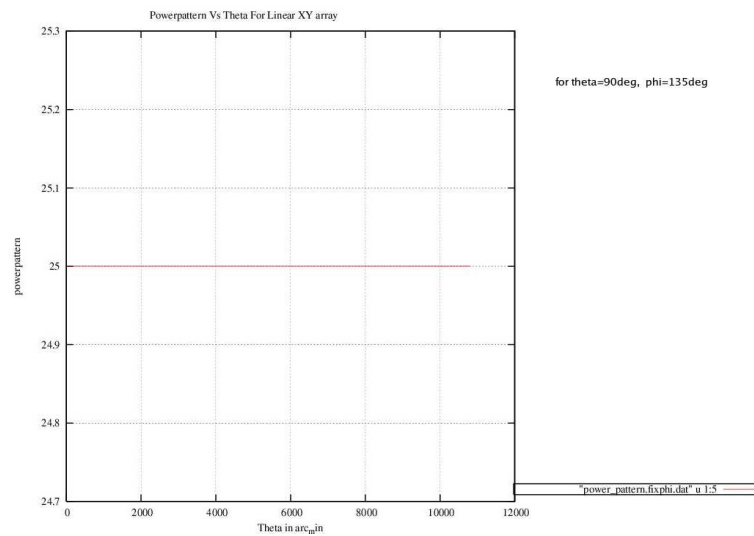


fig1.9

Broadside array produces disc shaped beams covering full 360deg in plane normal to axis of array- i.e cylindrical symmetry.

Endfire array produces cigar shaped beam which has same shape in all planes containing axis of array i.e planar symmetry.

CHAPTER 2

PHASING OF BEAMS OF A LINEAR ARRAY:

The program is then modified to incorporate two attributes.

- 1) Direction for phasing the array beam. (θ_p, ϕ_p)
- 2) Direction for calculation of the power pattern. (θ_o, ϕ_o)

This change is incorporated for practical application of the array pattern. It is because, it might be necessary to phase the beam in a required direction, while the beam pattern may be necessary to be calculated in different direction for an antenna.

This is experimented for X-Y array with 5 elements with different θ_p values, and the power pattern being calculated about that angle.

Here $\lambda = 1 \text{ m}$, $N = 5$,

From below data we note that with $\lambda = 1 \text{ m}$, BWFN of the order of arcmin is obtained

θ_p (in deg)	ϕ_p (in deg)	θ_o (in deg)	ϕ_o (in deg)	Range of θ (in deg)	Range of ϕ (in deg)	Accuracy of θ (in arcmin)	Accura cy of ϕ (in arcmin)	BWFN <i>(measur ed)(in arcmin)</i>	BWFN <i>(calculat ed)(in arcmin)</i>
0	45	0	45	0.5	1	0.5	30	9.98	9.72
45	45	45	45	0.5	1	0.5	30	13	14.4
90	45	90	45	7	1	0.5	30	366	366

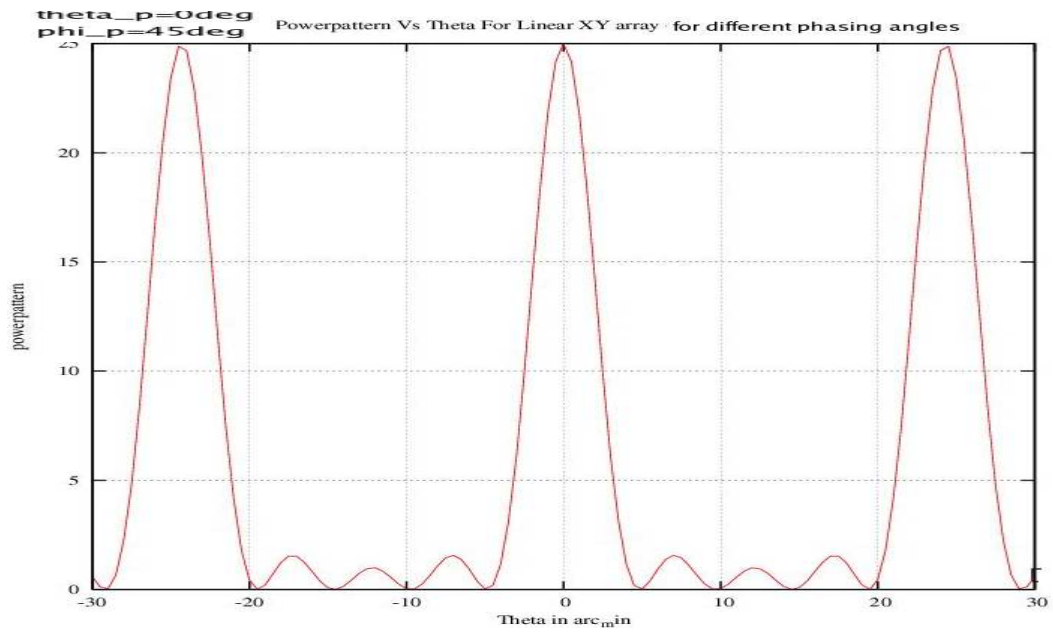


fig2.1

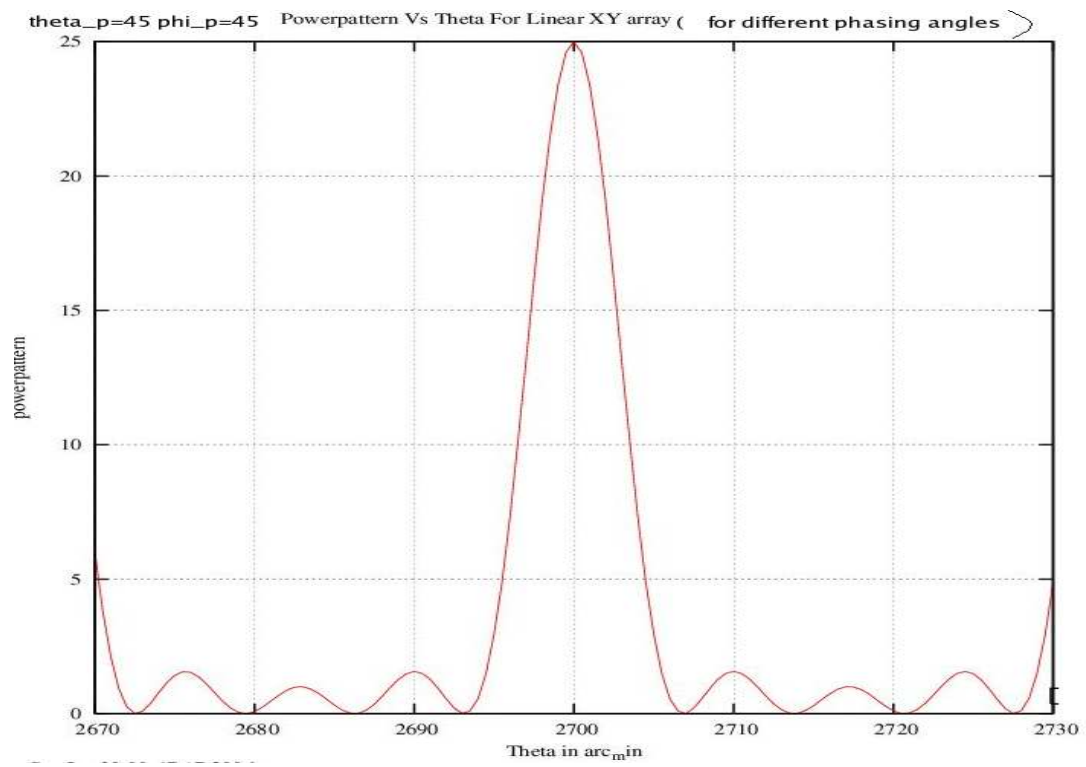


fig2.2

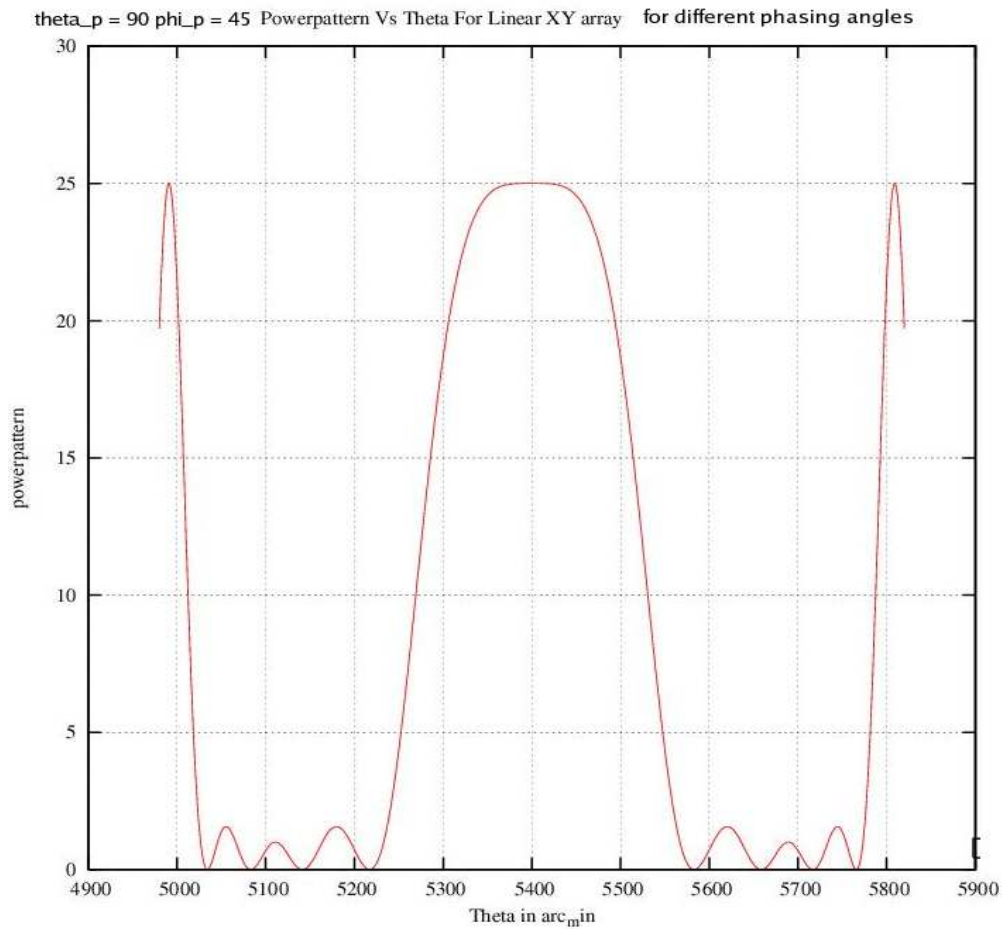
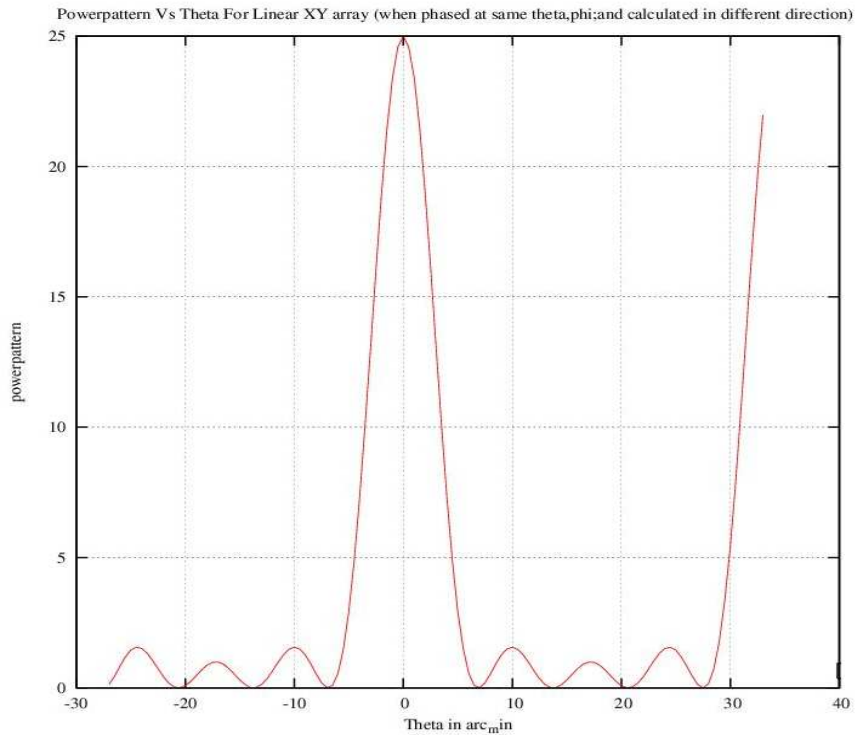


Fig 2.3

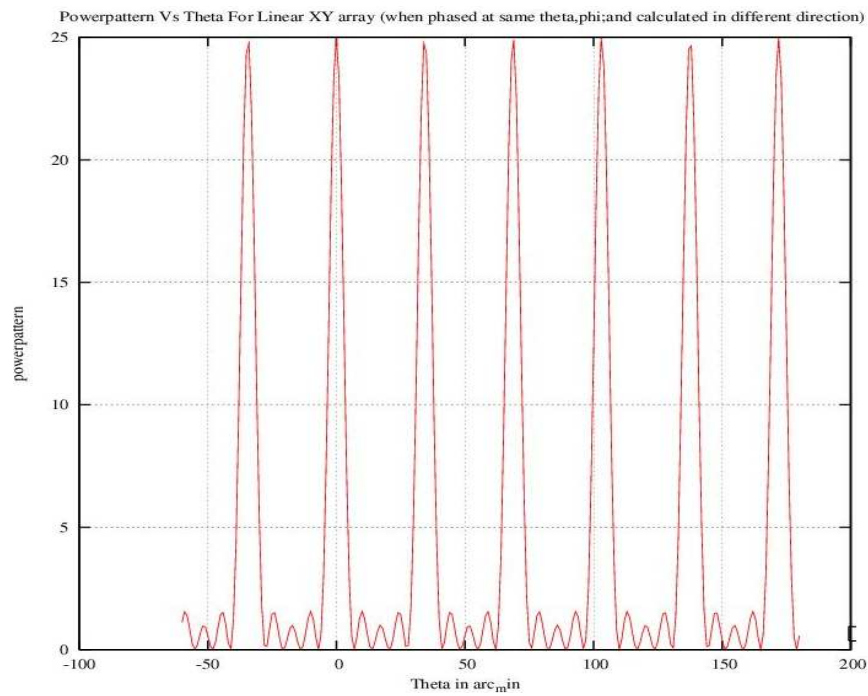
Next we keep $\theta_p = 0$, $\phi_p = 0$ fixed, and observe the nature of the power pattern by varying θ_0 , i.e the point of calculation of the power pattern is changed keeping the phasing of the beam at same theta_p



$$\theta_0 = 3 \text{ arc min}$$

$$\phi_0 = 0 \text{ deg}$$

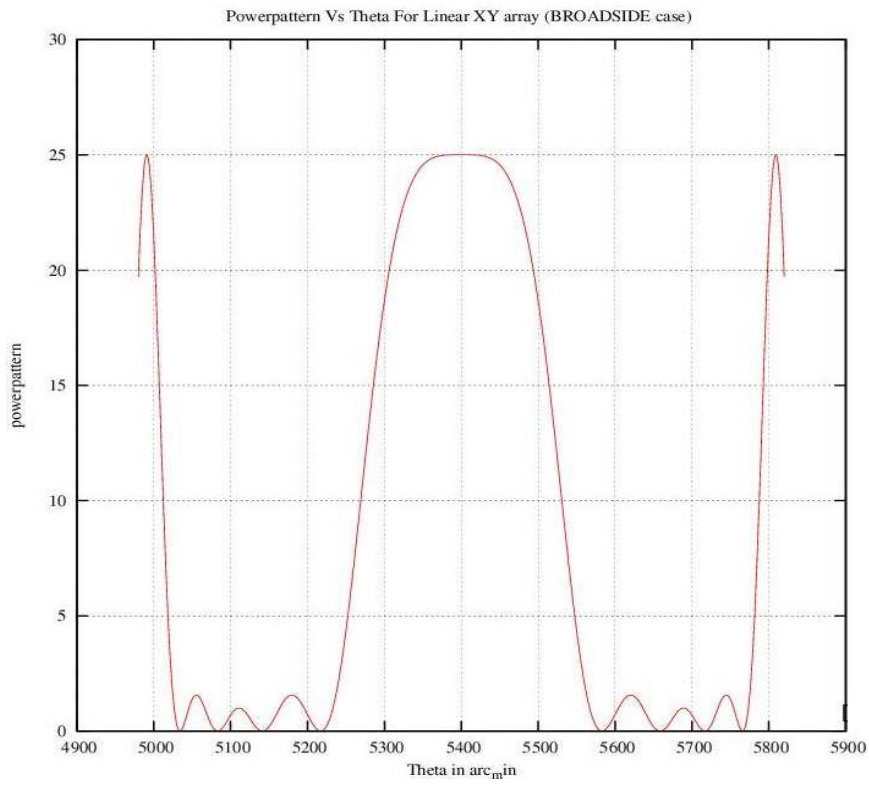
fig2.3



$$\theta_0 = 1 \text{ deg}$$

$$\phi_0 = 0 \text{ deg}$$

fig2.4



$\theta_0 = 5\text{deg}$
 $\phi_0 = 0\text{deg}$

fig2.5

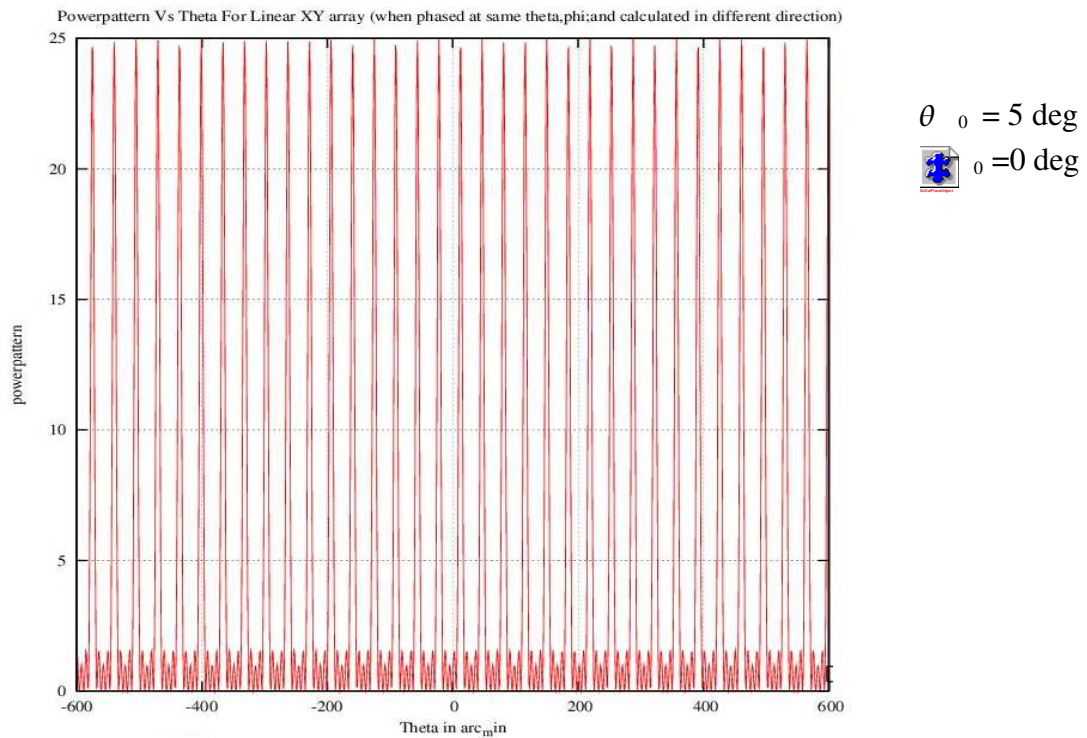


fig2.6

Now we make the following observation:

We keep $\theta_0 = 0 \text{ deg}$, $\phi_0 = 0 \text{ deg}$, $N=5$, $\lambda = 70 \text{ m}$

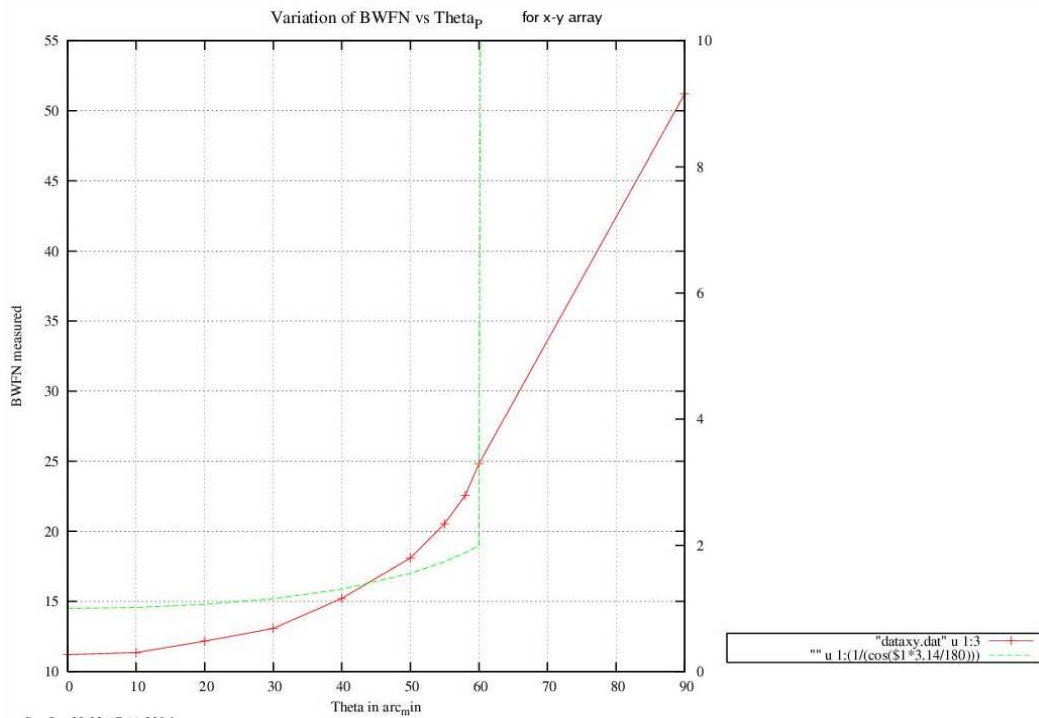
Changing the phasing angle θ_p , with $\phi_p = 0$ kept constant, we measure the BWFN from the data file. Also the calculated value is noted. A plot of variation BWFN vs theta_p is then made. Array along X axis and X=Y line is considered.

FOR ARRAY ALONG X=Y LINE

θ_p (in deg)	ϕ_p (in deg)	<i>BWFN(in deg)</i>		<i>Main lobe position (in deg)</i>
		<i>measured</i>	<i>calculated</i>	
0	45	11.2	11.36	0
10	45	11.35	11.44	10
20	45	12.17	11.86	20.1
30	45	13.07	12.72	30.11
40	45	15.22	14.12	40.15
50	45	18.1	16.32	50
55	45	20.55	17.86	55.3
58	45	22.58	18.98	58
60	45	24.83	19.82	59.9
90	45	51.22	51.42	90

From above data table we note that as the phasing angle drifts more and more away from the value of θ_0 , the discrepancy in measured and calculated value of BWFN increases. This is because as θ_p increases, the power pattern gets folded about $\theta = 90^\circ$. This is shown in the plots below. Then at $\theta_p = 90^\circ$ the pattern again peaks.

fig 2.7



FOR ARRAY ALONG X AXIS

θ_p (in deg)	ϕ_p (in deg)	<i>BWFN(in deg)</i>		<i>Main lobe position (in deg)</i>
		<i>measured</i>	<i>calculated</i>	
0	0	15.97	16.08	0
10	0	16.2	16.14	10
20	0	16.97	16.68	29.98
30	0	21.1	19.64	30
50	0	24.73	22.48	49.91
55	0	30.18	24.46	55
58	0	34.48	25.84	57.91
60	0	43.97	26.9	59.9
90	0	60.97	61.38	90

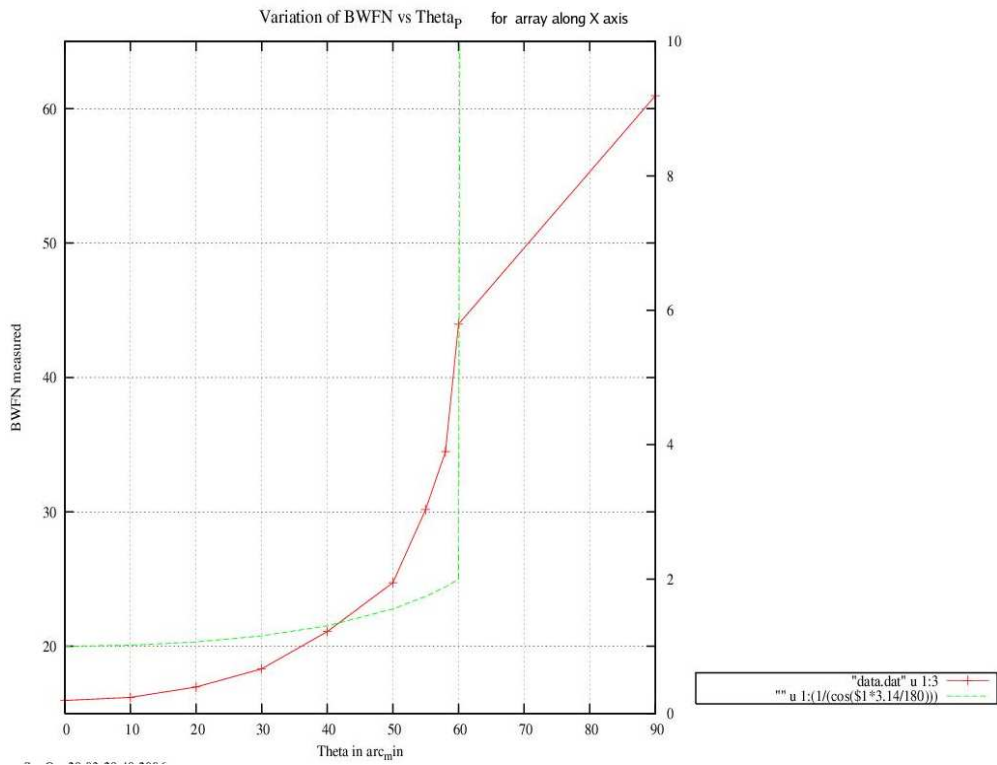


fig2.8

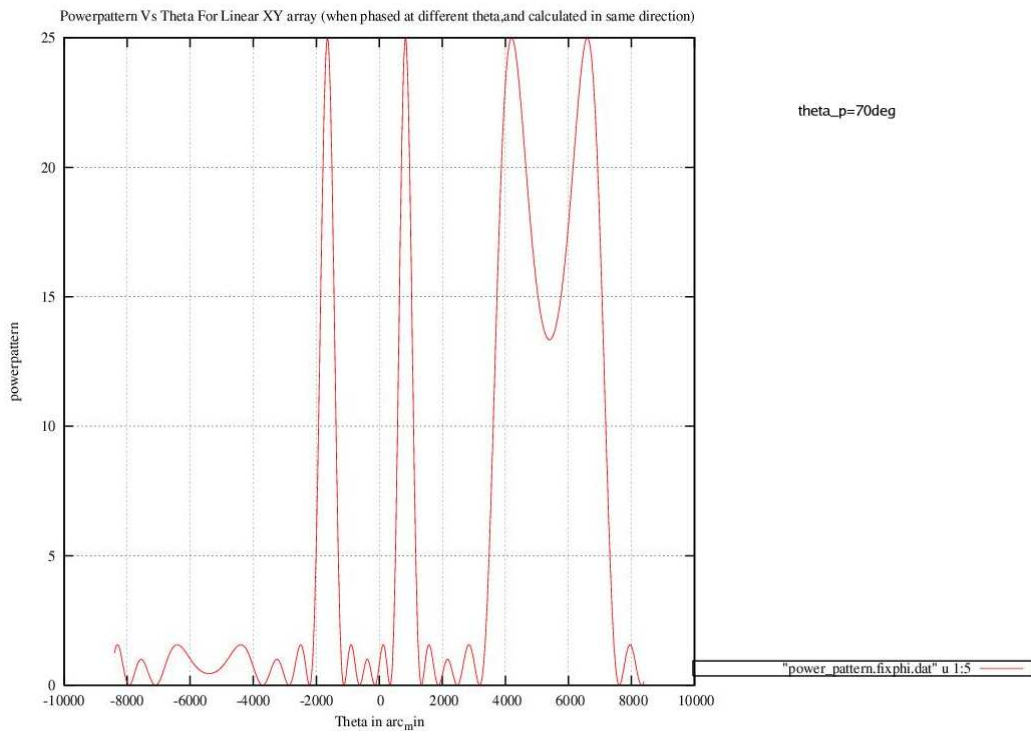


fig2.9

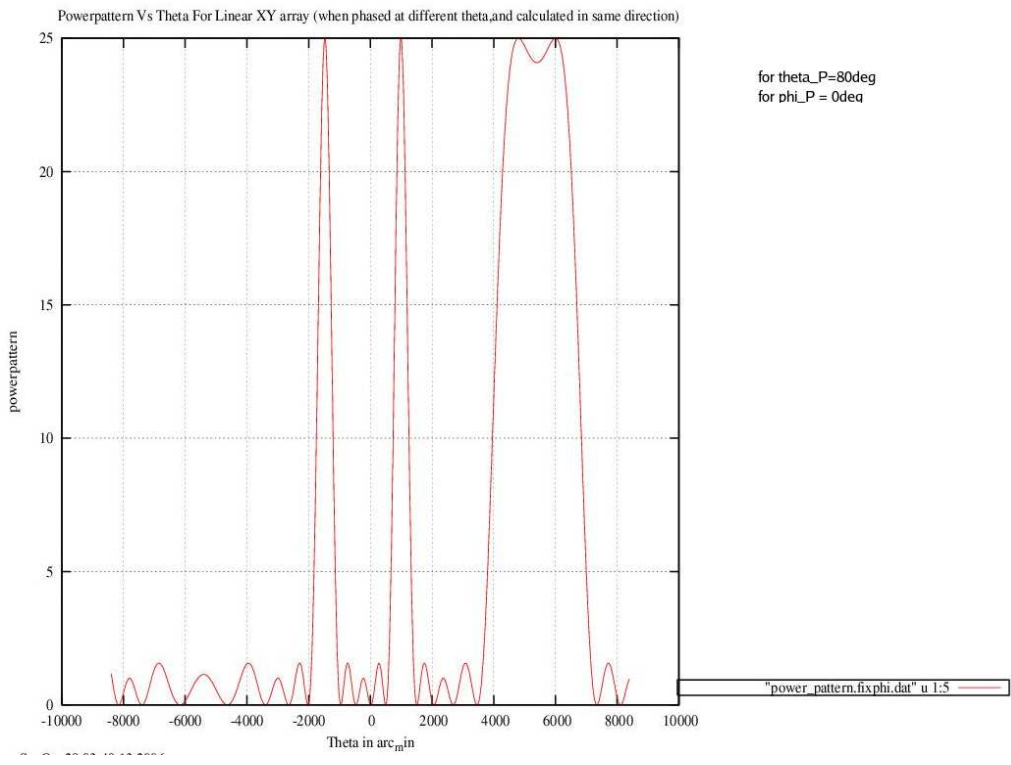


fig 2.10

CHAPTER 3

OBSERVATIONS ON GMRT ARRAY(CENTRAL SQUARE)

The Giant Meter Wave Radio Telescope(GMRT) consists of an array of 30 antennas,each antenna being 45m in diameter. The GMRT has a hybrid configuration with 14 of the antenna randomly distributed in a central region (approx 1 km across), called central square. The distribution of the antenna in the central square is “randomized” . The remaining antenna is distributed in a roughly Y shaped configuration with the length of each arm of the Y being approx 14km. The maximum baseline length between extreme arm antenna is approx 25km.

The central square antenna provide a large number of relatively short baselines. This is very useful for imaging large extended sources. The arm antennas on other hand are useful in imaging small sources where high angular resolution is essential.

We now study the power pattern vs angle θ for the GMRT array. We consider the 14 antenna of the central square with C2 as reference antenna. That the central square provides no grating lobes is confirmed on observing the power pattern over a wide range of θ .

The pole star is at 19deg elevation at GMRT . Therefore the zenith is 71deg from the polestar. In that case the maximum projected baseline is d (which is equal to 1km). If we phase the beam at an angle θ_p ,the projected baseline becomes $d*(90- \theta_p -19)$ deg.

The Half Power Beam Width (HPBW) is calculated using

$$\theta_{HPBW} \approx \lambda / (\text{maximum projected baseline})$$

where λ =wavelength of observation

The 1st table shows the HPBW (measured and calculated) for different phasing angles of the GMRT array.

WITH C2 AS REFERENCE ANTENNA.

θ_p (in deg)	ϕ_p (in deg)	θ_o (in deg)	ϕ_o (in deg)	Range of θ (in deg)	Accuracy of θ (in arc min)	HPBW (measured)(in arc min)	HPBW (calculated)(in arc min)	HPBW (cal)/ HPBW (meas)
10	0	10	0	0.3	0.1	5.4	7	1.3
30	0	30	0	0.3	0.1	3.43	4.5	1.33
45	0	45	0	0.3	0.1	2.82	3.8	1.35
71	0	71	0	0.3	0.1	2.61	3.43	1.3

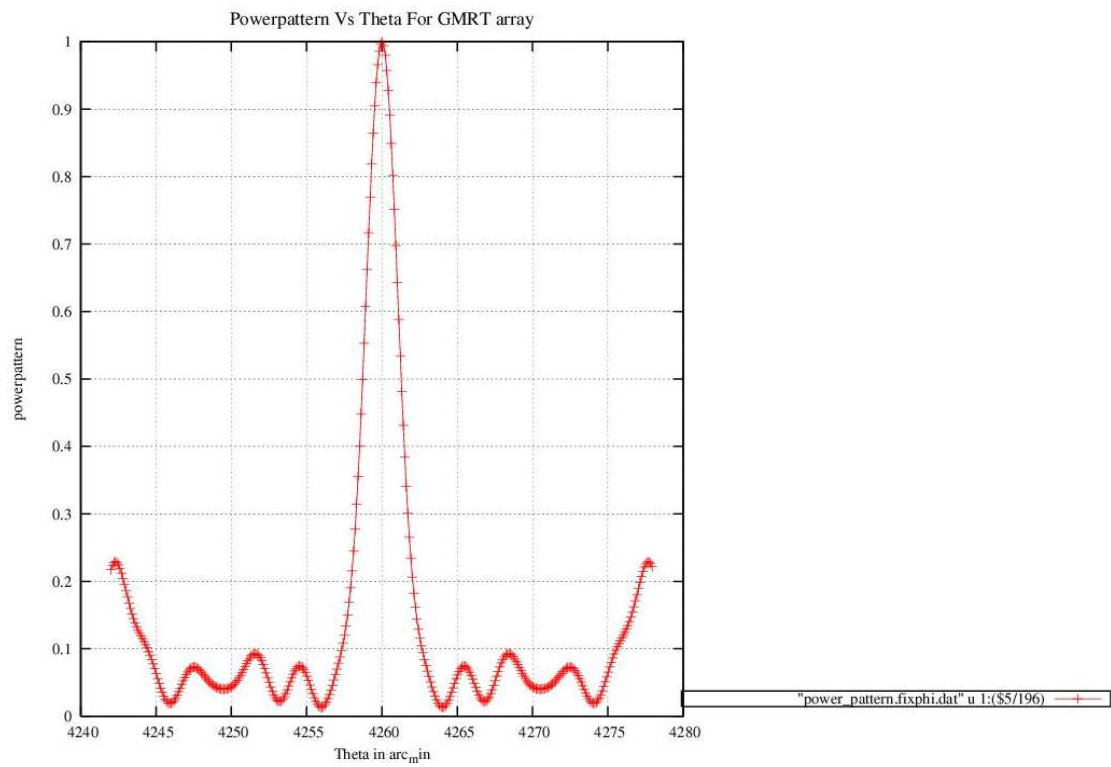


fig 3.1

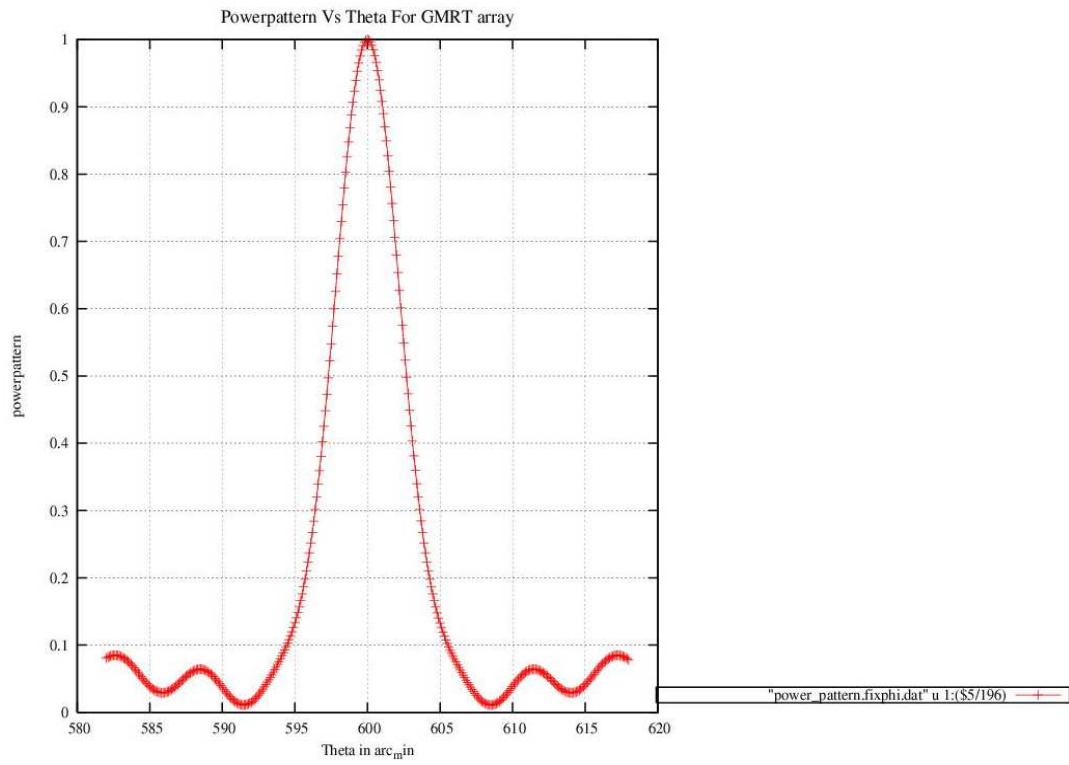


fig 3.2

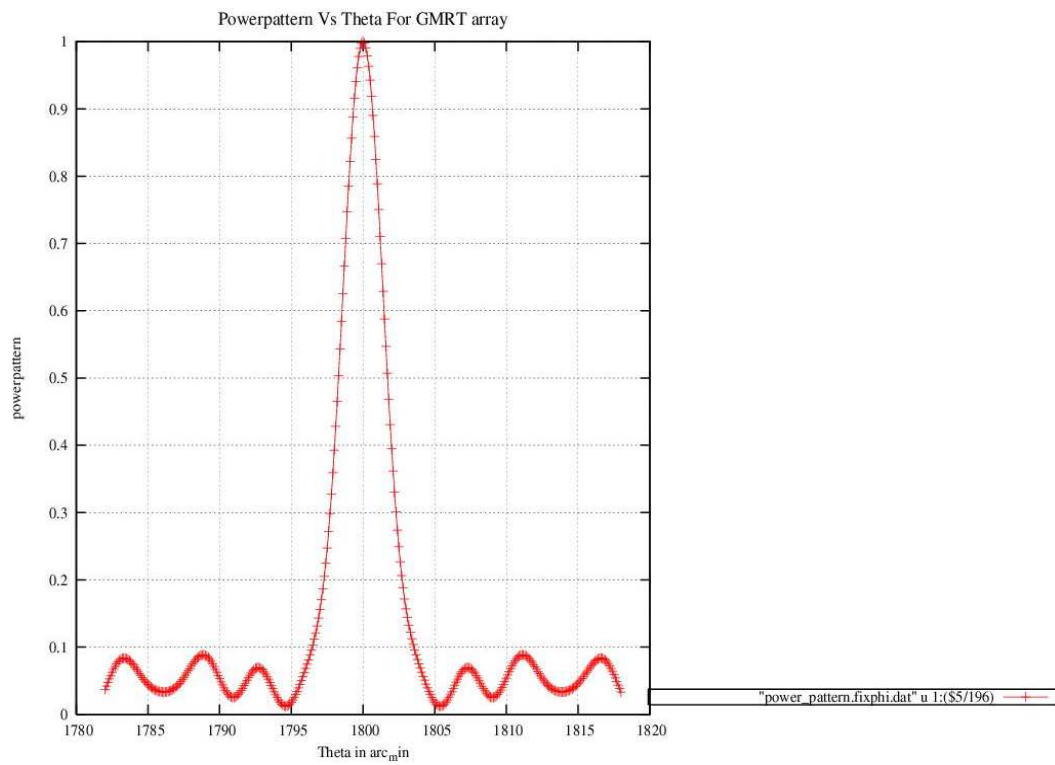


fig3.3

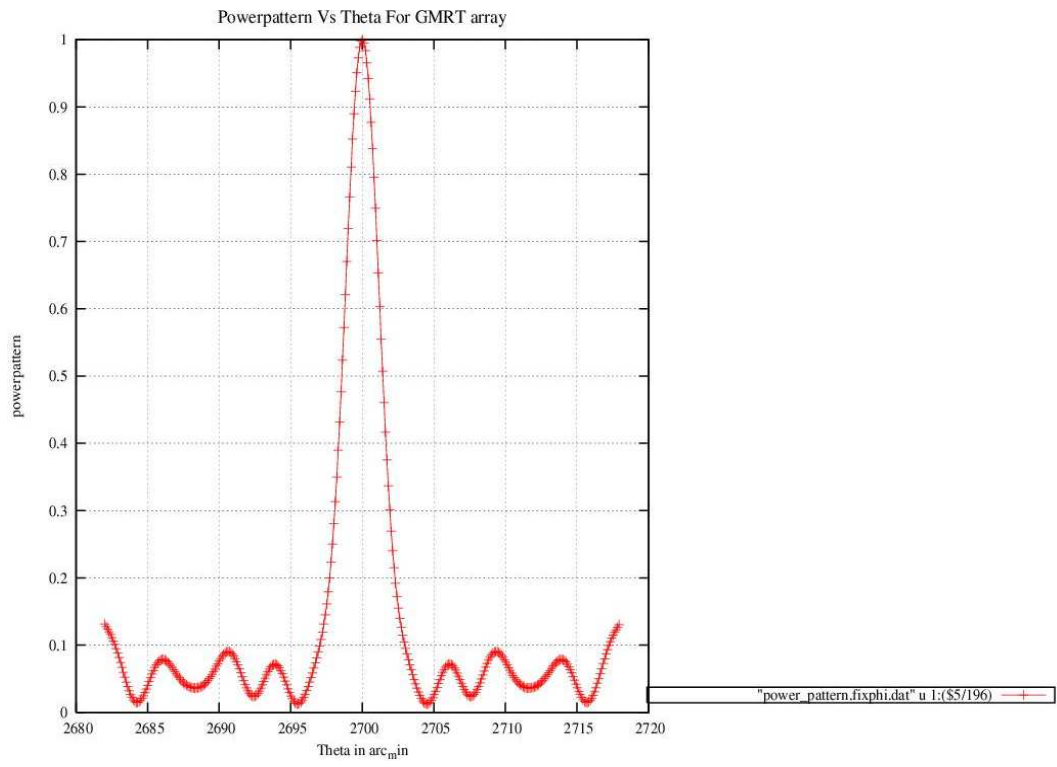


fig 3.4

WITH C13 AS REFERENCE ANTENNA

θ_p (in deg)	ϕ_p (in deg)	θ_o (in deg)	ϕ_o (in deg)	Range of θ (in deg)	Accuracy of θ (in arc min)	HPBW (measured)(in arc min)	HPBW (calculated)(in arc min)	HPBW (cal)/ HPBW (meas)
71	0	71	0	0.3	0.1	2.61	3.43	1.3

Thus we observe that on changing the reference antenna the discrepancy in the value HPBW measured and calculated still remains.

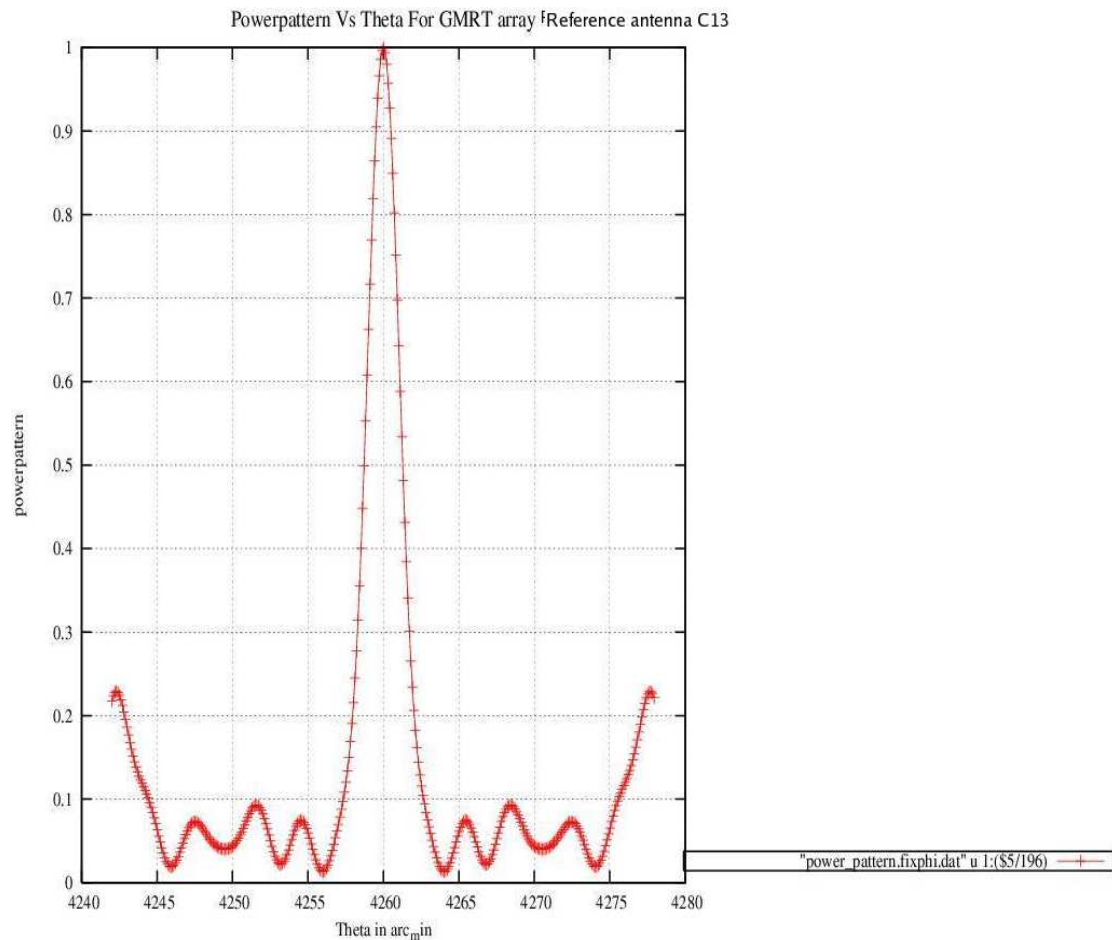


fig 3.5

Now we study the variation in the power pattern of the GMRT array with theta and phi variations, by varying the phasing angles. The pattern is calculated about the phasing angles. The range of theta and phi is kept constant throughout. The plots are taken for fixed phi (i.e variation in theta) and fixed theta (i.e variation in phi).

Range of theta = 5deg Range of phi = 5 deg
 Accuracy of theta = 1' accuracy of phi = 1'

1.

$\theta_p = 71 \text{ deg}$ $\phi_p = 0 \text{ deg}$
 $\theta_0 = 71 \text{ deg}$ $\phi_0 = 0 \text{ deg}$
 fixed theta = 71 deg fixed phi = 0 deg

a.

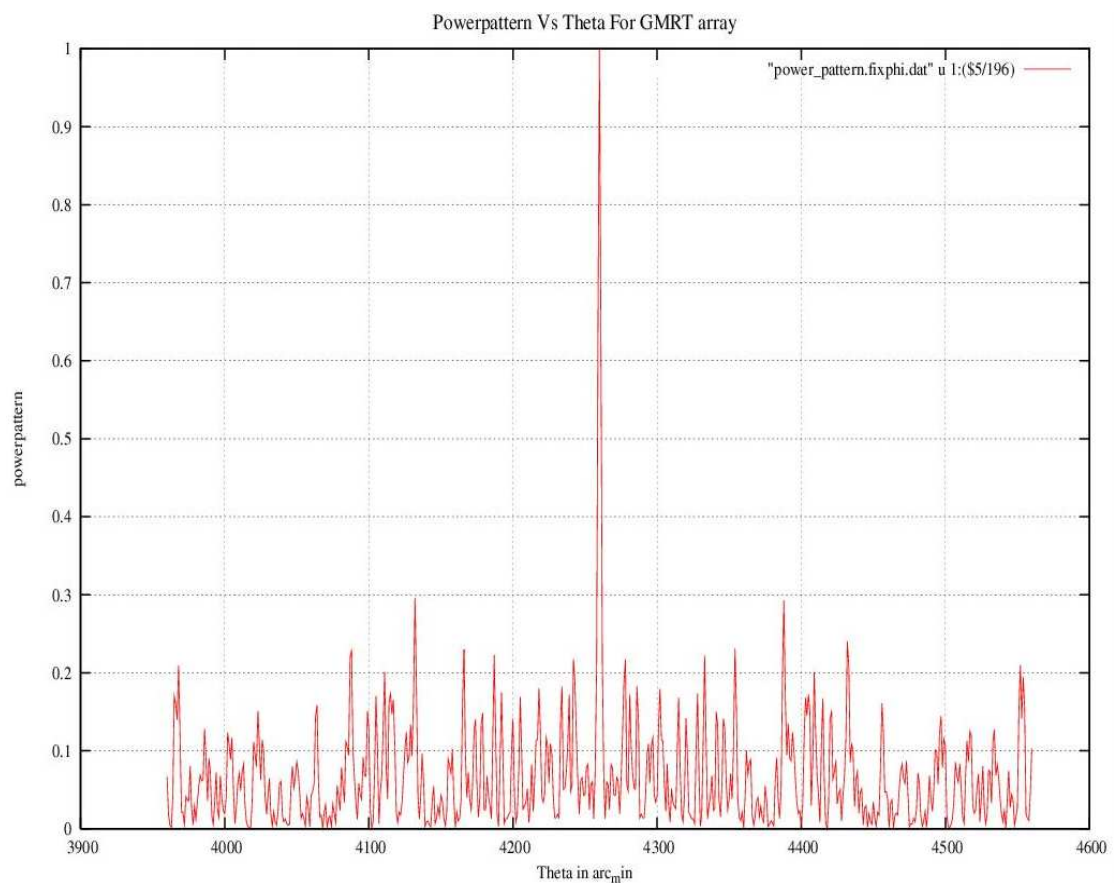


fig3.6

From above plot: side lobes with ~30% power occur 150' away from θ_p

b.

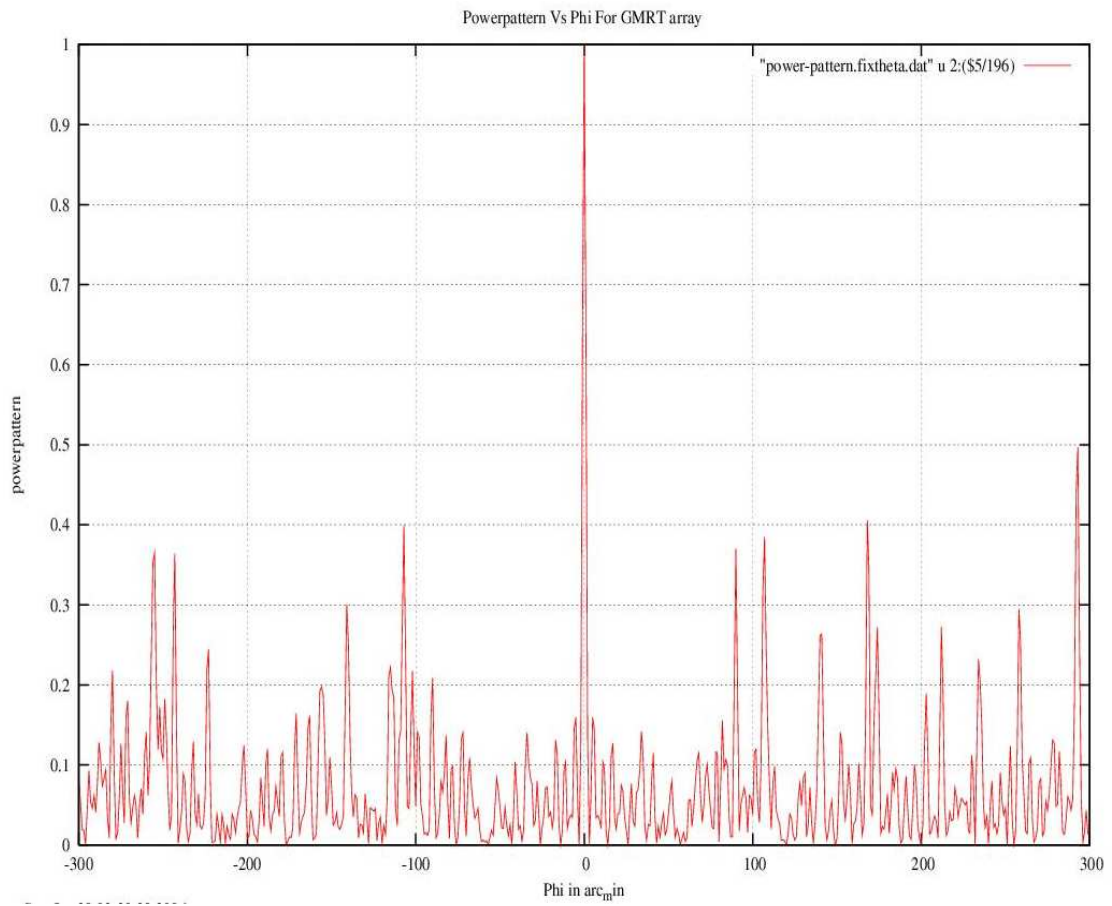






fig 3.7

From above plot:side lobes with ~40%power occur 100' away from ϕ_p

2.  $\theta_p = 71 \text{ deg}$  $\theta_p = 45 \text{ deg}$
 $\theta_0 = 71 \text{ deg}$  $\theta_0 = 45 \text{ deg}$
fixed theta = 71 deg fixed phi = 45 deg

a.

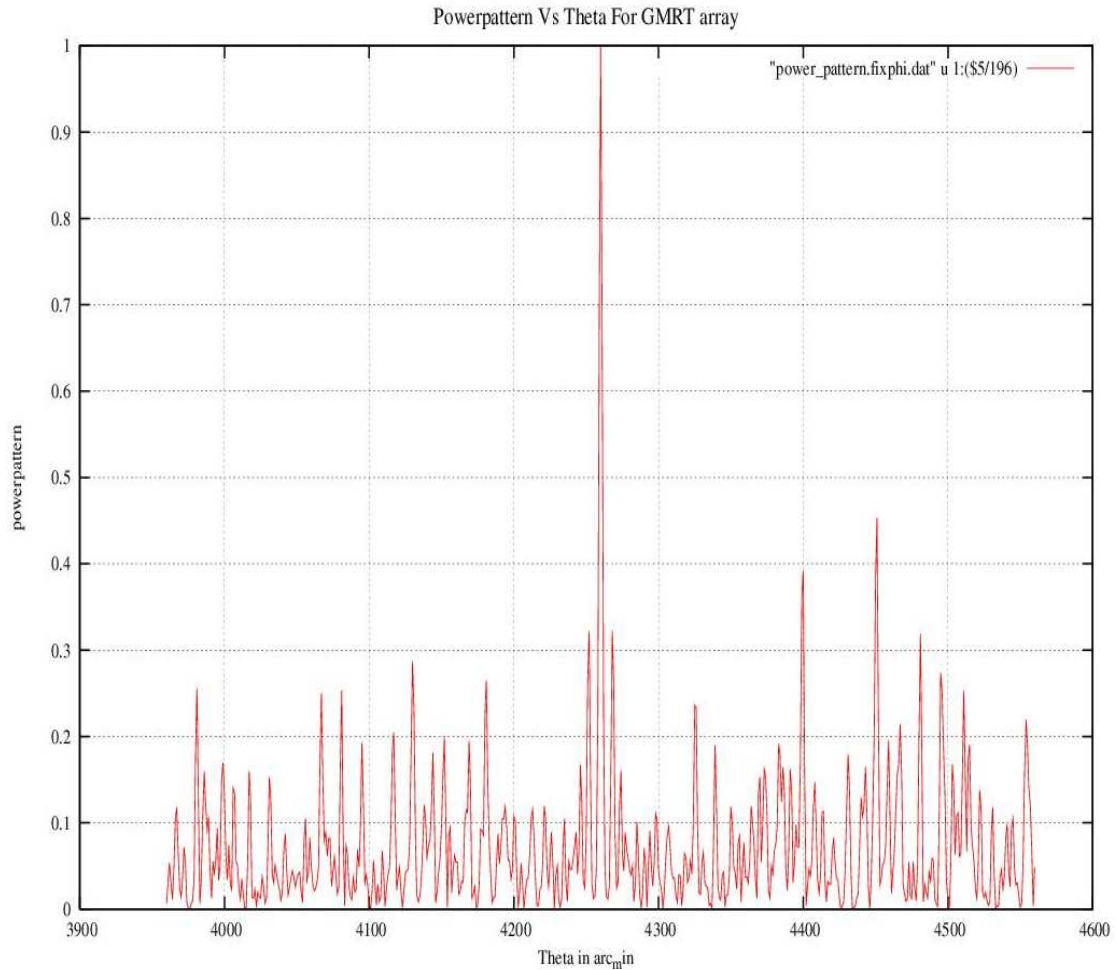


fig 3.8

From above plot: side lobes with ~33% power occur ~10' away from θ_p . Side lobe with 40% power occur ~200' away from θ_p

b.

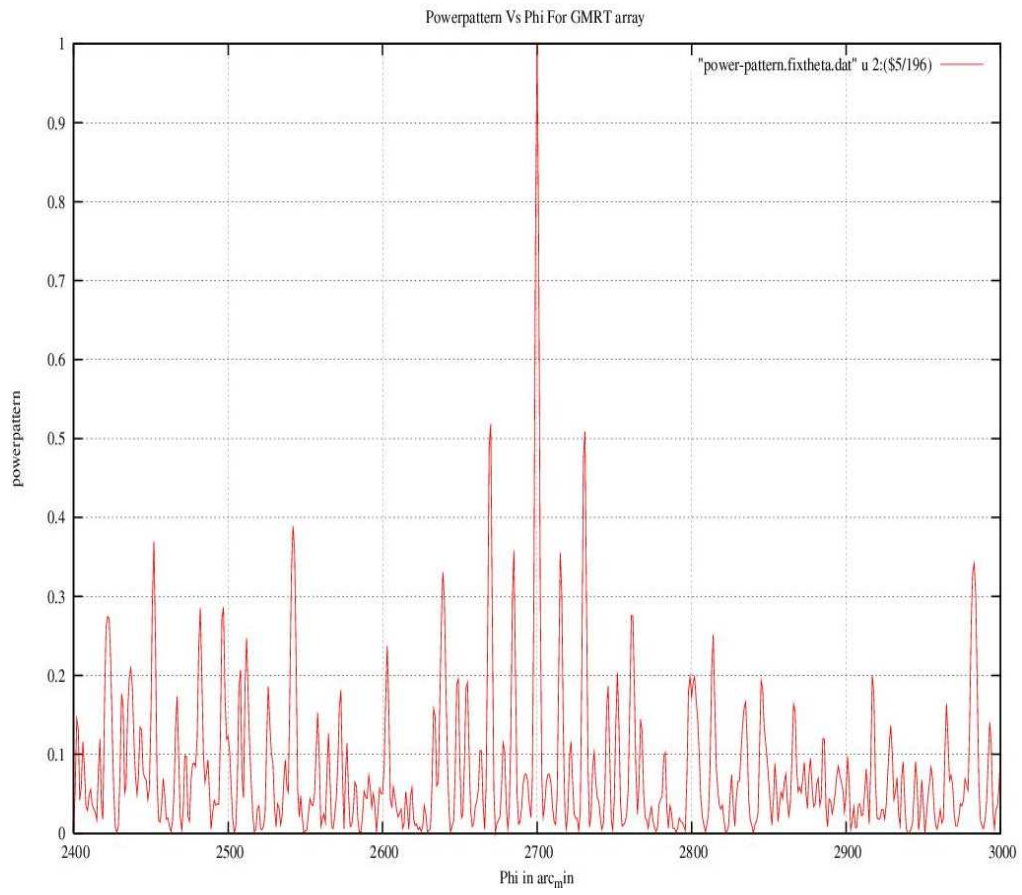


fig 3.9

From above plot: Side lobe with ~52% power occur ~30' away from ϕ_p

3. $\theta_p = 0 \text{ deg}$ $\phi_p = 0 \text{ deg}$
 $\theta_0 = 0 \text{ deg}$ $\phi_0 = 0 \text{ deg}$
fixed theta=0deg fixed phi=0 deg

a.

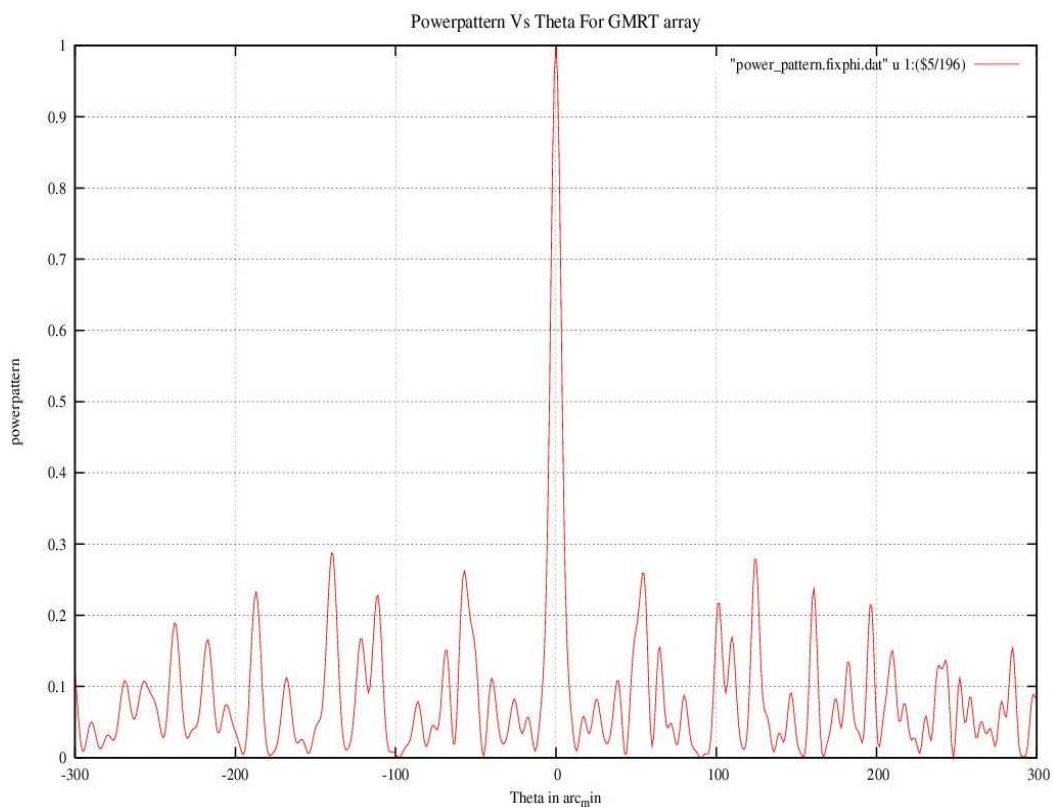


fig 3.10

From above plot:side lobes with 25% power occur 50 ' away from θ_p

b.

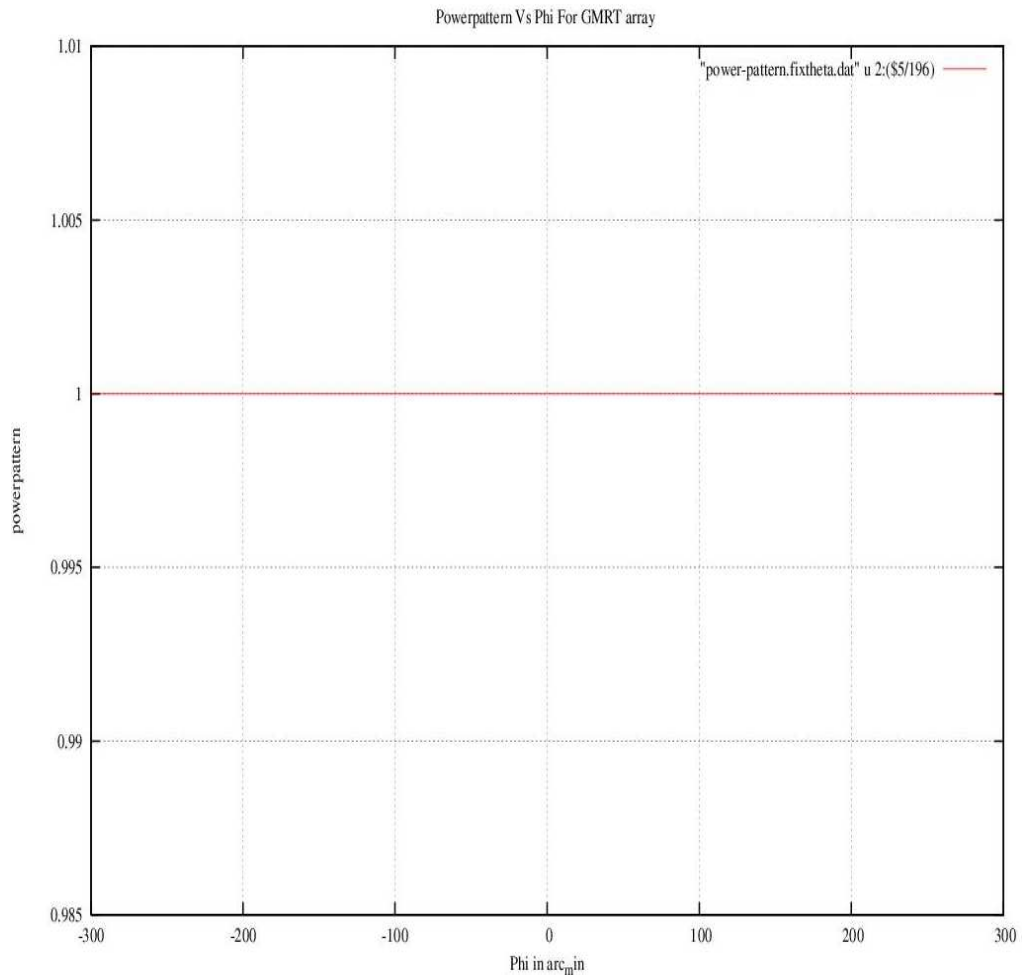


fig 3.11

From above plot: Mathematically When $\theta_p = 0$ deg the contribution of ϕ_p to the power equation is nullified,(as seen in the formula). Also,physically when $\theta_p = 0$,on varying phi we obtain a power pattern that is cylindrically symmetric.

4. $\theta_p = 0\text{deg}$ $\phi_p = 45\text{deg}$
 $\theta_0 = 0\text{ deg}$ $\phi_0 = 45\text{deg}$
fixed theta = 0 deg fixed phi= 45deg

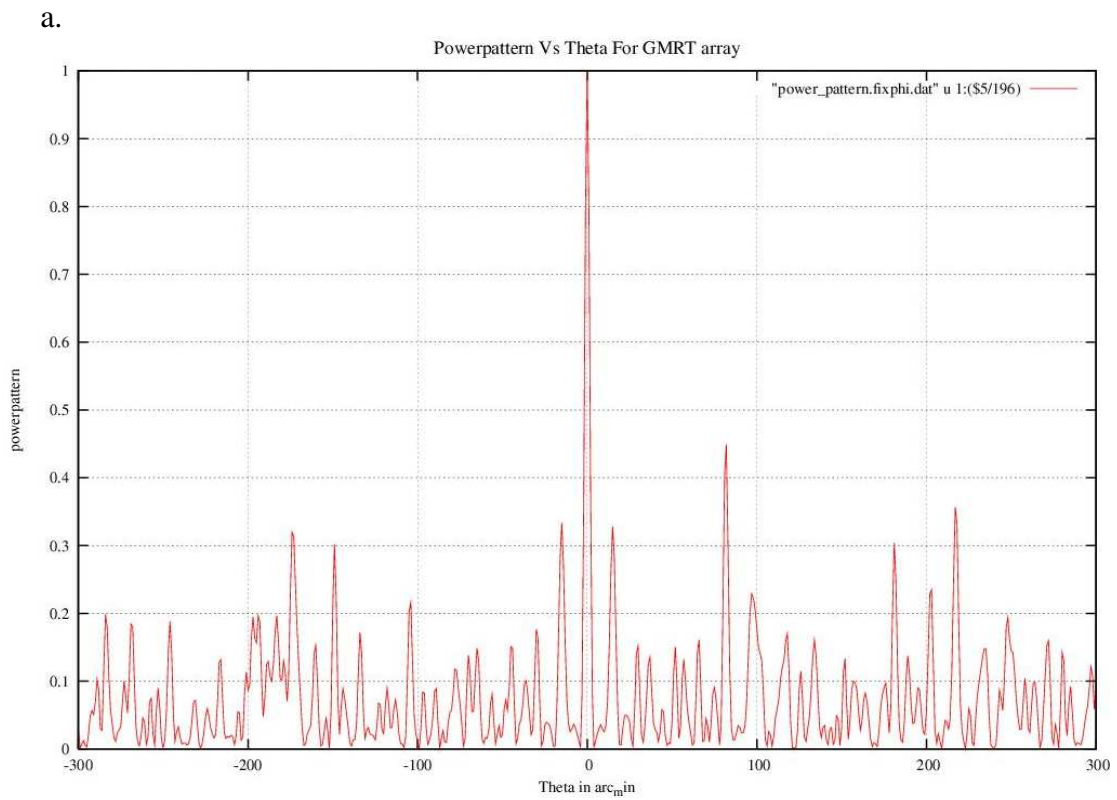


fig 3.12

From above plot: Side lobe with ~33% power occur ~20' away from θ_p

b.

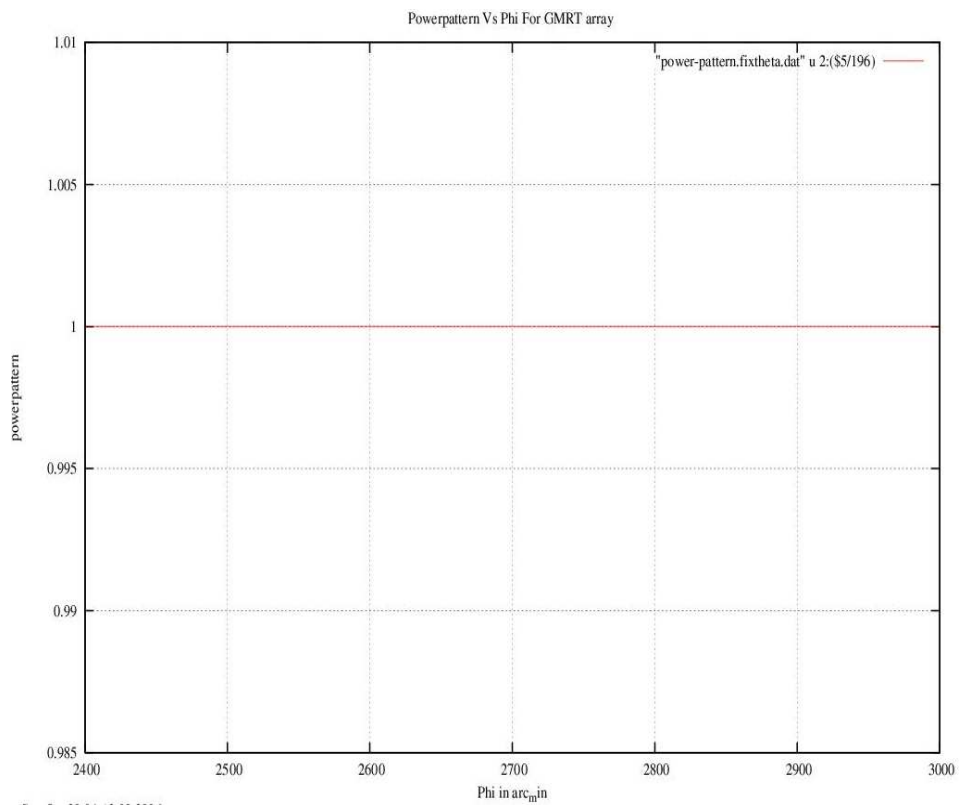


fig 3.13

Same as 3b.

5. $\theta_p = 45\text{deg}$ $\phi_p = 0\text{ deg}$
 $\theta_p = 45\text{deg}$ $\phi_0 = 0\text{ deg}$
fixed theta=45 deg fixed phi=0 deg
a.

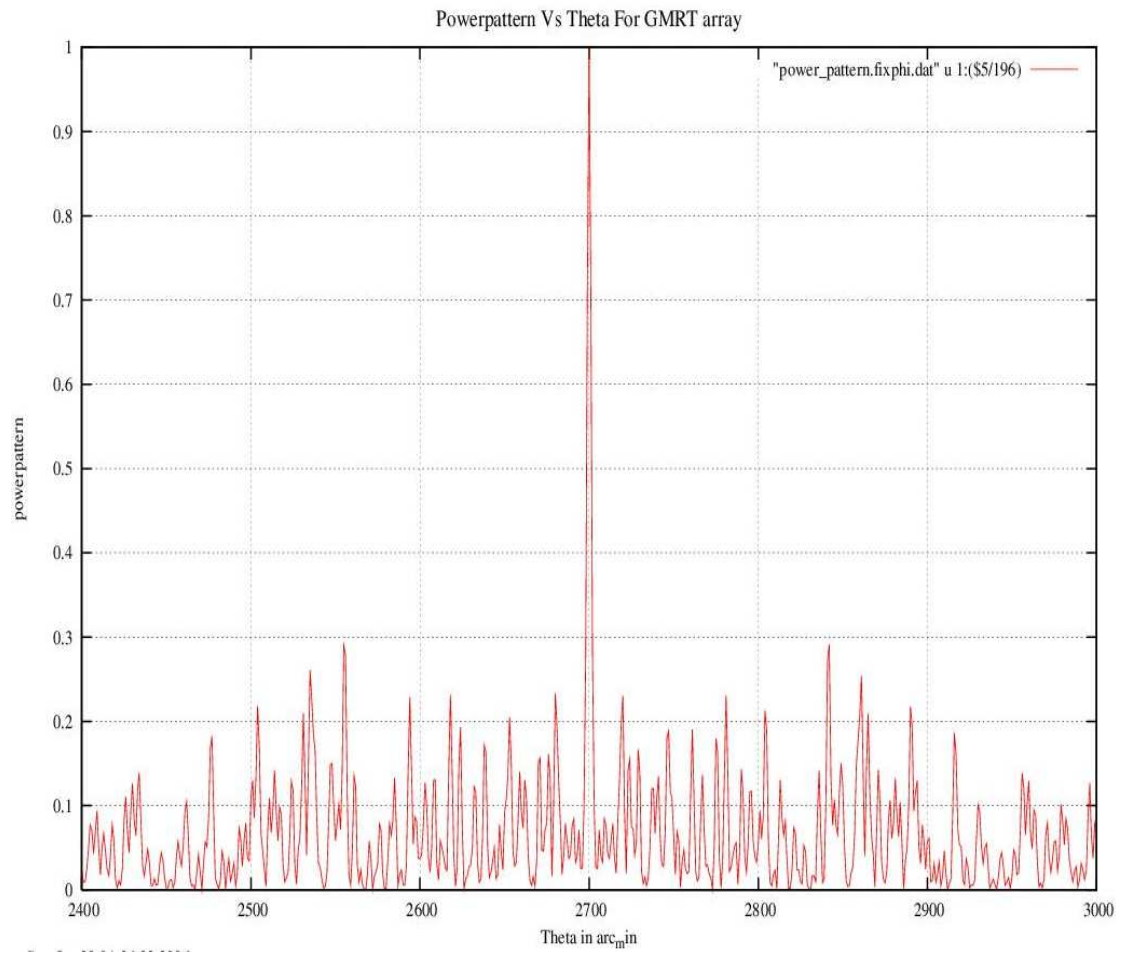


fig3.14

From above plot:side lobe power below 30% over the 5deg range.

b.

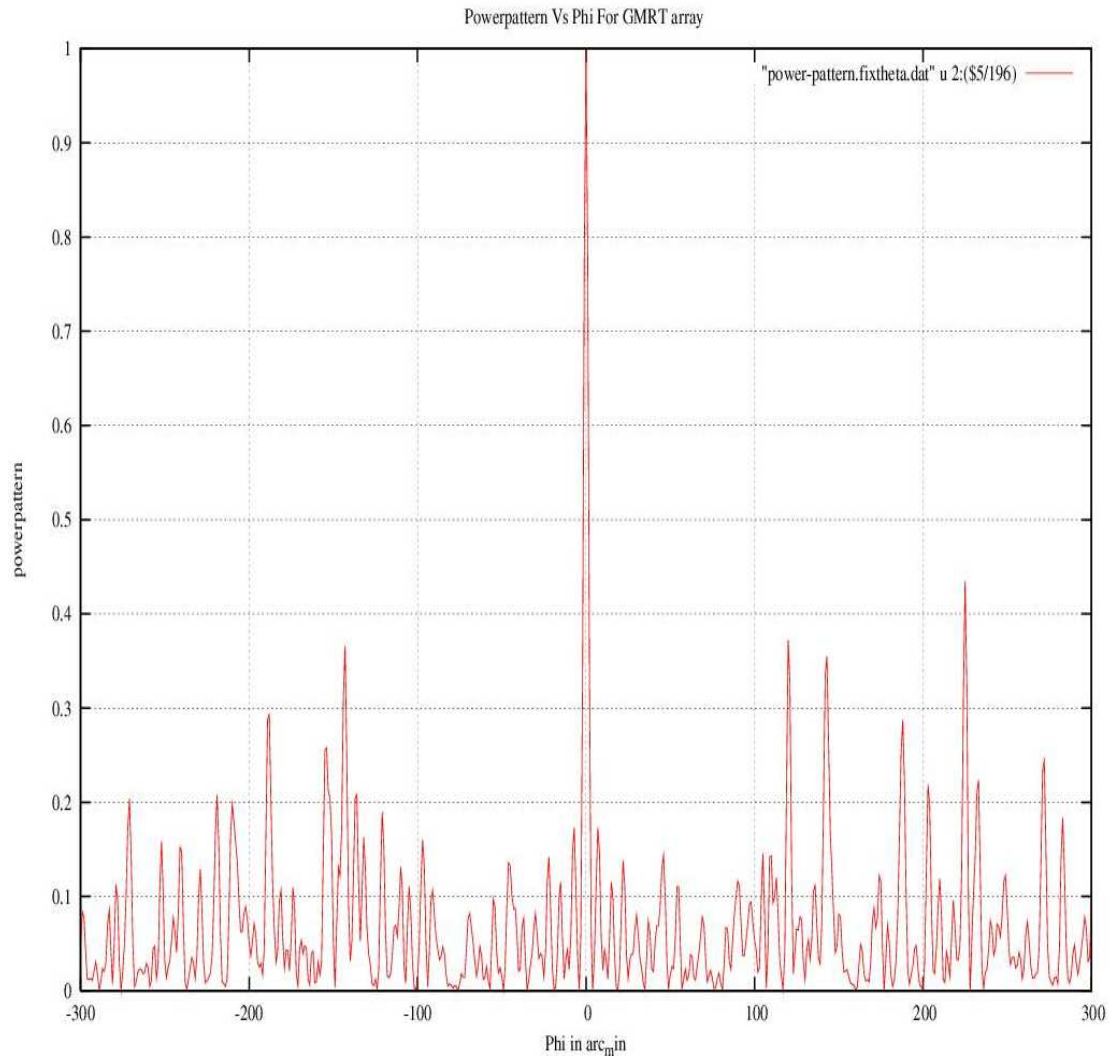


fig 3.15

From above plot:side lobe below 2% power over 100' range from ϕ_p

6. $\theta_p = 45\text{deg}$ $\phi_p = 45\text{ deg}$
 $\theta_p = 45\text{ deg}$ $\phi_0 = 45\text{ deg}$
fixed theta=45 deg fixed phi=45 deg

a.

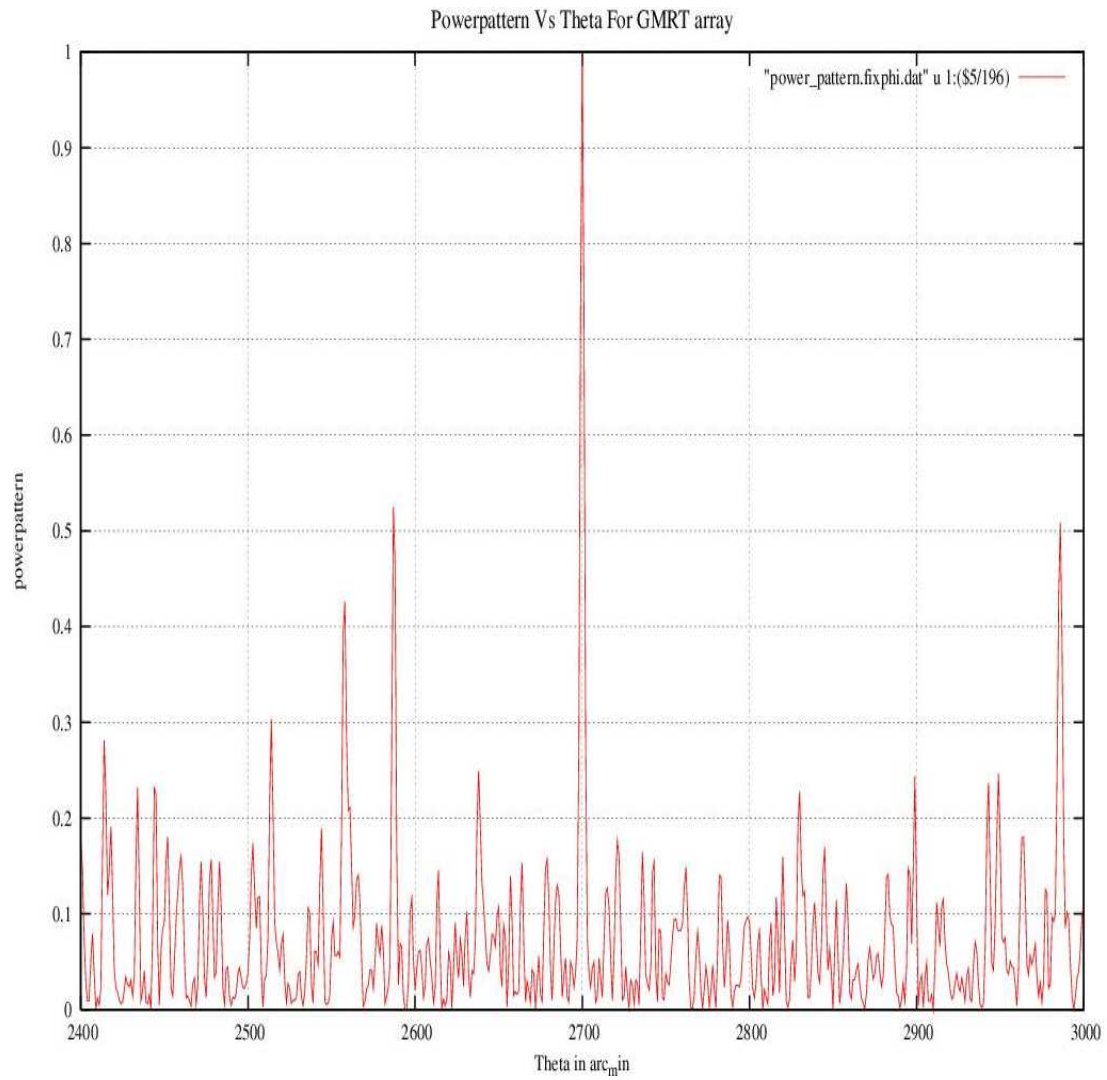


fig3.17

From above plot: side lobe with ~52% power 100' away from θ_p

b.

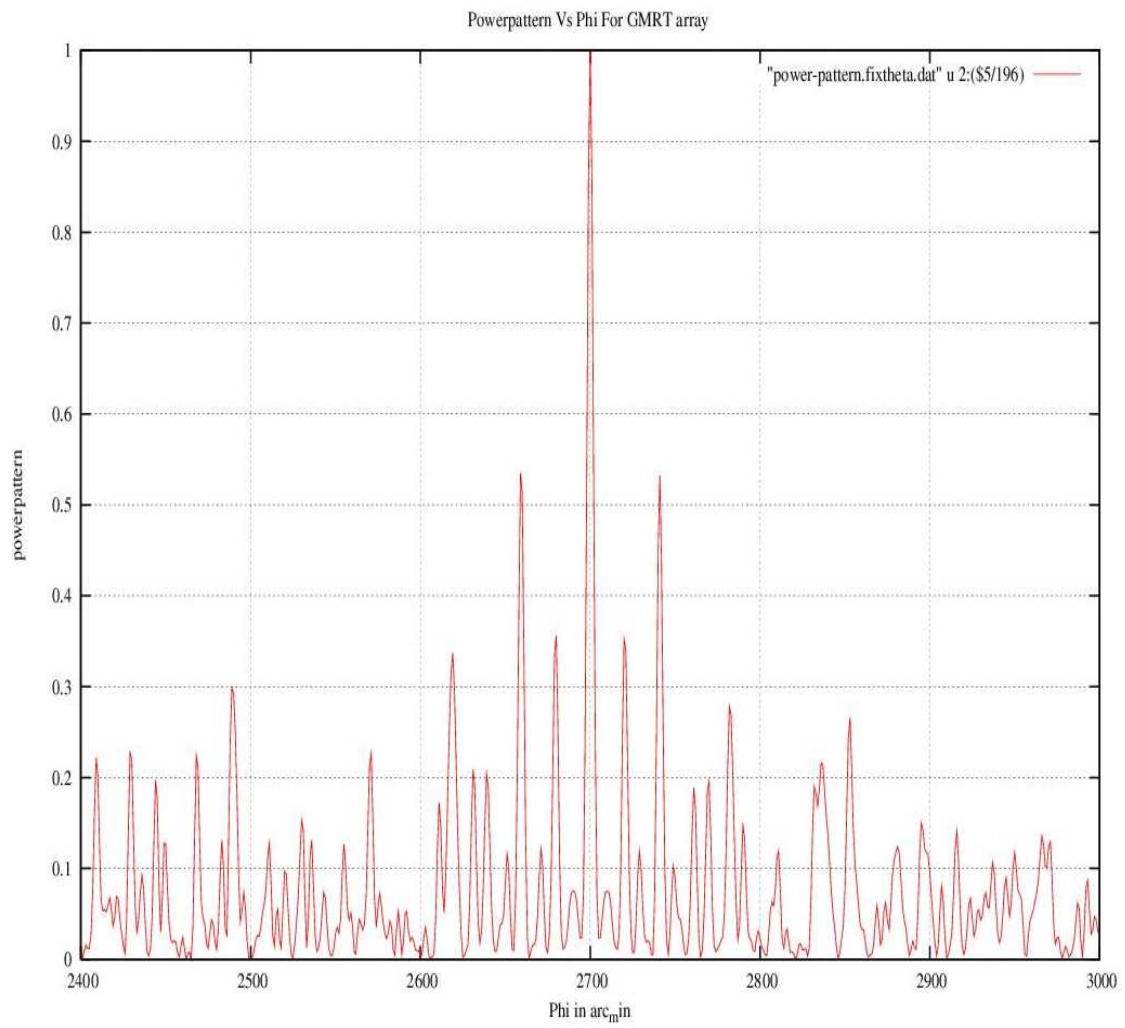


fig3.18

From above plot: side lobe with ~53% power 40' away from ϕ_p

We demonstrate a 3-D plot of the GMRT power pattern and make a contour plot with the given script file.

$$\theta_p = \theta_0 = 71 \text{ deg} \quad , \quad \phi_p = \phi_0 = 0 \text{ deg}$$

Range of theta = 0.5deg , Range of phi = 1deg
 Accuracy of theta = 0.2' , Accuracy of phi = 20'

The script file used to plot graph for the above parameters is :

```

set term post landscape enhanced colour colourtext "Times NewRoman" 10
set out "test.ps"
set autoscale
set xlabel 'Theta in arc_min'
set ylabel 'phi in arc_min'
set zlabel 'powerpattern'
set title 'Powerpattern Vs Theta For GMRT array' font "Times NewRoman, 12"
show title
set time
set grid
set contour
set hidden3d
splot "power_pattern.dat" u 1:2:($5/196) w l
    
```

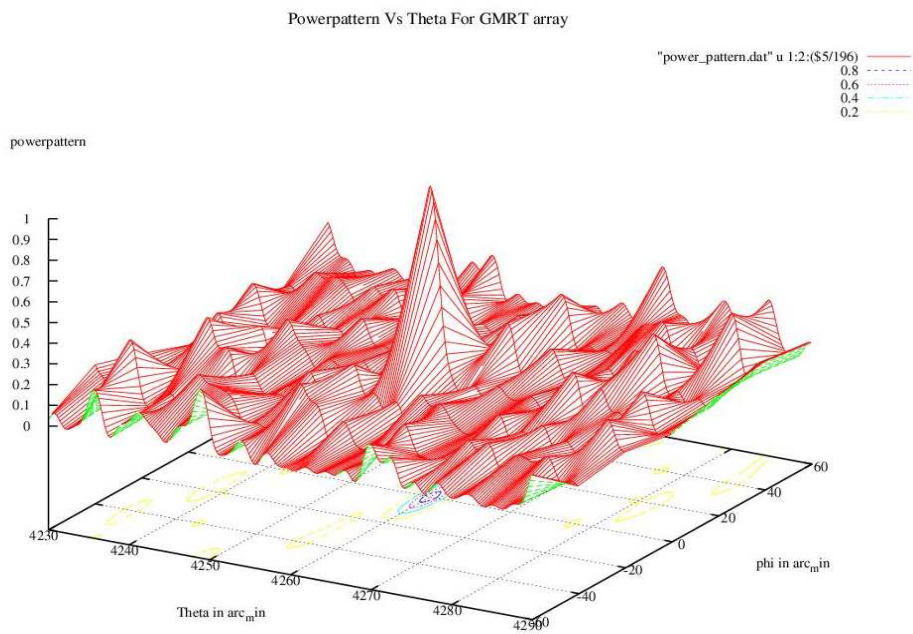


fig3.19

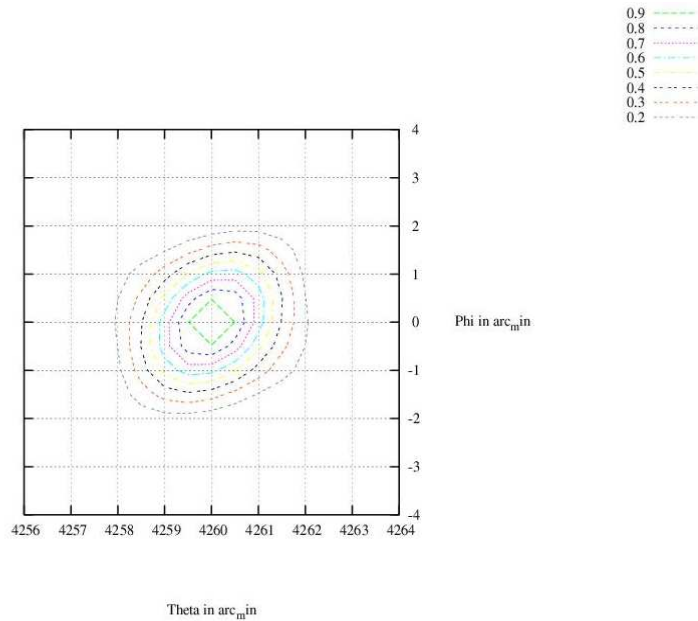


Fig 3.20

```

set term post landscape enhanced colour colourtext "Times NewRoman" 10
set out "test.ps"
set autoscale
set xlabel 'Theta in arc_min'
set ylabel 'phi in arc_min'
set zlabel 'powerpattern'
set title 'Powerpattern Vs Theta For GMRT array phased in any direction in
theta phi plane , showing contour' font "Times NewRoman, 12"
show title
set time
set grid
set contour
set dgrid3d 90,50,8
set cntrparam levels 2
set cntrparam levels incremental .9,.1,1
splot "power_theta_p0.dat" u 3:4:( $\$5/196$ ) w l, "power_theta_p1.dat" u 3:4:
( $\$5/196$ ) w l,"power_theta_p2.dat" u 3:4:( $\$5/196$ ) w l,"power_theta_p3.dat" u
3:4:( $\$5/196$ ) w l,"power_theta_p4.dat" u 3:4:( $\$5/196$ ) w l,"power_theta_p5.dat"
u 3:4:( $\$5/196$ ) w l,"power_theta_p6.dat" u 3:4:( $\$5/196$ ) w
l,"power_theta_p7.dat" u 3:4:( $\$5/196$ ) w l,"power_theta_p8.dat" u 3:4:( $\$5/196$ )
w l

```


CHAPTER 4

MODIFICATION OF THE C CODE

The program is modified with the concept in mind that multiple beams is needed to be generated to track an extended source. This is done by programming the code for multiple values of θ_p .

In the first stage, a for loop of θ_p is run over and above the loop of phi and theta. This generates 3 discrete beams phased at 0, 0.5, 1 deg in theta. The power of each of these phasing is put into 3 individual files and their plot taken together as shown.

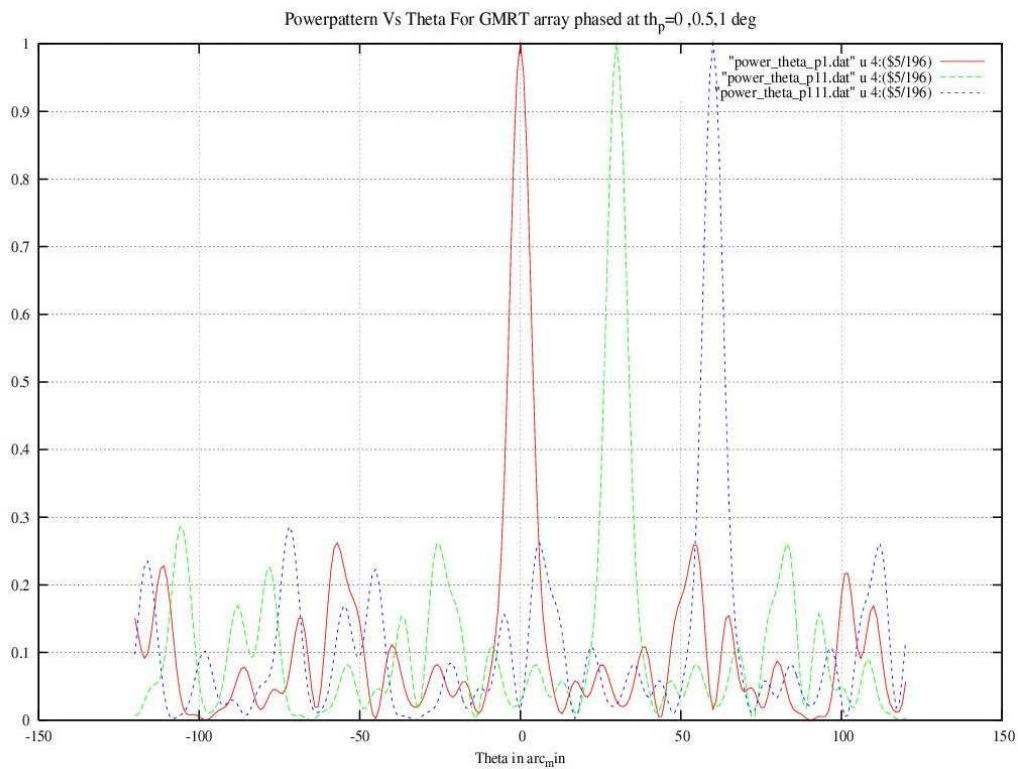


fig 4.1

The variation of power pattern with phi is as shown below.

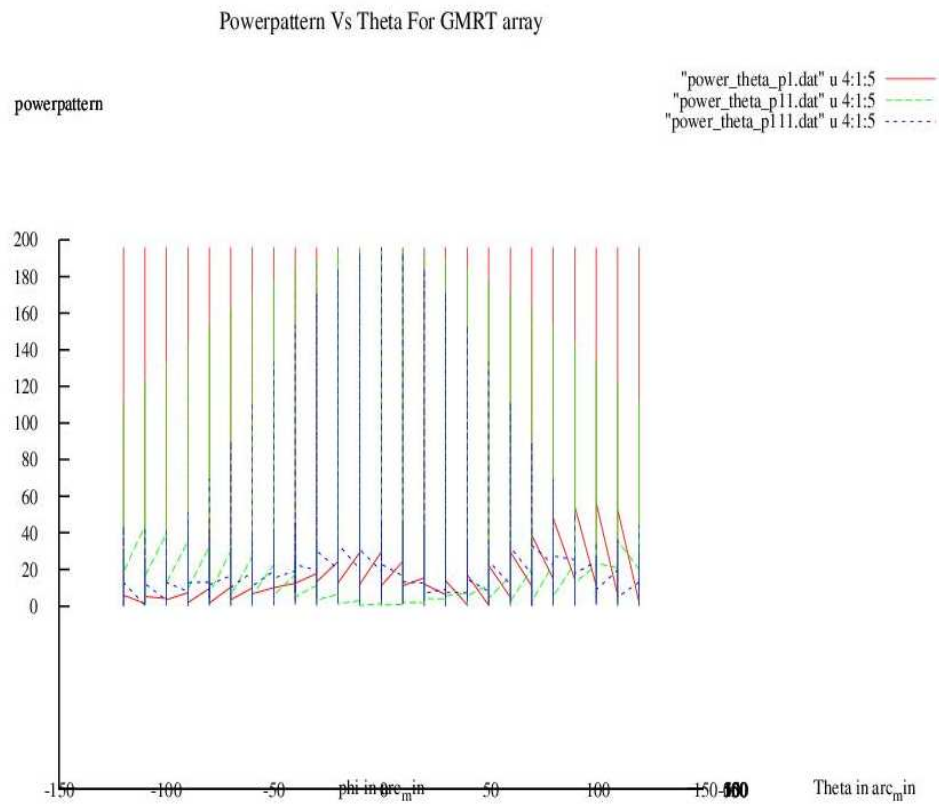


fig4.2

From the above plot we can see that when $\theta_p = 0$ deg there is no variation in the power pattern with phi. This can be explained from the equations below.

$$\begin{aligned}
 x_comp &= x[i] * (\sin(\min * th) * \cos(\min * phi) - \sin(\min * th_P * 60) * \cos(\min * phi_P * 60)); \\
 y_comp &= y[i] * (\sin(\min * th) * \sin(\min * phi) - \sin(\min * th_P * 60) * \sin(\min * phi_P * 60)); \\
 z_comp &= z[i] * (\cos(\min * th) - \cos(\min * th_P * 60));
 \end{aligned}$$

When $\theta_p = 0$, the contribution of phi part to the power components is constant. But when $\theta_p \neq 0$, the phi variations are as shown.

To generate multiple beams in 2-D we seek variation both in theta and phi, so that

beams phased at any θ_p and ϕ_p can be obtained. The separation of the phased beams in theta and phi is provided by the user. A corresponding data file is generated which stores the value of theta and phi .

If phasing directions are theta_P and phi_P and number of beams to be generated along either side of theta_P , phi_P =X(say) , then total number of multiple beams generated =(2X-1)*(2X-1).

Below are 2 examples of this:

1)Here,

theta_p=0deg,phi_P=0deg

theta_0=0deg,phi_0=0deg.

range of theta=2deg,range of phi=2 deg

accuracy in theta=10' ,accuracy in phi=10'

No. of multiple beams to be generated in theta and phi= 2

separation in between beams =60'.

The input file is as shown: ("input_thp_phip.dat")

# theta_P	phi_P
-60.000000	-60.000000
-60.000000	0.000000
-60.000000	60.000000
0.000000	-60.000000
0.000000	0.000000
0.000000	60.000000
60.000000	-60.000000
60.000000	0.000000
60.000000	60.000000

The 3-D plot is as shown:

The above plot along the axis is as shown:

Powerpattern Vs Theta For GMRT array phased in any direction in theta phi plane

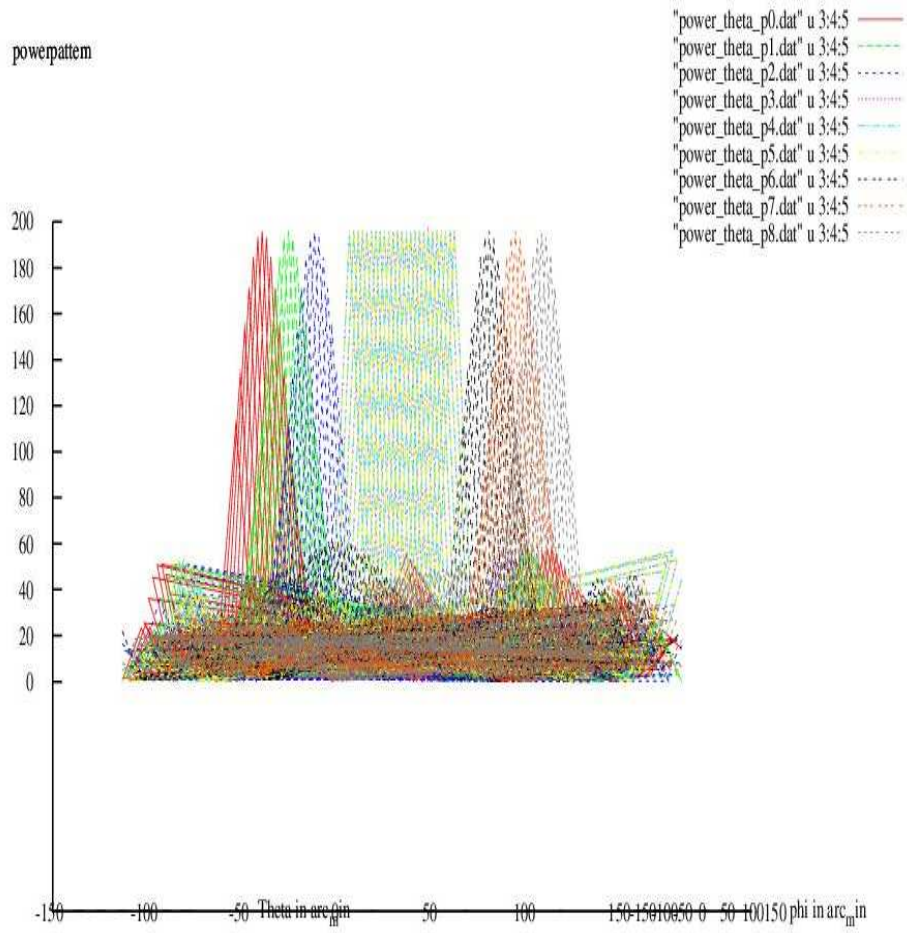


fig 4.3

Powerpattern Vs Theta For GMRT array phased in any direction in theta phi plane

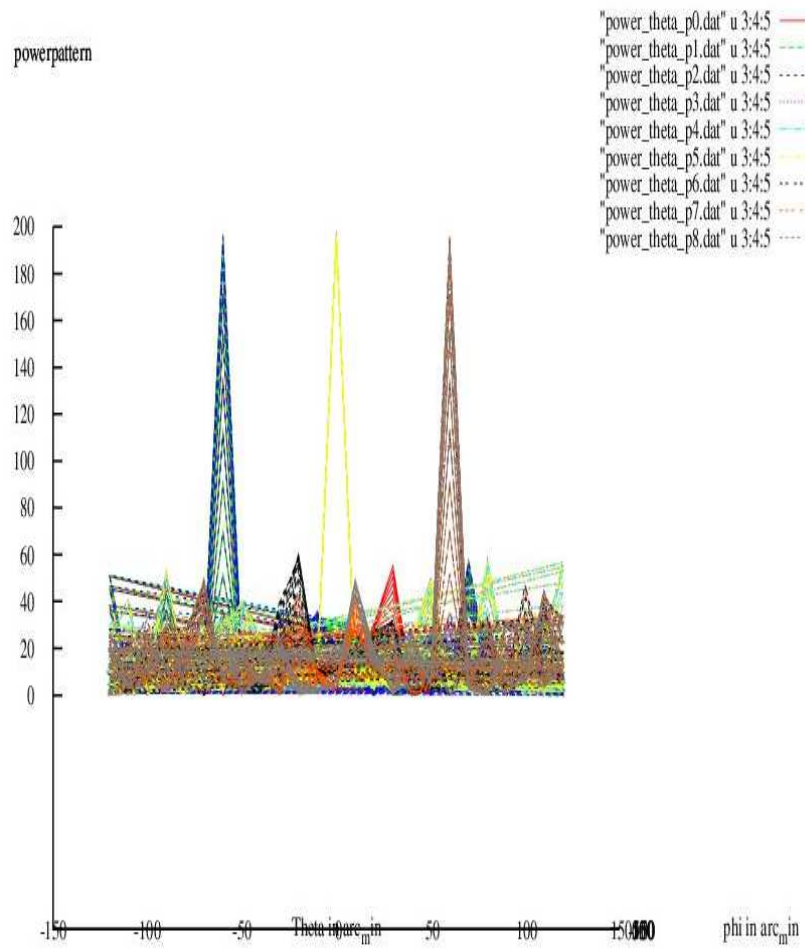


fig 4.4

Powerpattern Vs Theta For GMRT array phased in any direction in theta phi plane

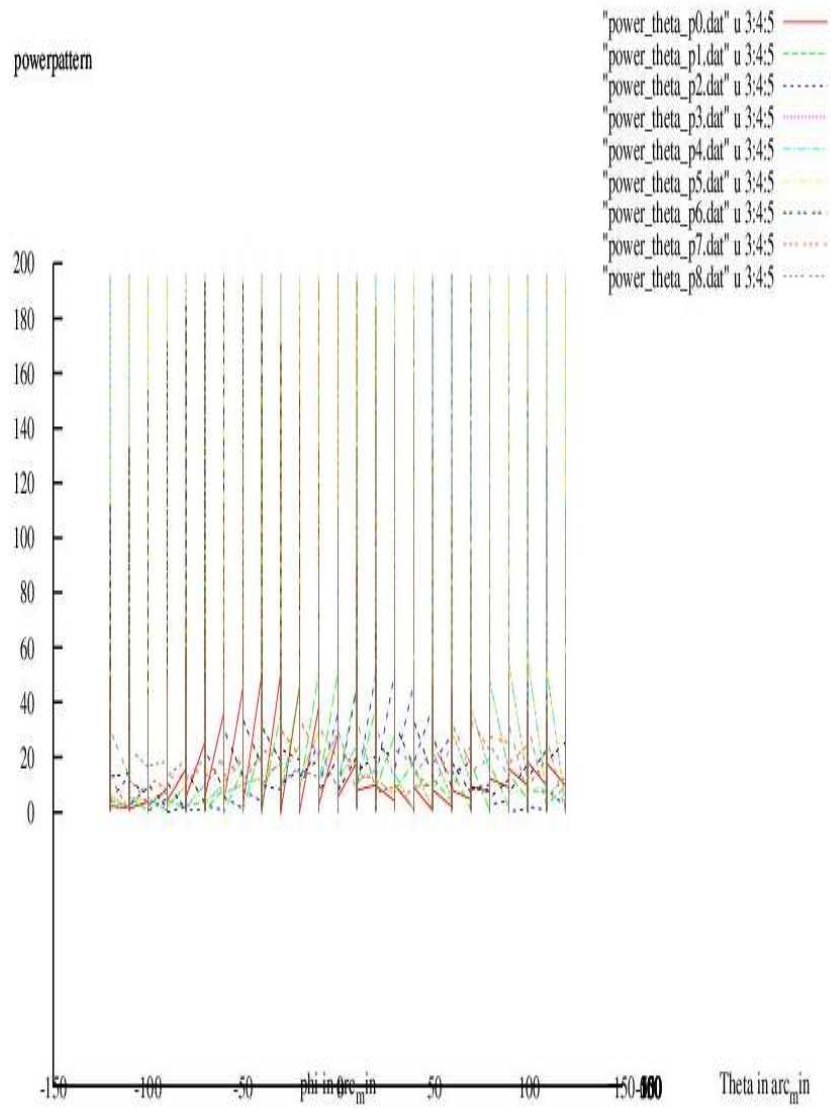


fig 4.5

The plot along phi axis is as shown:

The script file used for contour plot is as given:

```
set term post landscape enhanced colour colourtext "Times NewRoman" 10
set out "test.ps"
set autoscale
set xlabel 'Theta in arc_min'
set ylabel 'phi in arc_min'
set zlabel 'powerpattern'
set title 'Powerpattern Vs Theta For GMRT array phased in any direction in theta phi
plane , showing contour' font "Times NewRoman, 12"
show title
set time
set grid
set contour
set dgrid3d 90,50,8
set cntrparam levels 2
set cntrparam levels incremental .9,.1,1
splot "power_theta_p0.dat" u 3:4:($5/196) w l, "power_theta_p1.dat" u 3:4:($5/196)
w l,"power_theta_p2.dat" u 3:4:($5/196) w l,"power_theta_p3.dat" u 3:4:($5/196) w
l,"power_theta_p4.dat" u 3:4:($5/196) w l,"power_theta_p5.dat" u 3:4:($5/196) w
l,"power_theta_p6.dat" u 3:4:($5/196) w l,"power_theta_p7.dat" u 3:4:($5/196) w
l,"power_theta_p8.dat" u 3:4:($5/196) w l
```

The contour plot is as shown:

1) For 50% power levels:

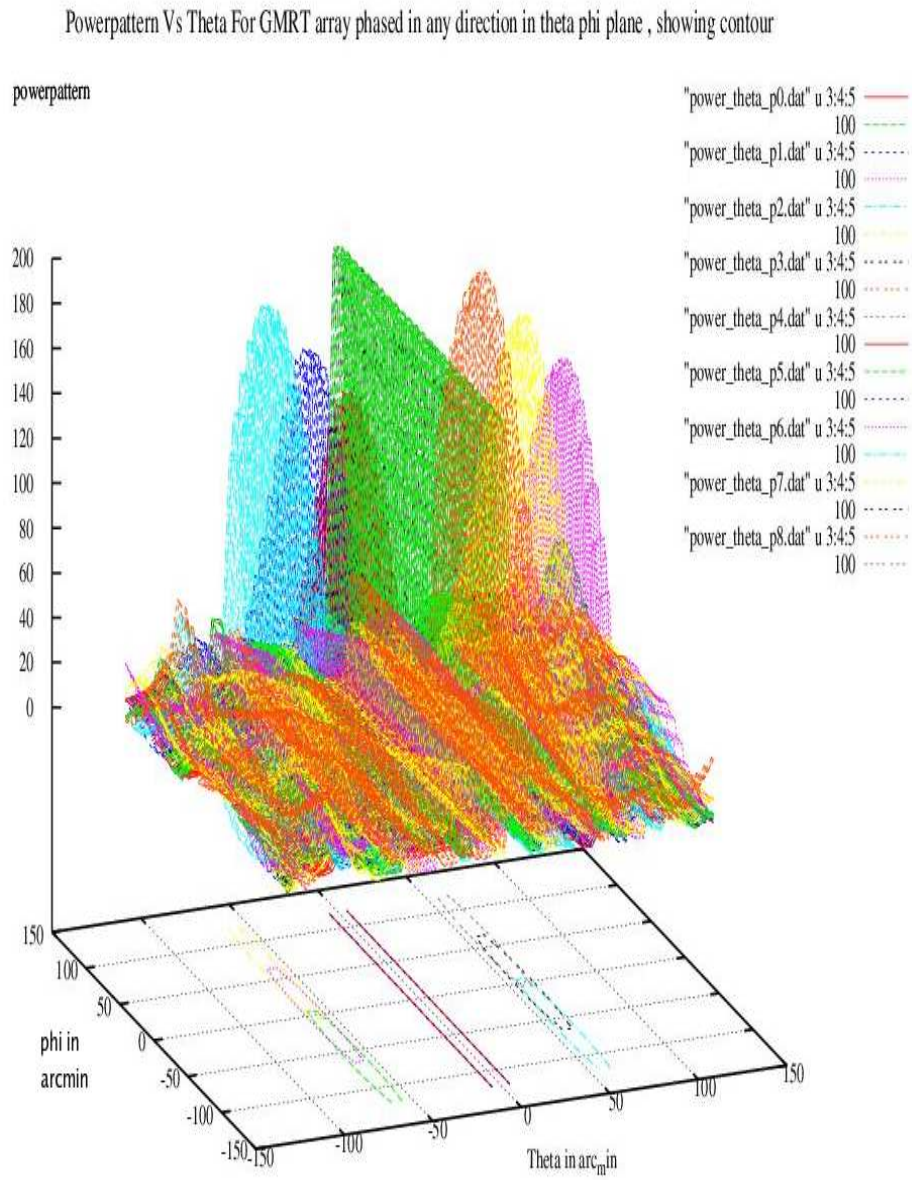


fig 4.6

2) Contour with 90% power levels

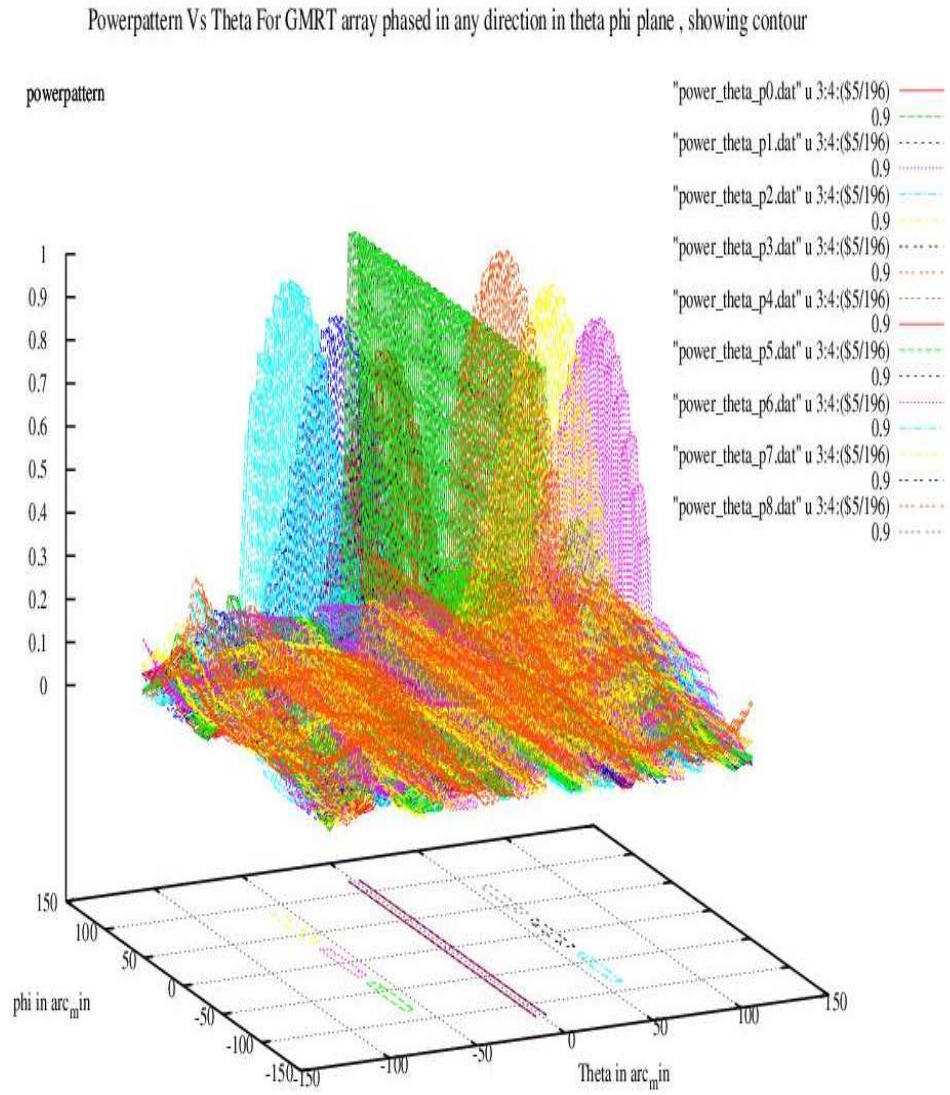


fig 4.7

With unset surface:

Powerpattern Vs Theta For GMRT array phased in any direction in theta phi plane , showing contour

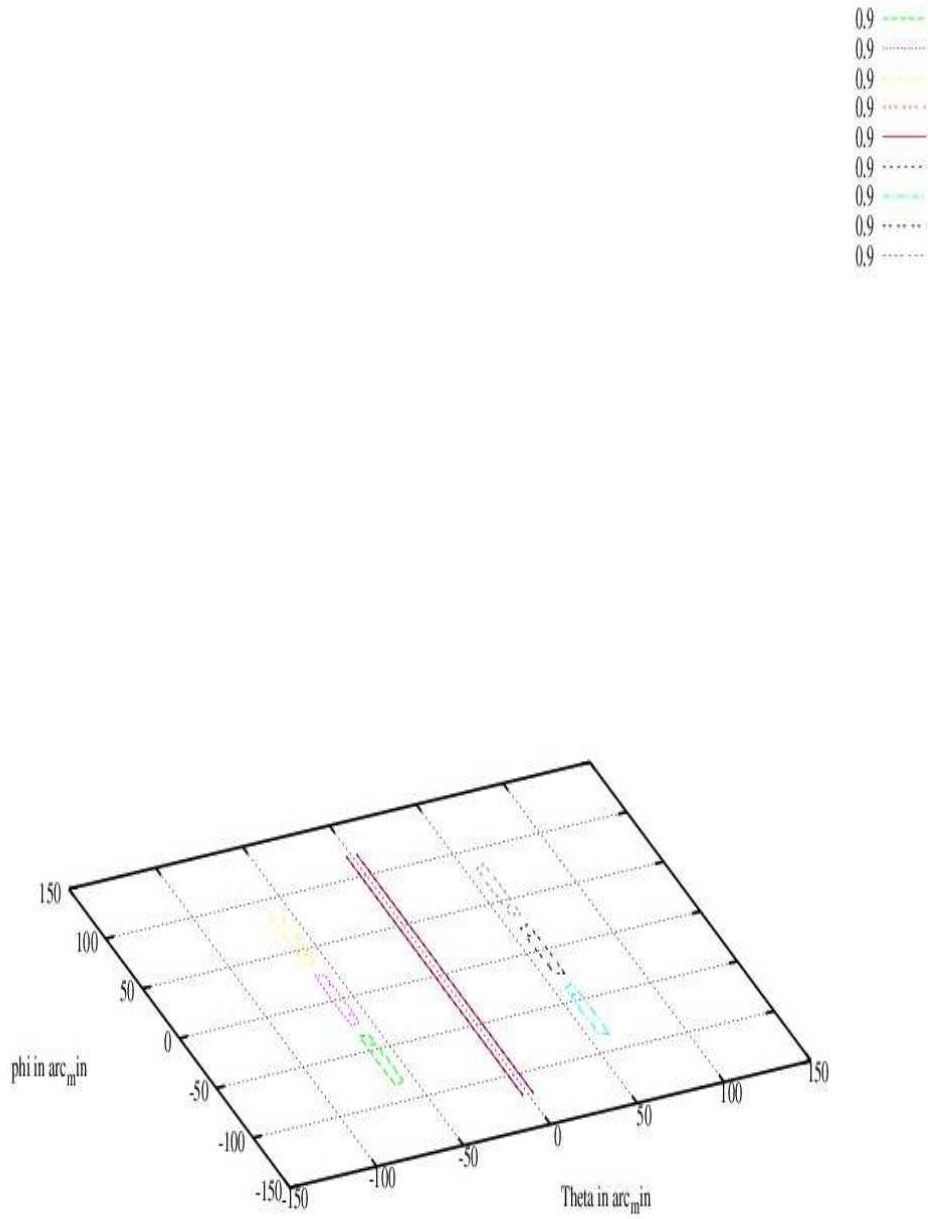


fig:4.8

2a)Here,

theta_p=71deg,phi_P=0deg

theta_0=71deg,phi_0=0deg.

range of theta=0.5deg, range of phi = 0.5 deg

accuracy in theta=0.2' ,accuracy in phi=0.2'

No. of multiple beams to be generated in theta and phi= 2

separation in between beams =2'. (both in theta and phi direction)

The input data file is as shown: ("input_thp_phip.dat")

# theta_P	phi_P
4257.000000	-3.000000
4257.000000	0.000000
4257.000000	3.000000
4260.000000	-3.000000
4260.000000	0.000000
4260.000000	3.000000
4263.000000	-3.000000
4263.000000	0.000000
4263.000000	3.000000

Powerpattern Vs Theta For GMRT array phased in any direction in theta phi plane , showing contour

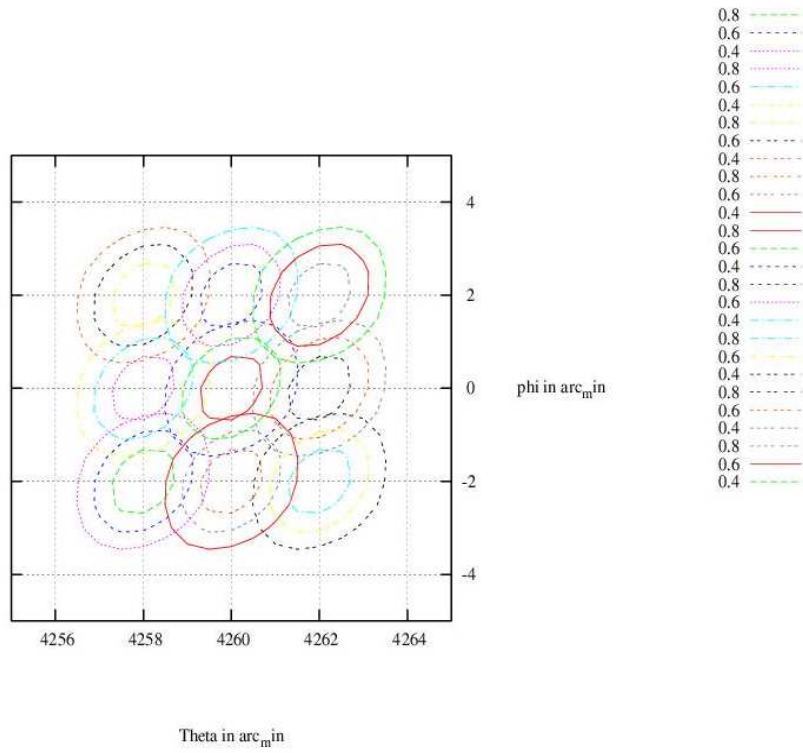


Fig 4.9

2b) $\theta_p=71^\circ, \phi_P=0^\circ$
 $\theta_0=71^\circ, \phi_0=0^\circ$.
 range of $\theta=0.2^\circ$, range of $\phi = 0.2^\circ$
 accuracy in $\theta=0.2'$, accuracy in $\phi=0.2'$
 No. of multiple beams to be generated in θ and $\phi=2$
 separation in between beams $=5'$. (both in θ and ϕ direction)

Powerpattern Vs Theta For GMRT array phased in any direction in theta phi plane , showing contour

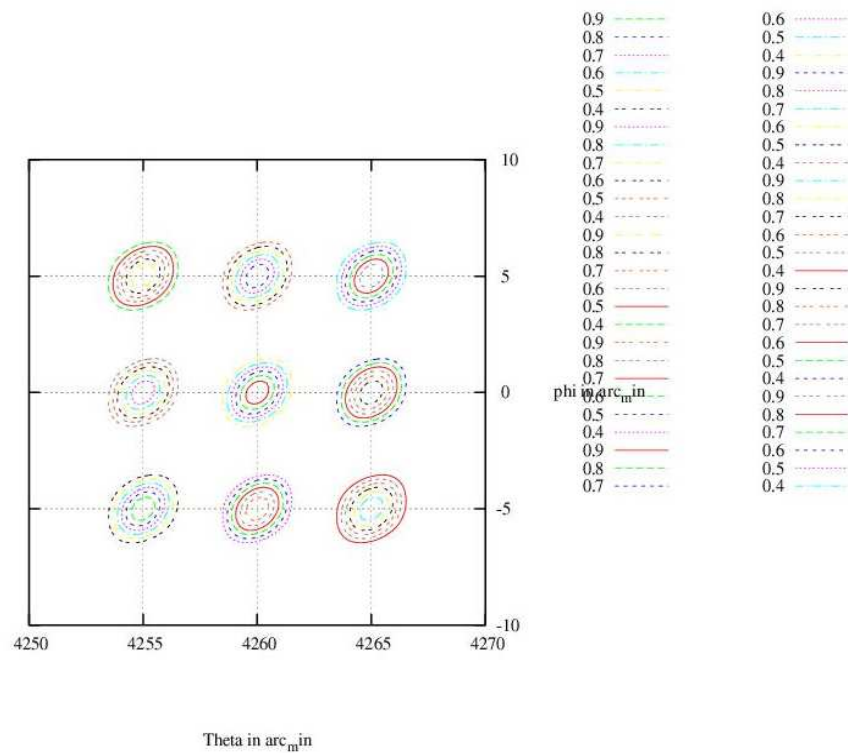


Fig 4.10

- 2c) $\theta_p=71\text{deg}, \phi_P=0\text{deg}$
 $\theta_0=71\text{deg}, \phi_0=0\text{deg}$.
 range of $\theta=0.2\text{deg}$, range of $\phi = 0.2\text{deg}$
 accuracy in $\theta=0.2'$, accuracy in $\phi=0.2'$
 No. of multiple beams to be generated in θ and $\phi=2$
 separation in between beams $=3'$. (both in θ and ϕ direction)

Powerpattern Vs Theta For GMRT array phased in any direction in theta phi plane , showing contour

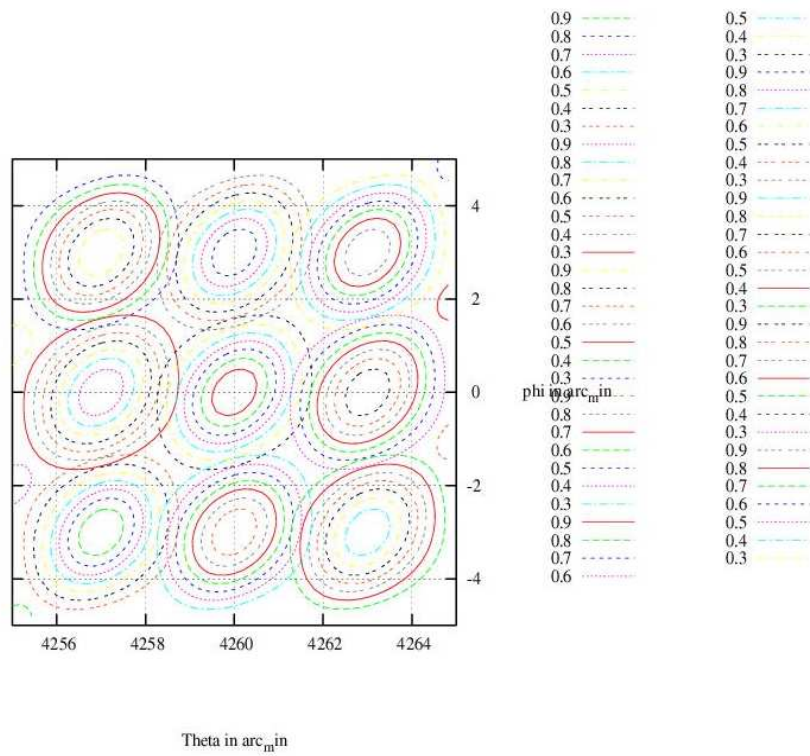


Fig 4.11

The script file for the above three graphs is:

```
set term post landscape enhanced colour colourtext "Times NewRoman" 10
set out "test.ps"
set autoscale
set xlabel 'Theta in arc_min'
set ylabel 'phi in arc_min'
set zlabel 'powerpattern'
set title 'Powerpattern Vs Theta For GMRT array phased in any direction in
theta phi plane , showing contour' font "Times NewRoman, 12"
show title
set time
set grid
set contour
set cntrparam levels 2
set cntrparam levels incremental .3,.1,1
unset surface
set xrange[4255:4265]
set yrange[-5:5]
set view 1,0,1,1
set size square
plot "power_theta_p0.dat" u 3:4:($5/196) w l, "power_theta_p1.dat" u 3:4:
($5/196) w l,"power_theta_p2.dat" u 3:4:($5/196) w l,"power_theta_p3.dat" u
3:4:($5/196) w l,"power_theta_p4.dat" u 3:4:($5/196) w l,"power_theta_p5.dat"
u 3:4:($5/196) w l,"power_theta_p6.dat" u 3:4:($5/196) w
l,"power_theta_p7.dat" u 3:4:($5/196) w l,"power_theta_p8.dat" u 3:4:($5/196)
w l
```

CHAPTER 5

OBSERVATION ON GMRT ARRAY(CENTRAL SQUARE + ARM ANTENNA):

We now extend the phased array of the GMRT from the central square symmetrically to the arm antennas. The power pattern are as observed in the plots below. The plot are taken both for fixed theta and fixed phi.

Here : $\lambda = 1\text{m}$

$$\theta_p = \theta_0 = 45 \text{ deg}$$

$$\phi_p = \phi_0 = 0 \text{ deg}$$

Range of theta = range of phi = 3 deg

Accuracy of theta = Accuracy of phi = 1'

1) For 17 antenna coordinates (Central Square + E02 +S01+W01)

a)

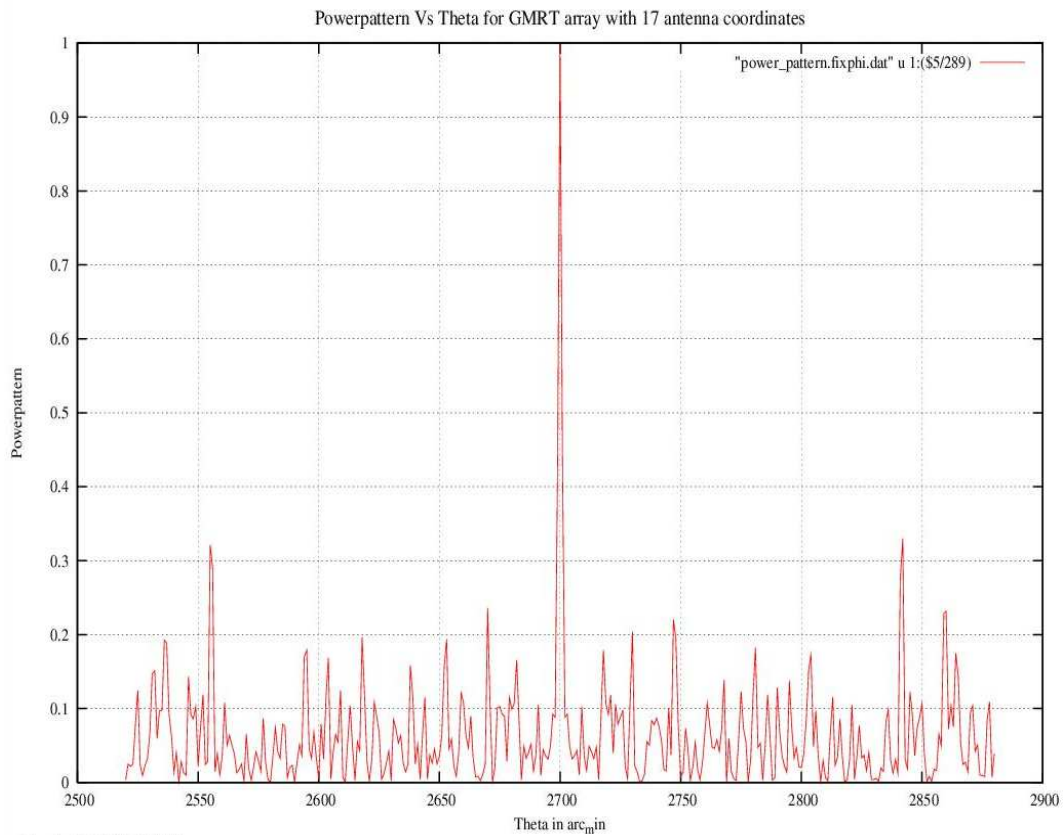


fig5.1

1 b)

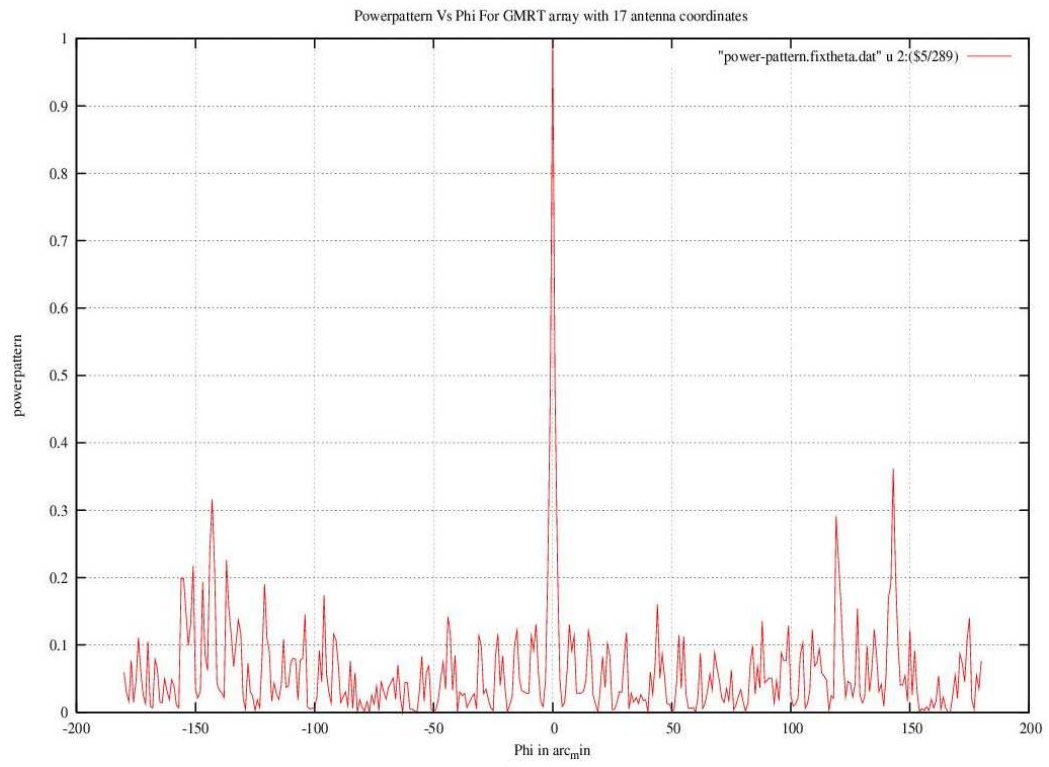


fig5.2

2) For 20 antenna coordinates (Central Square + E02 +E03+S01+S02+W01+W02)

a)

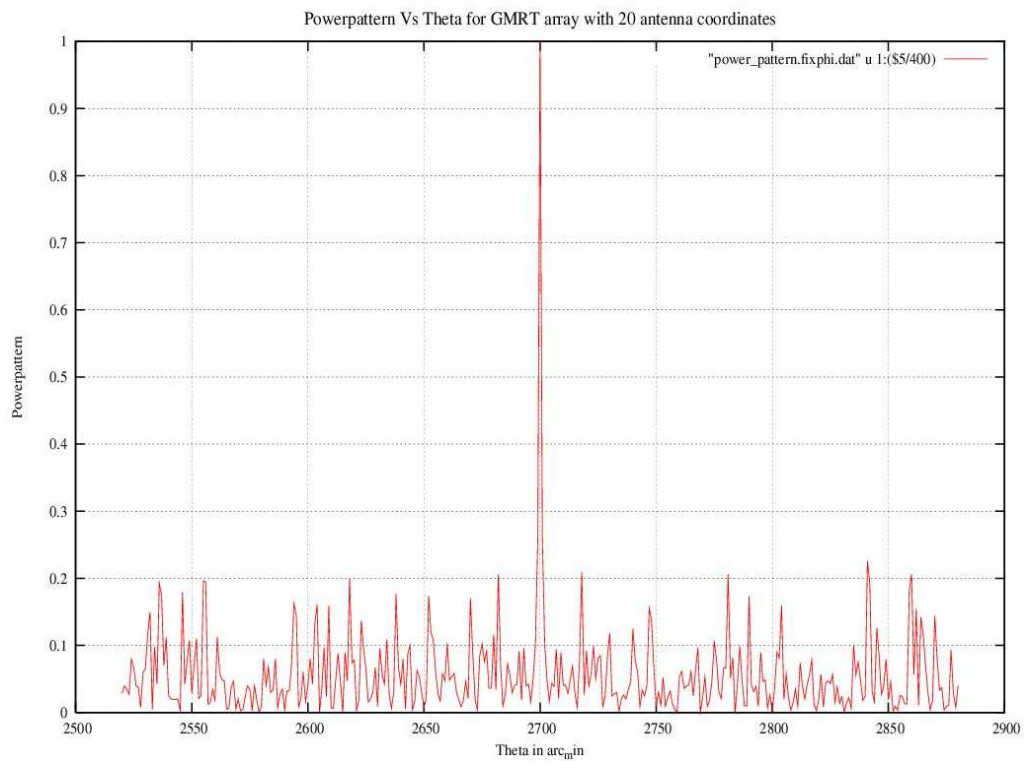


fig5.3

2b)

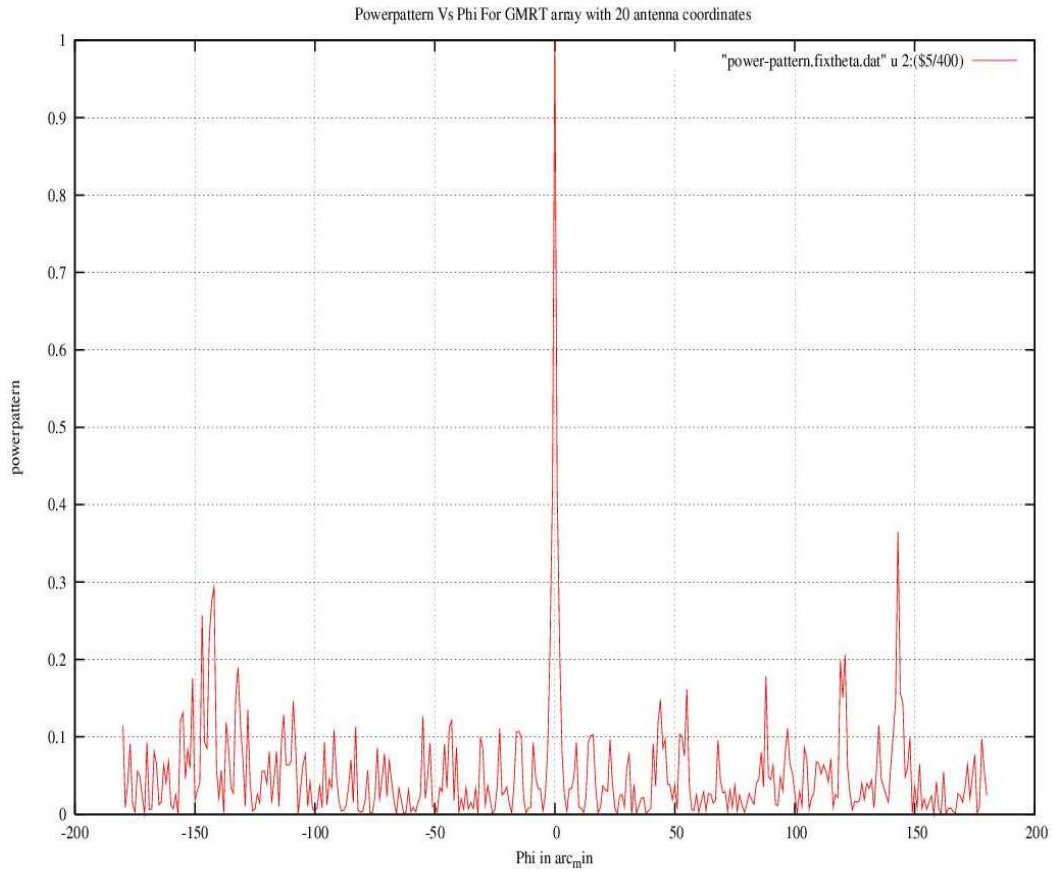


fig5.4

3)For 23 antenna coordinates (Central Square + E02
E03+E04+S01+S02+S03+W01+W02+W03)

a)

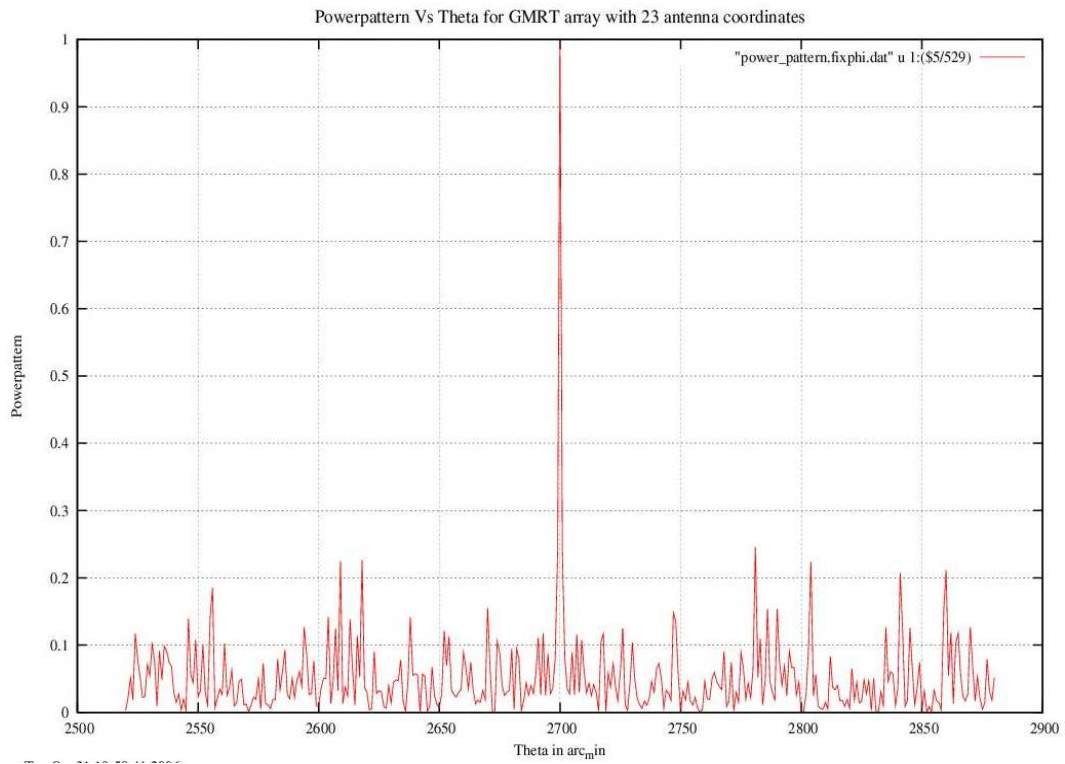


fig5.5

3 b)

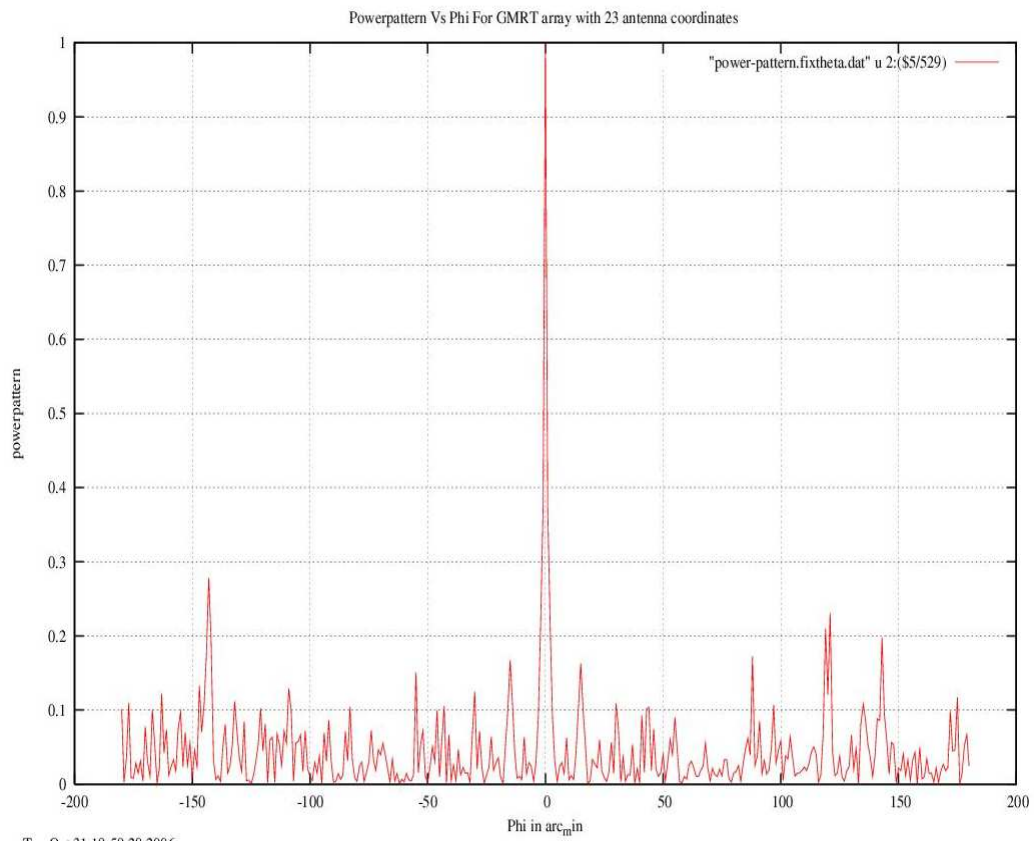


fig5.6

4)For 26 antenna coordinates (Central Square + E02
+E03+E04+E05+S01+S02+S03+S04+W01+W02+W03+W04)
a)

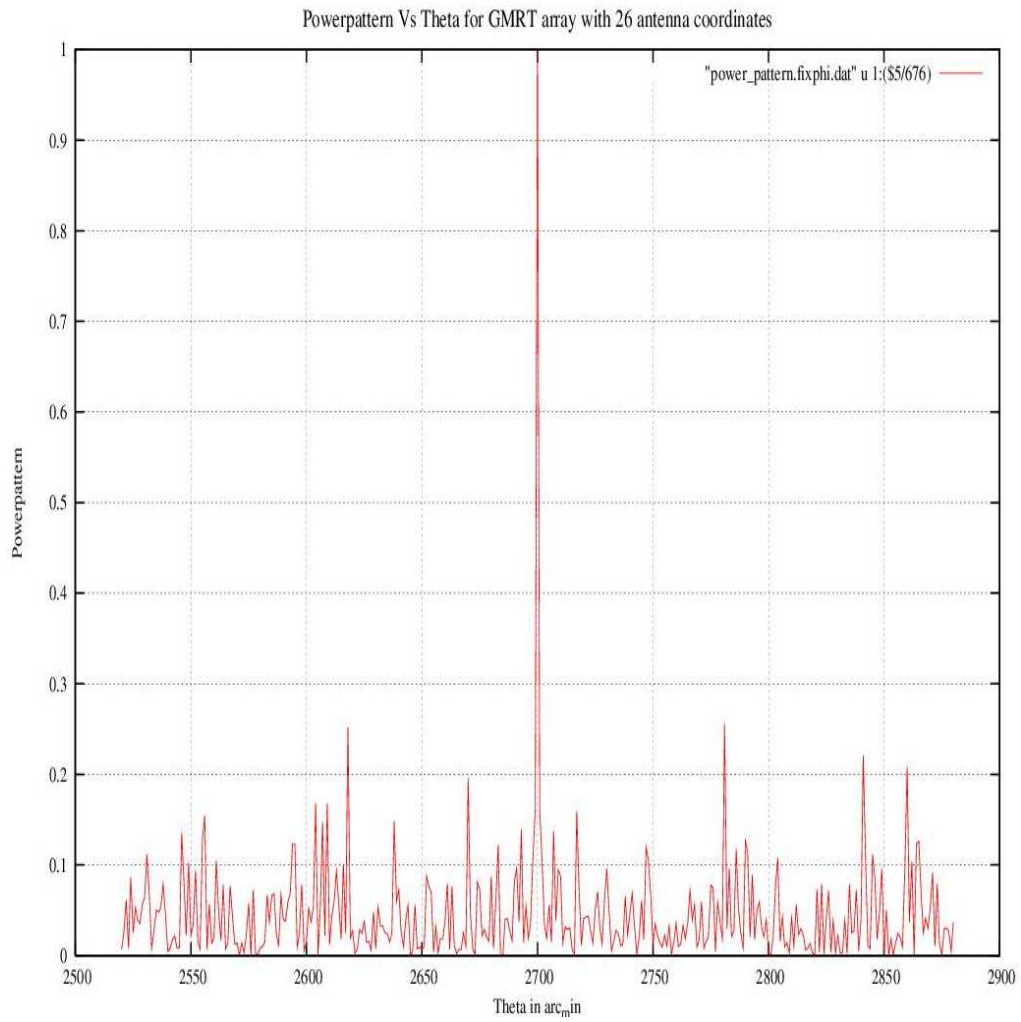


fig5.7

4 b)

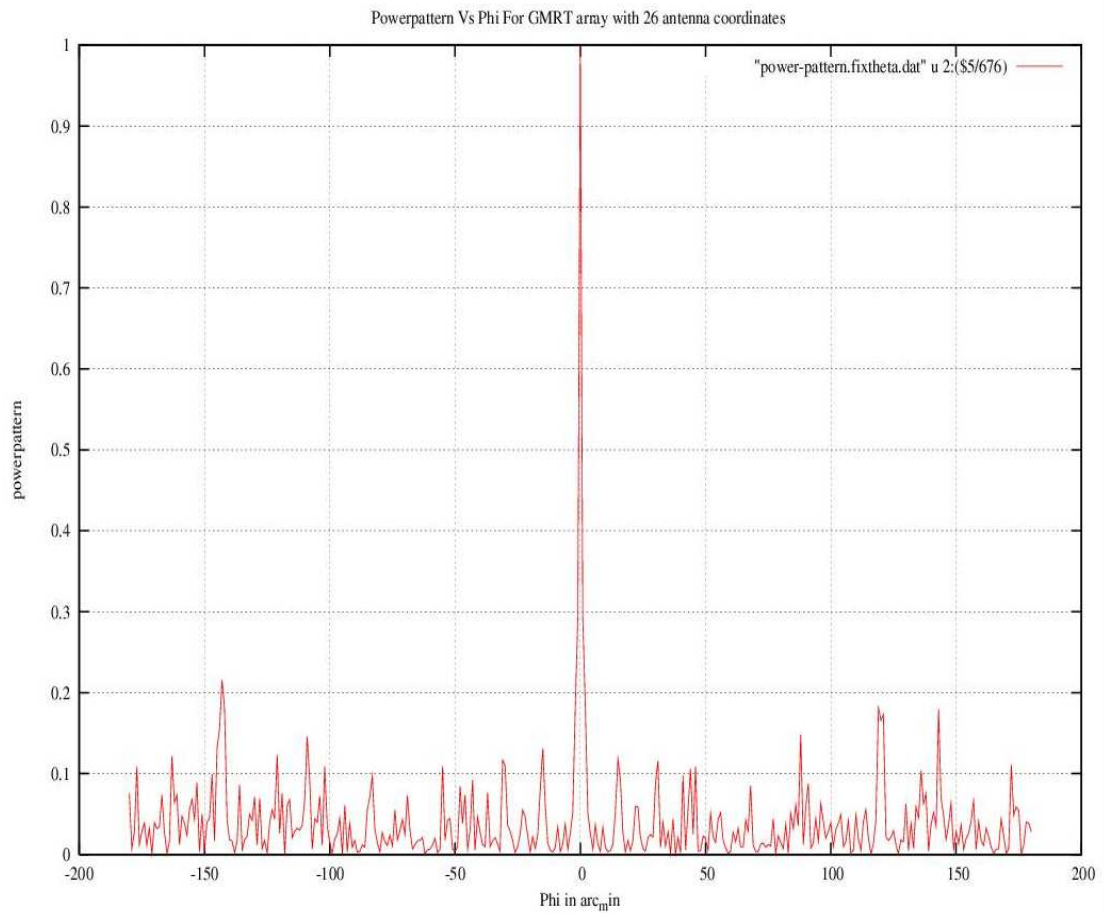


fig5.8

5) For 29 antenna coordinates (Central Square + E02
+E03+E04+E05+E06+S01+S02+S03+S04+S05+W01+W02+W03+W04+W05)

a)

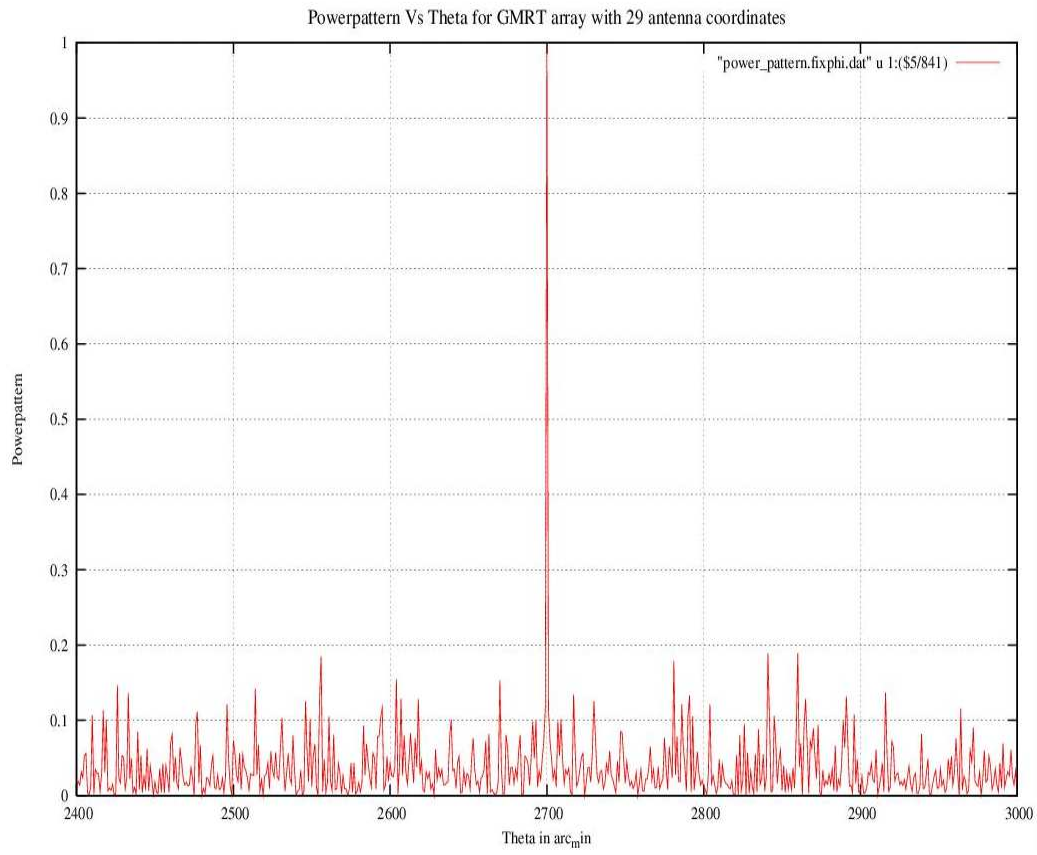


fig5.9

5 b)

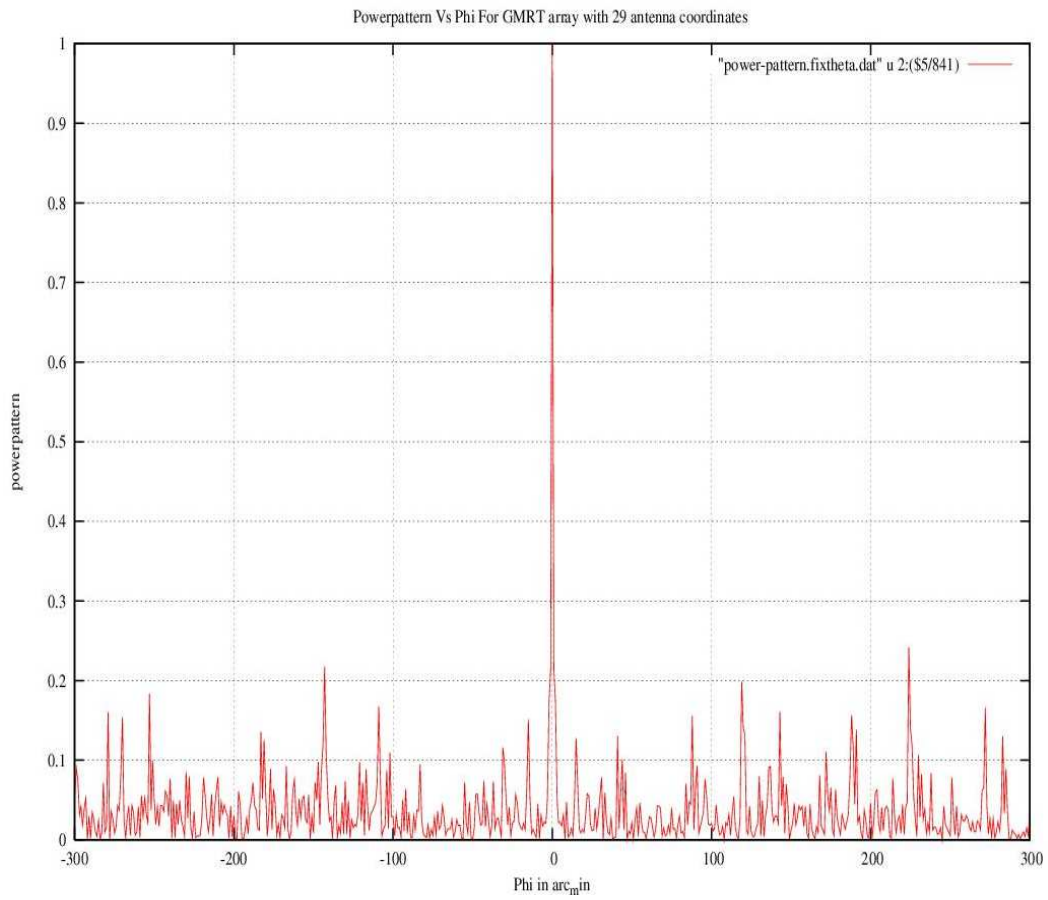


fig5.10

From the above plots it is noted that on increasing the phased antenna from the central square to the arm antenna there is not much change in the added noise levels which remains below 10%.

The Half Power Beam Width of the phased beams are as follows:

1) For 17 antenna coordinates : HPBW = 1.61'

Range of theta = 0.2 deg

Accuracy of theta = 0.05 '

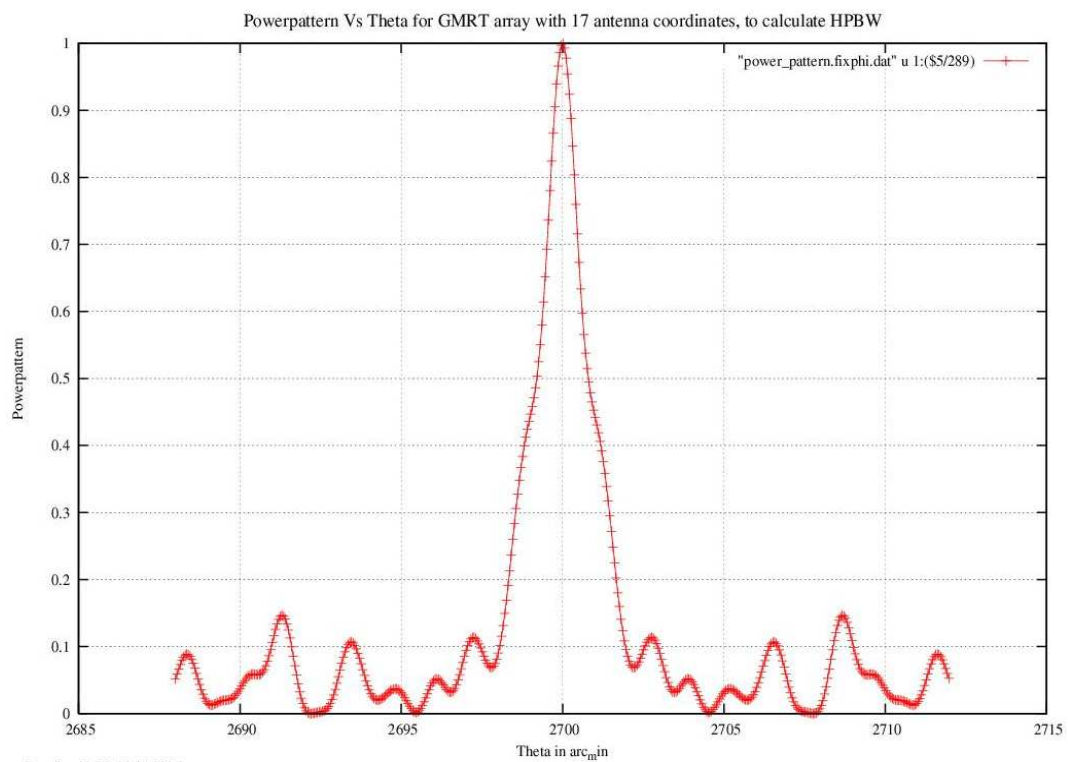


fig5.11

2) For 20 antenna coordinates : HPBW = 0.84 '

Range of theta = 0.2 deg

Accuracy of theta = 0.05 '

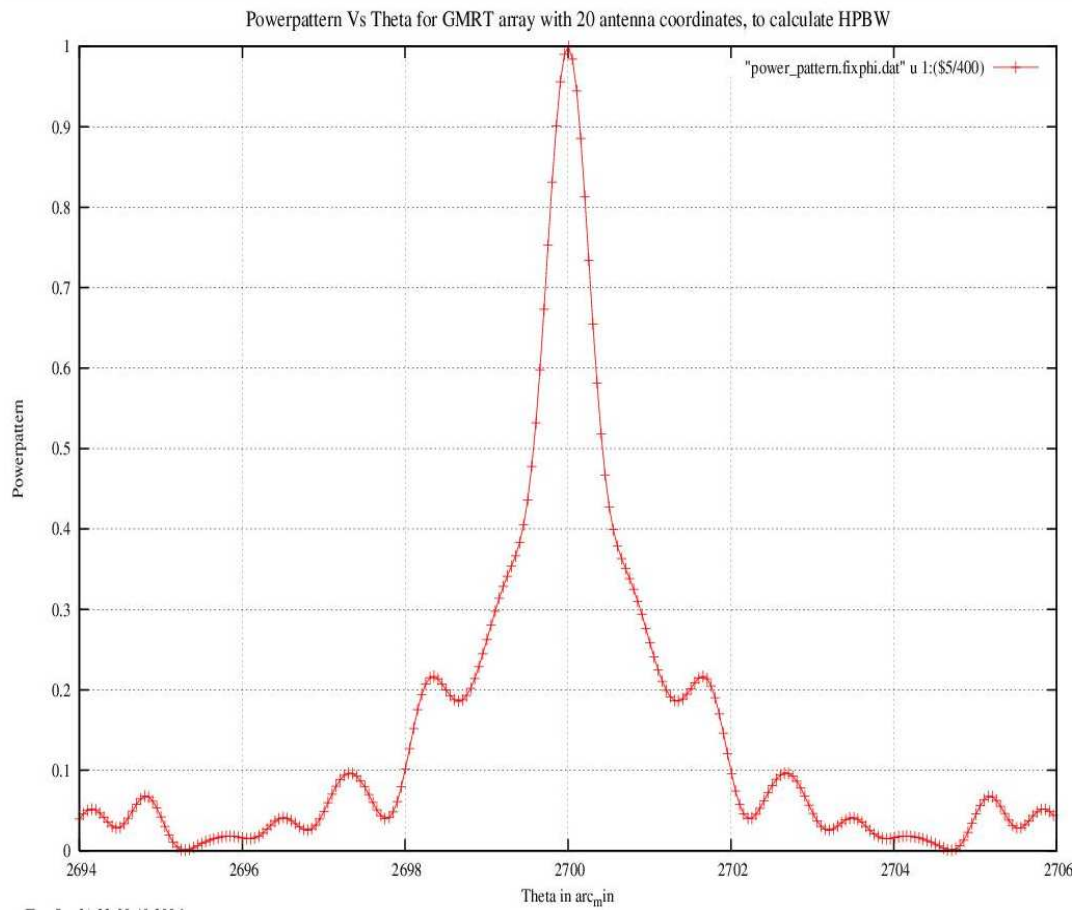


fig5.12

3) For 23 antenna coordinates : HPBW = 0.57 '

Range of theta = 0.1 deg

Accuracy of theta = 0 .05 '

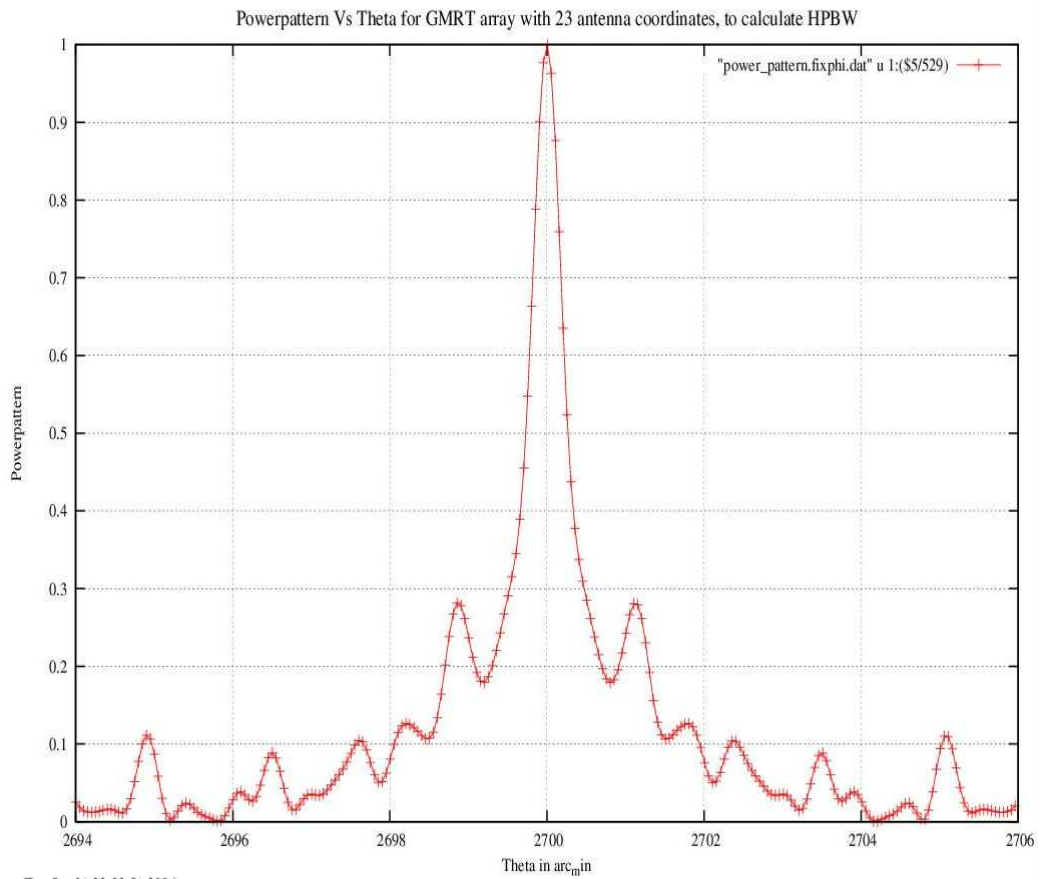


fig5.13

4) For 26 antenna coordinates : HPBW = 0.36 '

Range of theta = 0.05 deg

Accuracy of theta = 0.02 '

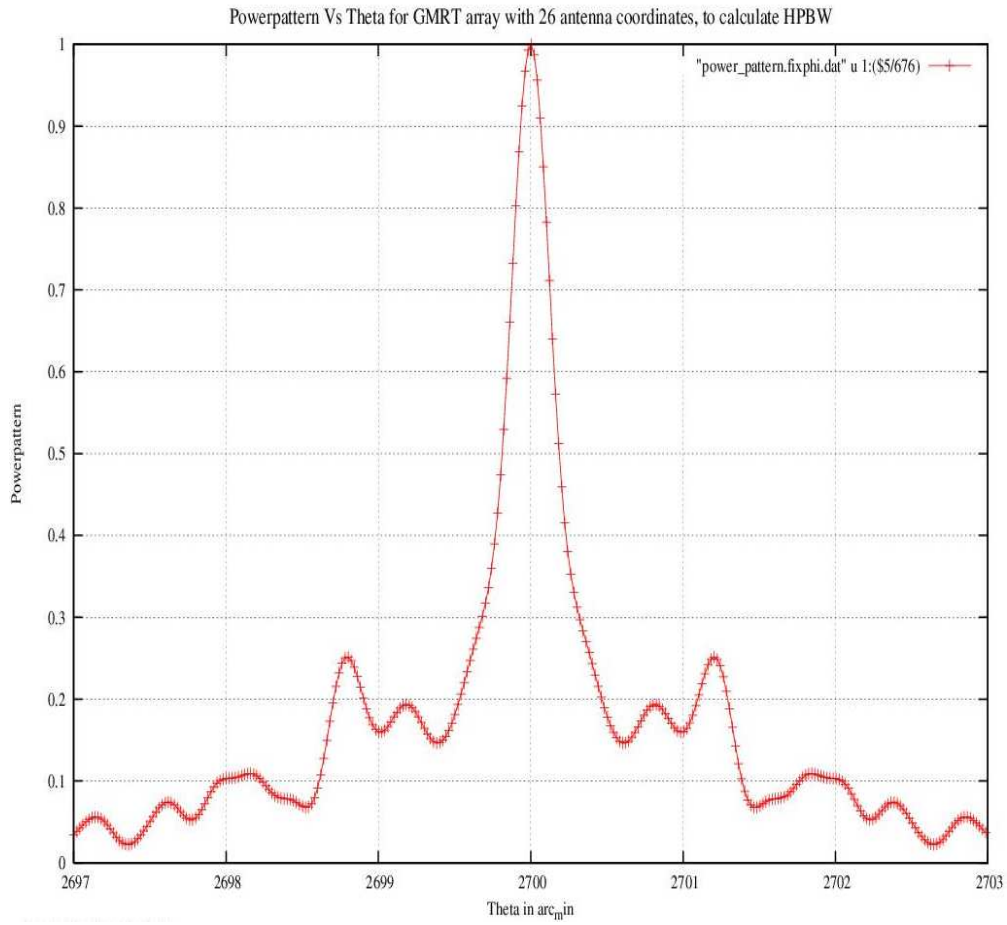


fig5.14

5) For 29 antenna coordinates : HPBW = 0.27 '

Range of theta = 0.02 deg

Accuracy of theta = 0.01 '

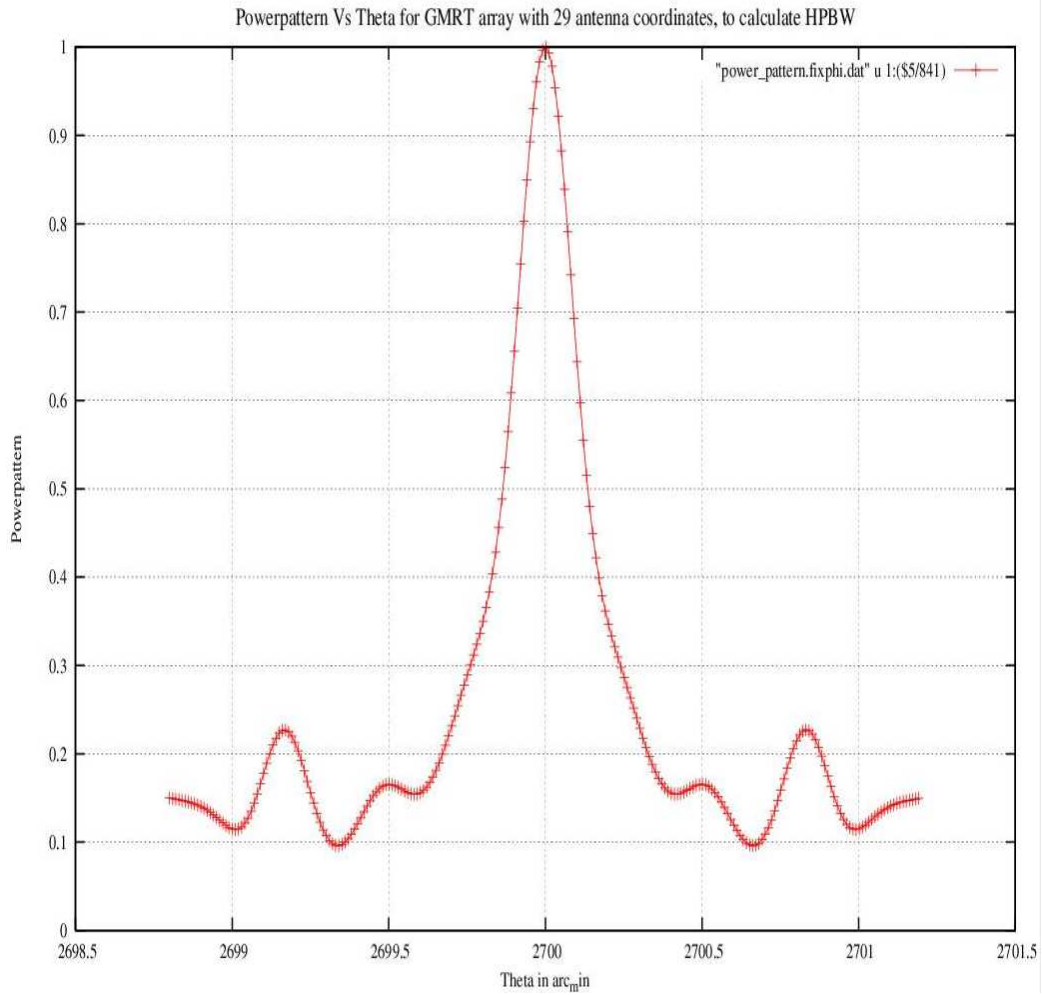


fig5.15

SCOPE OF IMPROVEMENT IN THE CODE

The final code of the program is successful in generating multiple beams in the 2-D sky for any direction of phasing. This program acts as a stand-alone in helping to view how array beam are phased for any linear or random distributed array of antenna.

- 1) The code can be improved to be made more user friendly in the task of entering the input data for calculation of the beam pattern. This can be done by writing a script file which will have the inputs given as arguments. Thus on running the script file each time the power gets calculated., and the user need not enter the same inputs time and again.
- 2) The equation for calculating the x component,y component,z component of the power pattern can be theoretically simplified . The incremental phase factors can be incorporated as a linear term over and above the existing phase factor. This makes the algorithm and the calculation of power factor much simpler.
- 3) The user may need to phase the beams finite beam width away from the central beam. The algorithm need to be modified accordingly

APPENDIX

LIST OF VARIABLES AND FLOWCHART

STRING VARIABLE:

fname[100] : to hold a variable file name

INTEGER VARIABLES:

n : Reference antenna co-ordinate

m : Number of antenna chosen

no_th : Number of beams to be generated along theta axis around phasing direction

no_phi : Number of beams to be generated along phi axis around phasing direction

s : Running index to demarcate the files where the generated multiple beam data is written

i : Running index for multiple use

j[30] : Array to hold the coordinates of the antenna selected in the sequence

cnt , count : Index used to obtain the antenna coordinates of the selected antenna

inp1 : Index for the array th_P when writing into the input data file

inp2 : Index for the array phi_P when writing into the input data file

inp3 : Index for the array th_P when reading into the input data file

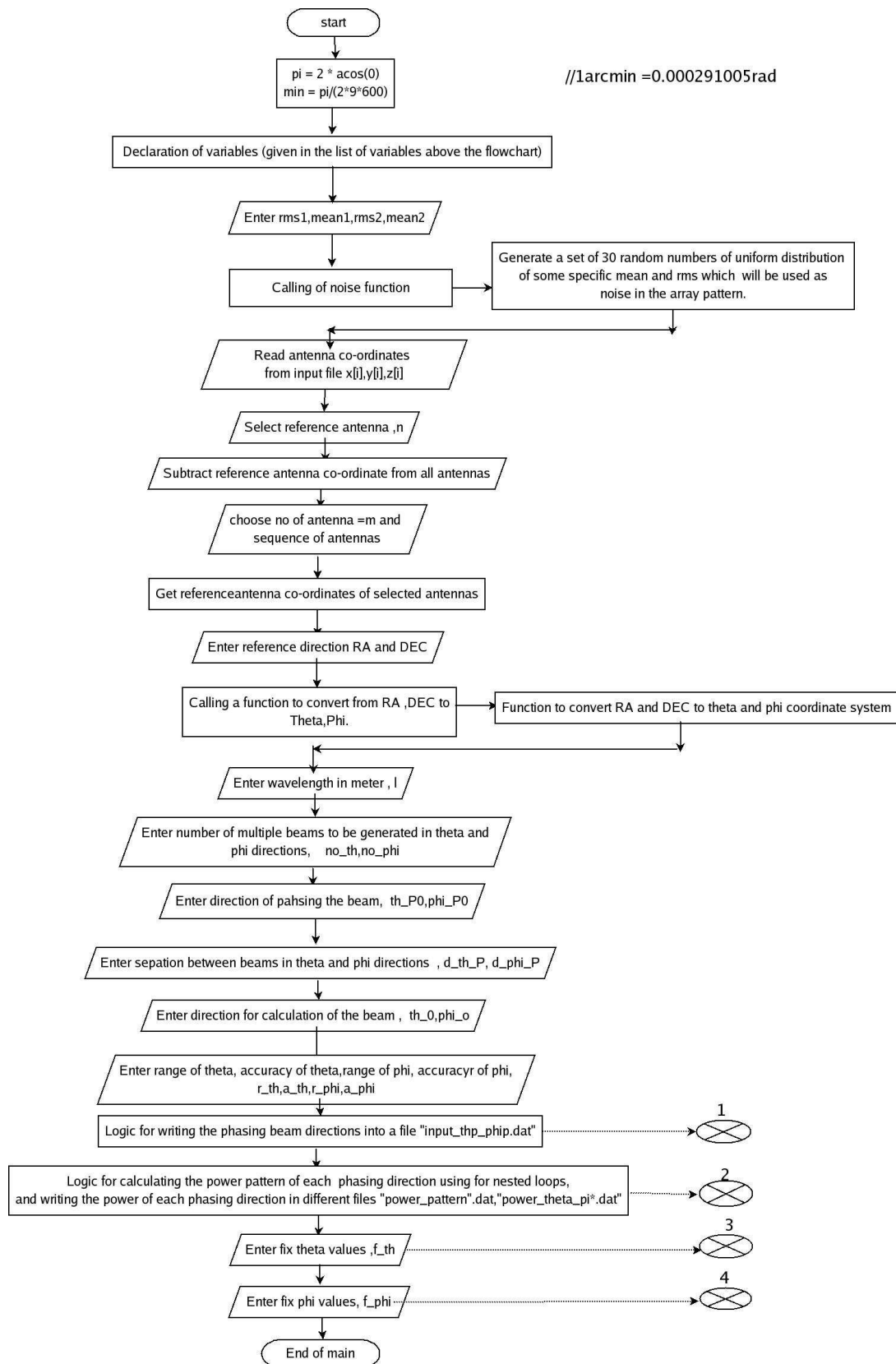
inp4 : Index for the array phi_P when writing into the input data file

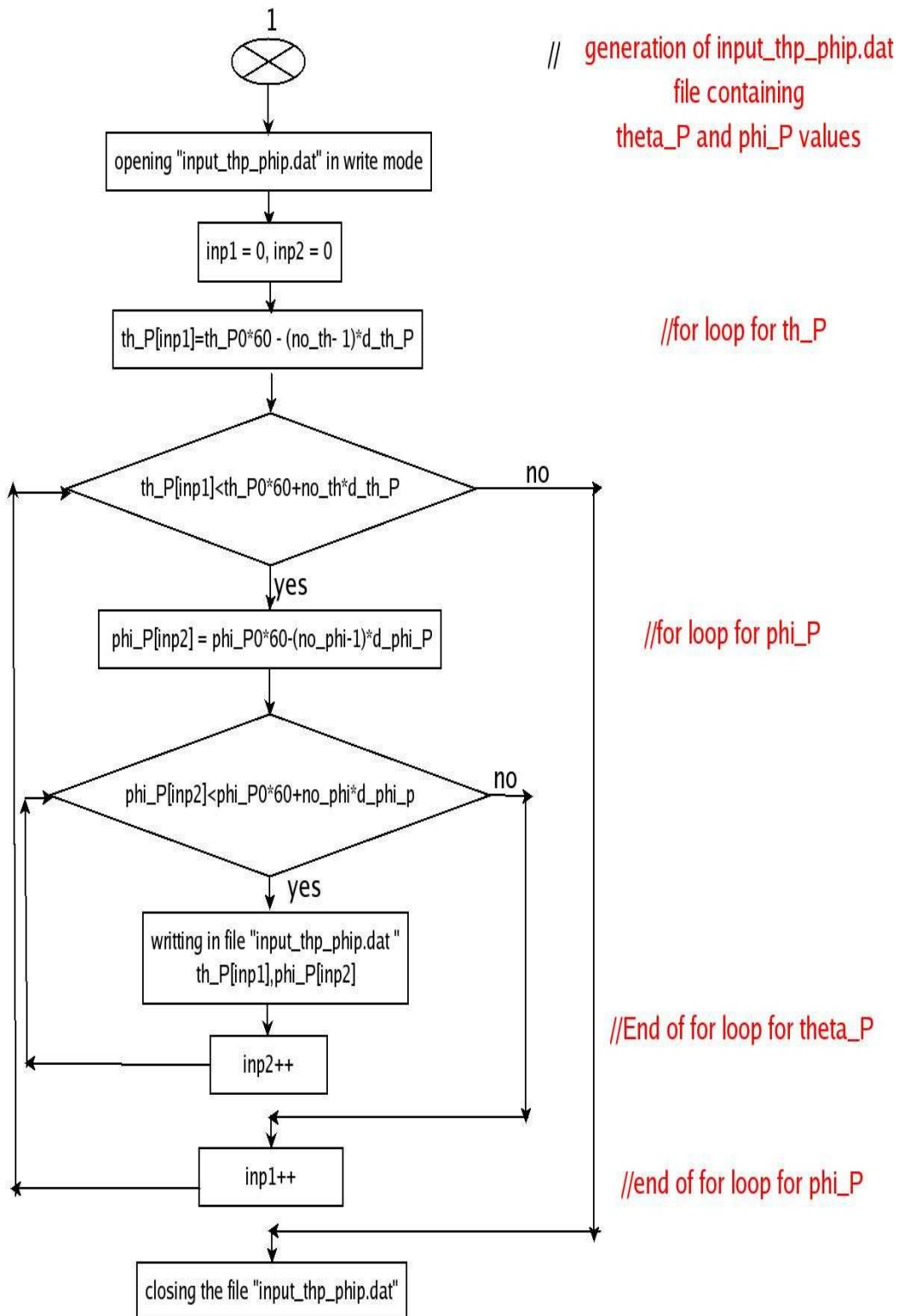
FLOAT VARIABLES

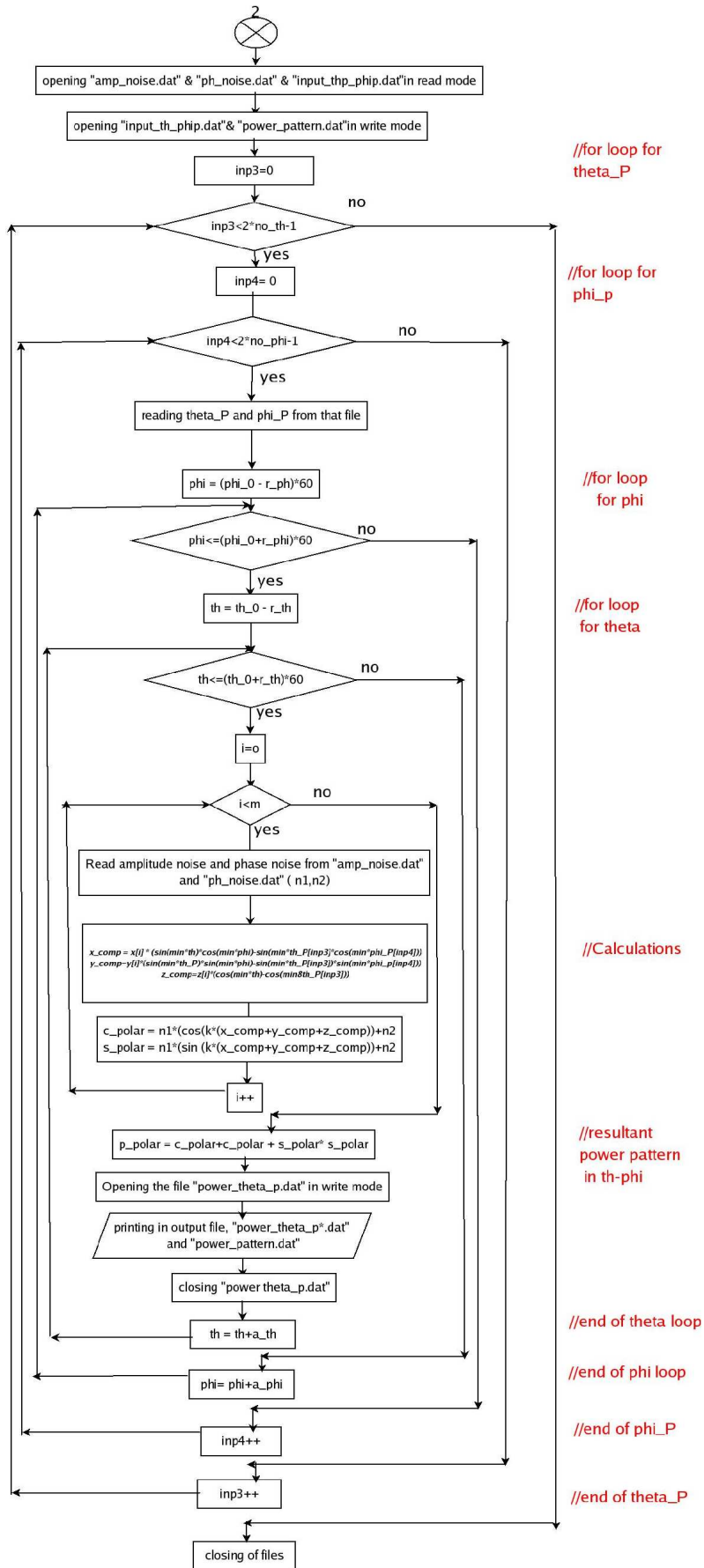
x[30] , y[30] , z[30] : Arrays to hold the coordinates of the antenna
x1[30] , y1[30] , z1[30] : Holds the antenna coordinates subtracted from the
reference antenna coordinates
l : Wavelength
th_P0 : Phasing direction of the beam in theta
phi_P0 : Phasing direction of the beam in phi
th_P[20] : Array to hold multiple value of phasing theta
phi_P[20] : Array to hold multiple value of phasing phi
d_th_P : Separation between phasing beams in direction of theta
d_phi_P : Separation between phasing beams in direction of phi
th_0 : Direction in theta about which beam pattern is calculated
phi_0 : Direction in phi about which beam pattern is calculated
th : Running value of theta according to the loop condition
phi : Running value of phi according to the loop condition
r_th : Range of theta
a_th : Accuracy of theta
r_phi : Range of phi
a_phi : Accuracy of phi
f_th : Fixed theta value ,the phi variation is noted keeping this theta fixed
f_phi : Fixed phi value ,the theta variation is noted keeping this phi fixed
rms1 : RMS value of amplitude of noise
mean1 : Mean value of amplitude of noise
rms2 : RMS value of phase of noise
mean2 :Mean value of phase of noise
RA(RA_hr,RA_min,RA_sec) ,DEC(DEC_deg,DEC_min,DEC_sec) :
Coordinates in the RA-DEC system

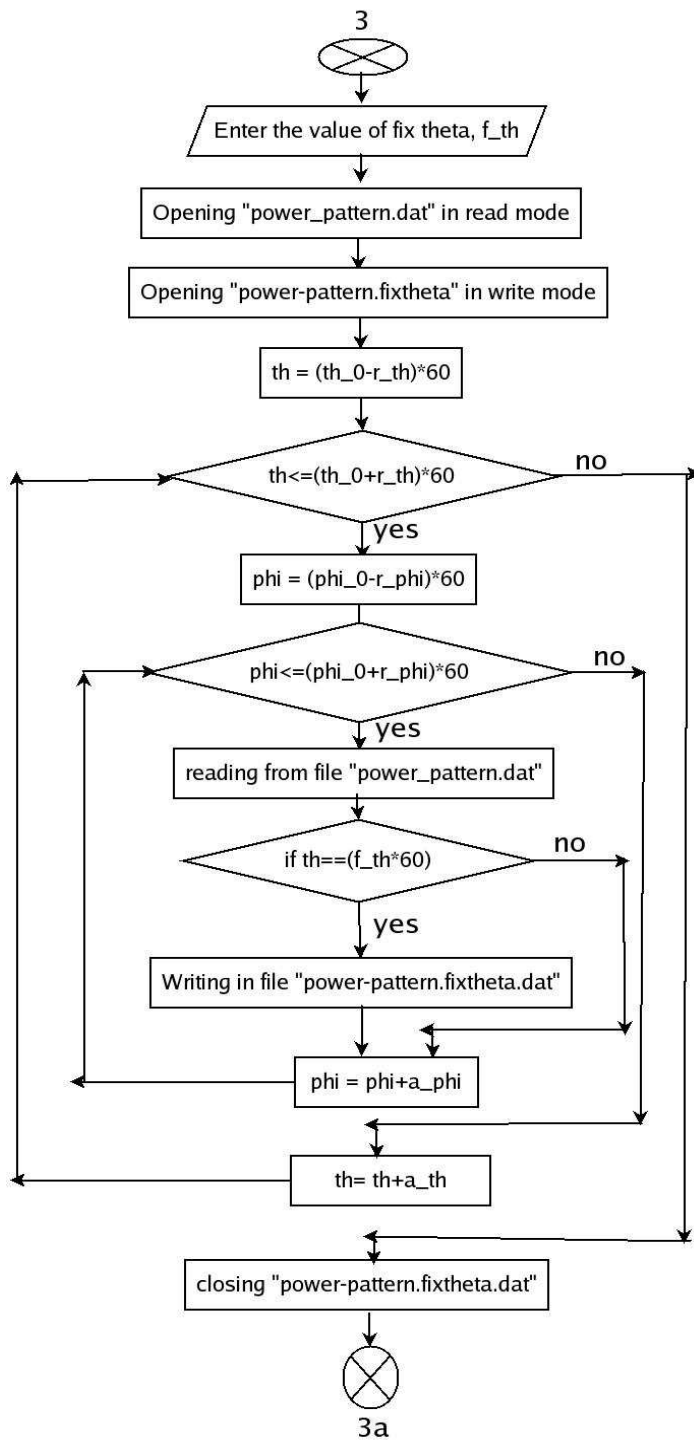
FLOWCHART:

1) The flowchart of the program as an algorithm is given below:







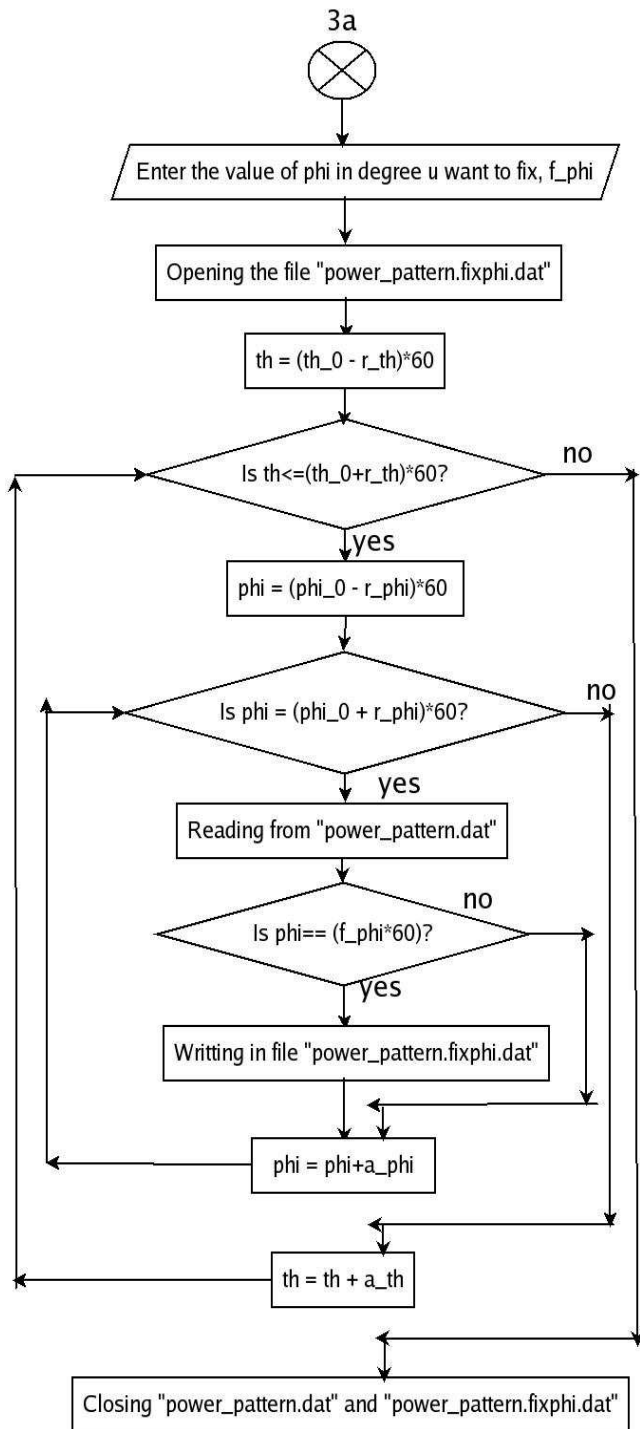


//loop for theta

//loop for phi

//end of phi loop

//end of theta loop



//loop for theta

//loop for phi

//end of phi loop

//end of theta loop

gmrt_arraypattern.c

COMMENTS::

1. Endfire did not work.
2. Beam phasing in one direction and calculation in another direction is not possible.
3. Multiple beam generation is not possible here.

STEPS TO ACHIEVE THE FINAL CODE::

1. Multiple beam generation concept brought in.
2. Initially in three specific directions of theta the beam is phased which is supplied by the programmer and not by the user. The power pattern is written in three individual files and this file generation is not flexible.

PROGRAM NAME :: **arraypattern_version1.c**

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```
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/*COMMENTS:: 1.PHASING OF BEAM FOR ENDFIRE CASE IS POSSIBLE HERE.
              2.CONCEPT OF PHASING THE BEAM AND CALCULATION OF THE PATTERN IN
              ANY DIRECTION BROUGHT IN.

/* Make a program which will calculate the phased array beam in the presence of
noise */

#include<stdio.h>
#include <math.h>

#define pi 2*acos(0)
#define min pi/(2*9*600) //1 arc-minute = 0.000291005 rad

int RA_DEC_to_th_phi(float RA_hr,float RA_min,float RA_sec,float DEC_deg,float DEC_min,float DEC_sec);
int noise(float rms1,float rms2,float mean1,float mean2);

int main()
{
    int n,m;
    float x[30],y[30],z[30],l,th_P,phi_P,th_0,phi_0,r_th,a_th,r_phi,a_phi,f_th,f_phi,
    rms1,mean1,rms2,mean2,RA_hr,RA_min,RA_sec,DEC_deg,DEC_min,DEC_sec;

    system("clear");
    printf("enter the rms-s and means you want for uniformly distributed 30 random numbers will be used as noise to the beam pattern corresponding to amplitude and phase respectively ");
    scanf("%f%f%f%f",&rms1,&mean1,&rms2,&mean2);
    noise(rms1,rms2,mean1,mean2);

    // Read the antenna co-ordinates from the input file
    FILE *ifp = NULL;
    int i;
    float x1[30], y1[30], z1[30];

    ifp = fopen("input.dat","r"); //input.dat is the input file for antenna co-ordinates in x,y,z plane.
    for(i=0;i<30;i++)
        fscanf(ifp,"%f%f%f\n",&x[i],&y[i],&z[i]);
    fclose(ifp);

    // Select the reference antenna co-ordinate from the list of the antennas
    printf("choose the antenna number you want as reference n = ");
    scanf("%d",&n);
    printf("the antenna co-ordinates you select as reference are\n");
    printf("%5.2f %5.2f %5.2f\n",x[n-1],y[n-1],z[n-1]);

    // Subtract the ref antenna co-ordinate from the all antennas
    printf("\nthe corrected co-ordinates in meters are:\n");
    for(i=0;i<30;i++)
    {
        x1[i] = x[i] - x[n-1];
        y1[i] = y[i] - y[n-1];
        z1[i] = z[i] - z[n-1];
        printf("%5.2f %5.2f %5.2f\n",x1[i],y1[i],z1[i]);
    }

    // Get the No of antenna as well as the antenna nos
    printf("choose the number (must be <= 30) of antennas m = ");
    scanf("%i",&m);
}
```

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```

int j[30];
printf("\n choose the sequence of antennas. You must input the reference antenna.\n");
for(i=0;i<m;i++)
{
scanf("%i",&j[i]);
printf("%i\n",j[i]);
}

// Get the antenna co-ordinates of the selected antennas
int cnt;
int count = 0;
printf("the corresponding antenna co-ordinates are\n");
for(i=0;i<30;i++)
{
for(cnt=0;cnt<m;cnt++)
{
if(i == j[cnt])
{
printf("%5.2f %5.2f %5.2f\n",x1[i],y1[i],z1[i]);
x[count] = x1[i]; y[count] = y1[i]; z[count] = z1[i];
count++;
}
}
}
printf("enter the reference direction (in sky) in RA(RA_hr,RA_min,RA_sec) and DEC(DEC_deg,DEC_min,DEC_sec) co-ordinate system");
scanf("%f %f %f %f %f %f", &RA_hr, &RA_min, &RA_sec, &DEC_deg, &DEC_min, &DEC_sec);
RA_DEC_to_th_phi(RA_hr, RA_min, RA_sec, DEC_deg, DEC_min, DEC_sec);

printf("\nenter the wavelength l in meter = ");
scanf("%f", &l);

printf("enter Theta_P and Phi_P which specify the direction for phasing the array ");
scanf("%f %f", &th_P, &phi_P);

printf("enter Theta_0 and Phi_0 which specify the direction for calculation of pattern ");
scanf("%f %f", &th_0, &phi_0);
printf("\nenter the range you want in degree and the accuracy you want in arc-minute in the direction of theta or polar angle\n");
scanf("%f %f", &r_th, &a_th);
printf("\nenter the range you want in degree and the accuracy you want in arc-minute in the direction of phi or azimuth\n");
scanf("%f %f", &r_phi, &a_phi);

FILE *ifp2;
FILE *ifp3;
FILE *ofp;

float x_comp, y_comp, z_comp, p_polar, p_RA_DEC, th, phi, RA, DEC, RA_0, DEC_0, k, c_polar, s_polar, c_RA_DEC, s_RA_DEC, n1, n2;
ifp2 = fopen("amp_noise.dat", "r");
ifp3 = fopen("ph_noise.dat", "r");
ofp = fopen("power_pattern.dat", "w");
fprintf(ofp, "#theta phi RA DEC power(th,phi) power(RA,DEC)\n");
k = 2*pi/l;
for(th=(th_0 - r_th)*60; th<=(th_0 + r_th)*60; th=(th + a_th))//so th and py are in arc-minutes
{
for(phi=(phi_0 - r_phi)*60; phi<=(phi_0 + r_phi)*60; phi=(phi + a_phi))//1 degree = 60 arc-minute

```

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```

{
    RA = 4 * phi; // RA is expressed in Second since 1 arc-minute = 4 sec
    DEC = (90*60) - th; //DEC is expressed in arc-minute
    RA_0 = 4 * phi_0; //Reference RA direction
    DEC_0 = (90*60) - th_0; //Reference DEC direction
    c_polar = 0;
    s_polar = 0;
    c_RA_DEC = 0;
    s_RA_DEC = 0;
    for(i=0;i<m;i++)
    {
        fscanf(ifp2,"%f",&n1);
        fscanf(ifp3,"%f",&n2);

        x_comp=x[i]*(sin(min*th)*cos(min*phi)-sin(min*th_P*60)*cos(min*phi_P*6
0));
        y_comp=y[i]*(sin(min*th)*sin(min*phi)-sin(min*th_P*60)*sin(min*phi_P*6
0));
        z_comp=z[i]*(cos(min*th)-cos(min*th_P*60));

        c_polar += n1*(cos(k*( x_comp+ y_comp+ z_comp)) + n2); //the external n
for amp_noise and internal for ph_noise. Here both are taken from an uniform ra
ndom distribution. But for more sophisticated version we can add different distr
ibution, in principle they should be also different for phase and amplitude.
        s_polar += n1*(sin(k*( x_comp+ y_comp+ z_comp)) + n2); //sin in polar c
o-ordinate

        c_RA_DEC += n1*(cos(k*(x[i]*(sin(min*DEC)*cos(min*phi)-sin(min*DEC_0)*
cos(min*phi_0))+y[i]*(sin(min*DEC)*sin(min*phi)-sin(min*DEC_0)*sin(min*phi_0))+z
[i]*(cos(min*DEC)-cos(min*DEC_0)))) + n2); //cosine in RA DEC co-ordinate system

        s_RA_DEC += n1*(sin(k*(x[i]*(sin(min*DEC)*cos(min*phi)-sin(min*DEC_0)*
cos(min*phi_0))+y[i]*(sin(min*DEC)*sin(min*phi)-sin(min*DEC_0)*sin(min*phi_0))+z
[i]*(cos(min*DEC)-cos(min*DEC_0)))) + n2);
    }

    p_polar = (c_polar * c_polar)+(s_polar * s_polar); //resultant power patt
ern in th-phi
    p_RA_DEC = (c_RA_DEC * c_RA_DEC) +(s_RA_DEC * s_RA_DEC); //resultant powe
r pattern in RA-DEC

    fprintf (ofp,"%f %f %f %f %f %f\n",th,phi,RA,DEC,p_polar,p_RA_DEC);
}
fprintf (ofp,"n");

}
fclose(ifp2);
fclose(ifp3);
fclose(ofp);

printf("enter the value of thita in degree u want to fix, fth=");
scanf("%f",&f_th);

FILE *ifp_6;
FILE *ofp_7;
char str[100];
ifp_6 = fopen("power_pattern.dat", "r");
ofp_7 = fopen("power-pattern.fixtheta.dat", "w");
fgets(str,100,ifp_6);

```

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<pre> fprintf (ofp_7, "#fixed-theta phi RA DEC power(th,phi) power(RA,DEC)\n"); for(th=(th_0-r_th)*60;th<=(th_0+r_th)*60;th=(th+a_th))//so th and py are in arc -minutes { for(phi=(phi_0-r_phi)*60;phi<=(phi_0+r_phi)*60;phi=(phi+a_phi))//1 degree = 60 arc-minute { fscanf(ifp_6, "%f %f %f %f %f %f", &th, &phi, &RA, &DEC, &p_polar, &p_RA_DEC); if(th == (f_th*60)) { fprintf(ofp_7, "%f %f %f %f %f %f\n", (f_th*60), phi, RA, DEC, p_polar, p_RA_DE C); printf("%f %f %f %f %f %f\n", (f_th*60), phi, RA, DEC, p_polar, p_RA_DEC); } } } fclose(ofp_7); printf("enter the value of phi in degree u want to fix, fphi ="); scanf("%f", &f_phi); FILE *ofp_8; ofp_8 = fopen("power_pattern.fixphi.dat", "w"); rewind(ifp_6); fgets(str, 100, ifp_6); fprintf (ofp_7, "#theta fixed-phi RA DEC power(th,phi) power(RA,DEC)\n"); for(th=(th_0-r_th)*60;th<=(th_0+r_th)*60;th=(th+a_th))//so th and py are in ar c-minutes { for(phi=(phi_0-r_phi)*60;phi<=(phi_0+r_phi)*60;phi=(phi+a_phi))//1 degree = 60 arc-minute { fscanf(ifp_6, "%f %f %f %f %f %f", &th, &phi, &RA, &DEC, &p_polar, &p_RA_DEC); if(phi == (f_phi*60)) { fprintf(ofp_8, "%f %f %f %f %f %f\n", th, (f_phi*60), RA, DEC, p_polar, p_RA_DE C); printf("%f %f %f %f %f %f\n", th, (f_phi*60), RA, DEC, p_polar, p_RA_DEC); } } } fclose(ifp_6); fclose(ofp_8); return(0); } /* Generate a set of 30 random number of uniform distribution of some specific m ean and rms which will be used as noise in the array pattern*/ int noise(float rms1, float rms2, float mean1, float mean2) { FILE *ofp1 = NULL; FILE *ofp2 = NULL; ofp1 = fopen("amp_noise.dat", "w"); ofp2 = fopen("ph_noise.dat", "w"); </pre>		

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```

int i;
float a,b;

for(i=0;i<30;i++)
{
a = rms1 * rand() + mean1;
b = rms2 * rand() + mean2;

fprintf(ofp1,"%f\n",a);
fprintf(ofp2,"%f\n",b);
printf("%d %f %f\n",i,a,b);
}

fclose(ofp1);
fclose(ofp2);
}

/* make a program to convert the RA & DEC co-ordinate system in to theta (elivati
ion) and phy (azimath) co-ordinate system */

int RA_DEC_to_th_phi(float RA_hr,float RA_min,float RA_sec,float DEC_deg,float D
EC_min,float DEC_sec)
{
float th_0, phi_0;
phi_0 = (RA_sec * 15)/(60*60) + (RA_min * 15)/(60) + (RA_hr * 15);//phi calcula
ted in degree : 1 hr = 15 deg, 1 min = 15 arc-min, 1 sec = 15 arc-sec
th_0 = 90 - ((DEC_deg) + (DEC_min/60) + (DEC_sec/(60*60)));// th in Degree

printf("Theta_0=%f\n",th_0);
printf("Phi_0=%f\n",phi_0);

return(0);
}

```

PROGRAM NAME:: arraypattern_version2.c

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```
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/*COMMENTS : 1. MULTIPLE BEAM GENERATION IN THETA DIRECTION CONCEPT BROUGHT
IN
2. THREE FIXED VALUES OF THETA_P SET IN THE PROGRAM, AND IF-ELS
E LOOP WRITES DATA INTO 3 INDIVIDUAL FILES. */

#include<stdio.h>
#include <math.h>

#define pi 2*acos(0)
#define min pi/(2*9*600) //1 arc-minute = 0.000291005 rad

int RA_DEC_to_th_phi(float RA_hr,float RA_min,float RA_sec,float DEC_deg,float DEC_min,float DEC_sec);
int noise(float rms1,float rms2,float mean1,float mean2);

int main()
{
    int n,m;
    float x[30],y[30],z[30],l,th_P,phi_P,th_0,phi_0,r_th,a_th,r_phi,a_phi,f_th,f_phi,
    rms1,mean1,rms2,mean2,RA_hr,RA_min,RA_sec,DEC_deg,DEC_min,DEC_sec;
    system("clear");
    printf("enter the rms-s and means you want for uniformly distributed 30 random numbers will be used as noise t
o the beam pattern corresponding to amplitude and phase respectively ");
    scanf("%f%f%f%f",&rms1,&mean1,&rms2,&mean2);
    noise(rms1,rms2,mean1,mean2);

    // Read the antenna co-ordinates from the input file
    FILE *ifp = NULL;
    int i;
    float x1[30], y1[30], z1[30];

    ifp = fopen("input.dat","r"); //input.dat is the input file for antenna co-ordinat
es in x,y,z plane.
    for(i=0;i<30;i++)
        fscanf(ifp,"%f%f%f\n",&x[i],&y[i],&z[i]);
    fclose(ifp);

    // Select the reference antenna co-ordinate from the list of the antennas
    printf("choose the antenna number you want as reference n = ");
    scanf("%d",&n);
    printf("the antenna co-ordinates you select as reference are\n");
    printf("%5.2f %5.2f %5.2f\n",x[n-1],y[n-1],z[n-1]);

    // Subtract the ref antenna co-ordinate from the all antennas
    printf("\nthe corrected co-ordinates in meters are\n");
    for(i=0;i<30;i++)
    { x1[i] = x[i] - x[n-1];
      y1[i] = y[i] - y[n-1];
      z1[i] = z[i] - z[n-1];
      printf("%5.2f %5.2f %5.2f\n",x1[i],y1[i],z1[i]);
    }

    // Get the No of antenna as well as the antenna nos
    printf("choose the number (must be <= 30) of antennas m = ");
    scanf("%i",&m);
    int j[30];
    printf("\n choose the sequence of antennas. You must input the reference antenna.\n");

```

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```

for(i=0;i<m;i++)
{
scanf("%i",&j[i]);
printf("%i\n",j[i]);
}

// Get the antenna co-ordinates of the selected antennas
int cnt;
int count = 0;
printf("the corresponding antenna co-ordinates are\n");
for(i=0;i<30;i++)
{
for(cnt=0;cnt<m;cnt++)
{
if(i == j[cnt])
{
printf("%5.2f %5.2f %5.2f\n",x1[i],y1[i],z1[i]);
x[count] = x1[i]; y[count] = y1[i]; z[count] = z1[i];
count++;
}
}
}
printf("enter the reference direction (in sky) in RA(RA_hr,RA_min,RA_sec) and DEC(DEC_deg,DEC_min,DEC_sec) co-ordinate system");
scanf("%f %f %f %f %f %f", &RA_hr, &RA_min, &RA_sec, &DEC_deg, &DEC_min, &DEC_sec);
RA_DEC_to_th_phi(RA_hr,RA_min,RA_sec,DEC_deg,DEC_min,DEC_sec);

printf("\nenter the wavelength l in meter = ");
scanf("%f",&l);
printf("enter Phi_P which specify the direction for phasing the array ");
scanf("%f",&phi_P);
printf("enter Theta_0 and Phi_0 which specify the direction for calculation of pattern ");
scanf("%f %f",&th_0,&phi_0);
printf("\nenter the range you want in degree and the accuracy you want in arc-minute in the direction of theta or polar angle\n");
scanf("%f %f",&r_th,&a_th);
printf("\nenter the range you want in degree and the accuracy you want in arc-minute in the direction of phi or azimath\n");
scanf("%f %f",&r_phi,&a_phi);

FILE *ifp2;
FILE *ifp3;
FILE *ofp;
FILE *ofp1;
FILE *ofp11;
FILE *ofp111;

float x_comp,y_comp,z_comp,p_polar,p_RA_DEC,th,phi,RA,DEC,RA_0,DEC_0,k,c_polar,s_polar,c_RA_DEC,s_RA_DEC,n1,n2;
ifp2 = fopen("amp_noise.dat","r");
ifp3 = fopen("ph_noise.dat","r");
ofp = fopen("power_pattern.dat","w");
ofp1 =fopen("power_theta_pl.dat","w");
ofp11=fopen("power_theta_pll.dat","w");
ofp111=fopen("power_theta_pll1.dat","w");

fprintf (ofp,"#theta theta_P phi RA DEC power(th,phi) power(RA,DEC)\n");
fprintf (ofp1,"#phi phi_p theta_P theta power(th,phi)\n");
fprintf (ofp11,"#phi phi_p theta_P theta power(th,phi)\n");
fprintf (ofp111,"#phi phi_p theta_P theta power(th,phi)\n");

```

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```

k = 2*pi/l;
for(th_P=0;th_P<=1;th_P+=.5)
{
for(phi=(phi_0 - r_phi)*60;phi<=(phi_0 + r_phi)*60;phi=(phi + a_phi))//1 degree
= 60 arc-minute
{
for(th=(th_0 - r_th)*60;th<=(th_0 + r_th)*60;th=(th + a_th))//so th and py are
in arc-minutes
{
RA = 4 * phi; // RA is expressed in Second since 1 arc-minute = 4 sec
DEC = (90*60) - th; //DEC is expressed in arc-minute
RA_0 = 4 * phi_0; //Reference RA direction
DEC_0 = (90*60) - th_0; //Reference DEC direction
c_polar = 0;
s_polar = 0;
c_RA_DEC = 0;
s_RA_DEC = 0;
for(i=0;i<m;i++)
{
fscanf(ifp2,"%f",&n1);
fscanf(ifp3,"%f",&n2);
x_comp=x[i]*(sin(min*th)*cos(min*phi)-sin(min*th_P*60)*cos(min*phi_P*6
0));
y_comp=y[i]*(sin(min*th)*sin(min*phi)-sin(min*th_P*60)*sin(min*phi_P*6
0));
z_comp=z[i]*(cos(min*th)-cos(min*th_P*60));

c_polar += n1*(cos(k*(x_comp+y_comp+z_comp)) + n2); //the external n fo
r amp_noise and internal for ph_noise. Here both are taken from an uniform rando
m distribution. But for more sophisticated version we can add different distribu
tion, in principle they should be also different for phase and amplitude.

s_polar += n1*(sin(k*(x_comp+y_comp+z_comp)) + n2); //sin in polar co-ord
inate

c_RA_DEC += n1*(cos(k*(x[i]*(sin(min*DEC)*cos(min*phi)-sin(min*DEC_0)*
cos(min*phi_0))+y[i]*(sin(min*DEC)*sin(min*phi)-sin(min*DEC_0)*sin(min*phi_0))+z
[i]*(cos(min*DEC)-cos(min*DEC_0)))) + n2); //cosine in RA DEC co-ordinate system

s_RA_DEC += n1*(sin(k*(x[i]*(sin(min*DEC)*cos(min*phi)-sin(min*DEC_0)*
cos(min*phi_0))+y[i]*(sin(min*DEC)*sin(min*phi)-sin(min*DEC_0)*sin(min*phi_0))+z
[i]*(cos(min*DEC)-cos(min*DEC_0)))) + n2);
}

p_polar = (c_polar * c_polar)+(s_polar * s_polar); //resultant power patt
ern in th-phi
p_RA_DEC = (c_RA_DEC * c_RA_DEC) +(s_RA_DEC * s_RA_DEC); //resultant powe
r pattern in RA-DEC

if(th_P==0)
{
fprintf (ofp1,"%f %f %f %f %f\n",phi,phi_P,th_P,th, p_polar
);
}
}
}

```

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-----------------	--------------------------------	----------

```

printf ("IFIOOPl %f %f %f %f %f\n",phi,phi_P,th_P,th,p_polar);
// WRITING DATA INTO 3 INDIVIDUAL FILES ,FOR EACH PHASING ANGLE
}

else if (th_P--.5)
{
fprintf (ofp11,"%f %f %f %f%f\n",phi,phi_P,th_P,th, p_polar );
printf ("IFIOOPl1 %f %f %f %f %f\n",phi,phi_P,th_P,th,p_polar);
}
else if(th_P--1)
{
fprintf (ofp111,"%f %f %f %f%f\n",phi,phi_P,th_P,th, p_polar );
printf ("IFIOOPl11 %f %f %f %f %f\n",phi,phi_P,th_P,th,p_polar);
}

fprintf (ofp,"%f %f %f %f %f %f %f\n",th,th_P,phi,RA,DEC,p_polar,p_RA
_DEC);

} //theta
fprintf (ofp,"\n");

} //phi
} //theta_p

fclose(ifp2);
fclose(ifp3);
fclose(ofp);
fclose(ofp1);
fclose(ofp11);
fclose(ofp111);

printf("enter the value of thita in degree u want to fix, fth =");
scanf("%f",&f_th);

FILE *ifp_6;
FILE *ofp_7;
char str[100];
ifp_6 = fopen("power_pattern.dat","r");
ofp_7 = fopen("power-pattern.fixtheta.dat","w");
fgets(str,100,ifp_6);
fprintf (ofp_7,"#fixed-theta phi RA DEC power(th,phi) power(RA,DEC)\n");

for(th=(th_0-r_th)*60;th<=(th_0+r_th)*60;th=(th+a_th))//so th and py are in arc
-minutes
{
for(phi=(phi_0-r_phi)*60;phi<=(phi_0+r_phi)*60;phi=(phi+a_phi))//1 degree = 60
arc-minute
{
fscanf(ifp_6,"%f %f %f %f %f %f",&th,&phi,&RA,&DEC,&p_polar,&p_RA_DEC);
if(th == (f_th*60))
{
fprintf(ofp_7,"%f %f %f %f %f %f\n",(f_th*60),phi,RA,DEC,p_polar,p_RA_DE
C);
printf("%f %f %f %f %f %f\n",(f_th*60),phi,RA,DEC,p_polar,p_RA_DEC);
}
}
}

```



```

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}
}
fclose(ofp_7);
printf("enter the value of phi in degree u want to fix, fphi =");
scanf("%f", &f_phi);

FILE *ofp_8;

ofp_8 = fopen("power_pattern.fixphi.dat", "w");
rewind(ifp_6);
fgets(str, 100, ifp_6);
fprintf(ofp_8, "#theta fixed-phi RA DEC power(th,phi) power(RA,DEC)\n");
for(th=(th_0-r_th)*60; th<=(th_0+r_th)*60; th=(th+a_th))//so th and py are in ar
c-minutes
{
for(phi=(phi_0-r_phi)*60; phi<=(phi_0+r_phi)*60; phi=(phi+a_phi))//1 degree = 60
arc-minute
{
fscanf(ifp_6, "%f %f %f %f %f", &th, &phi, &RA, &DEC, &p_polar, &p_RA_DEC);
if(phi == (f_phi*60))
{
fprintf(ofp_8, "%f %f %f %f %f %f\n", th, (f_phi*60), RA, DEC, p_polar, p_RA_DE
C);
printf("%f %f %f %f %f\n", th, (f_phi*60), RA, DEC, p_polar, p_RA_DEC);
}
}
}
fclose(ifp_6);
fclose(ofp_8);
return(0);
}

/* Generate a set of 30 random number of uniform distribution of some specific m
ean and rms which will be used as noise in the array pattern*/

int noise(float rms1, float rms2, float mean1, float mean2)
{
FILE *ofp1 = NULL;
FILE *ofp2 = NULL;
ofp1 = fopen("amp_noise.dat", "w");
ofp2 = fopen("ph_noise.dat", "w");
int i;
float a, b;

for(i=0; i<30; i++)
{
a = rms1 * rand() + mean1;
b = rms2 * rand() + mean2;

fprintf(ofp1, "%f\n", a);
fprintf(ofp2, "%f\n", b);
printf("%d %f %f\n", i, a, b);
}

fclose(ofp1);
fclose(ofp2);
}

```

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```
/* make a program to convert the RA & DEC co-ordinate system in to theta (elivat  
ion) and phy (azimath) co-ordinate system */  
  
int RA_DEC_to_th_phi(float RA_hr,float RA_min,float RA_sec,float DEC_deg,float D  
EC_min,float DEC_sec)  
{  
    float th_0, phi_0;  
    phi_0 = (RA_sec * 15)/(60*60) + (RA_min * 15)/(60) + (RA_hr * 15);//phi calcula  
ted in degree : 1 hr = 15 deg, 1 min = 15 arc-min, 1 sec = 15 arc-sec  
    th_0 = 90 - ((DEC_deg) + (DEC_min/60) + (DEC_sec/(60*60)));// th in Degree  
  
    printf("Theta_0=%f\n",th_0);  
    printf("Phi_0=%f\n",phi_0);  
  
    return(0);  
}
```

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PROGRAM NAME:: arraypattern_version3.c

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```
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/* COMMENTS: FINAL CODE OF THE PROGRAM
              MULTIPLE BEAM GENERATION OVER 2D SKY PLANE POSSIBLE*/

#include<stdio.h>
#include <math.h>

#define pi 2*acos(0)
#define min pi/(2*9*600) //1 arc-minute = 0.000291005 rad

int RA_DEC_to_th_phi(float RA_hr,float RA_min,float RA_sec,float DEC_deg,float D
EC_min,float DEC_sec);
int noise(float rms1,float rms2,float mean1,float mean2);

int main()
{
    int n,m,no_th,no_phi,s,inp1,inp2,inp3,inp4;
    float x[30],y[30],z[30],l,th_P0,phi_P0,th_P[20],phi_P[20],d_th_P,d_phi_P,th_0,p
hi_0,r_th,a_th,r_phi,a_phi,f_th,f_phi,rms1,mean1,rms2,mean2,RA_hr,RA_min,RA_sec,
DEC_deg,DEC_min,DEC_sec;
    char fname[100];
    system("clear");
    printf("enter the rms-s and means you want for uniformly distributed 30 random numbers will be used as noise t
o the beam pattern corresponding to amplitude and phase respectively ");
    scanf("%f%f%f%f",&rms1,&mean1,&rms2,&mean2);
    noise(rms1,rms2,mean1,mean2);

    // Read the antenna co-ordinates from the input file
    FILE *ifp = NULL;
    int i;
    float x1[30], y1[30], z1[30];

    ifp = fopen("input.dat","r"); //input.dat is the input file for antenna co-ordinat
es in x,y,z plane.
    for(i=0;i<30;i++)
        fscanf(ifp,"%f%f%f\n",&x[i],&y[i],&z[i]);
    fclose(ifp);

    // Select the reference antenna co-ordinate from the list of the antennas
    printf("choose the antenna number you want as reference n = ");
    scanf("%d",&n);
    printf("the antenna co-ordinates you select as reference are\n");
    printf("%5.2f %5.2f %5.2f\n",x[n-1],y[n-1],z[n-1]);

    // Subtract the ref antenna co-ordinate from the all antennas
    printf("\nthe corrected co-ordinates in meters are:\n");
    for(i=0;i<30;i++)
    { x1[i] = x[i] - x[n-1];
      y1[i] = y[i] - y[n-1];
      z1[i] = z[i] - z[n-1];
      printf("%5.2f %5.2f %5.2f\n",x1[i],y1[i],z1[i]);
    }

    // Get the No of antenna as well as the antenna nos
    printf("choose the number (must be <= 30) of antennas m = ");
    scanf("%i",&m);
    int j[30];
    printf("\n choose the sequence of antennas. You must input the reference antenna. \n");

```

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```

for(i=0;i<m;i++)
{
scanf("%i",&j[i]);
printf("%i\n",j[i]);
}

// Get the antenna co-ordinates of the selected antennas
int cnt;
int count = 0;
printf("the corresponding antenna co-ordinates are\n");
for(i=0;i<30;i++)
{
for(cnt=0;cnt<m;cnt++)
{
if(i == j[cnt])
{
printf("%5.2f %5.2f %5.2f\n",x1[i],y1[i],z1[i]);
x[count] = x1[i]; y[count] = y1[i]; z[count] = z1[i];
count++;
}
}
}
printf("enter the reference direction (in sky) in RA(RA_hr,RA_min,RA_sec) and DEC(DEC_deg,DEC_min,DEC_sec) co-ordinate system ");
scanf("%f %f %f %f %f %f",&RA_hr,&RA_min,&RA_sec,&DEC_deg,&DEC_min,&DEC_sec);
RA_DEC_to_th_phi(RA_hr,RA_min,RA_sec,DEC_deg,DEC_min,DEC_sec);

printf("\nenter the wavelength l in meter = ");
scanf("%f",&l);
printf("\nenter the no. of multiple beams to be generated in a given range of theta and given range of phi = ");
scanf("%d %d",&no_th,&no_phi);

printf("\nenter theta_p and Phi_P which specify the direction for phasing the array ");
scanf("%f %f",&th_P0,&phi_P0);
printf("\nenter separation in th_p (in arc min) and phi_p (in arc min) at which the beams are to be separated in the direction ");
scanf("%f %f",&d_th_P,&d_phi_P);
printf("\nenter Theta_0 and Phi_0(in deg) which specify the direction for calculation of pattern ");
scanf("%f %f",&th_0,&phi_0);
printf("\nenter the range you want (in degree) and the accuracy you want in (arc-minute) in the direction of theta or polar angle \n");
scanf("%f %f",&r_th,&a_th);
printf("\nenter the range you want (in degree) and the accuracy you want (in arc-minute) in the direction of phi or azimuth \n");
scanf("%f %f",&r_phi,&a_phi);

FILE *ifp2;
FILE *ifp3;
FILE *ofp;
FILE *ofpnl;
FILE *ifpn;

float x_comp,y_comp,z_comp,p_polar,p_RA_DEC,th,phi,RA,DEC,RA_0,DEC_0,k,c_polar,s_polar,c_RA_DEC,s_RA_DEC,n1,n2;
ifp2 = fopen("amp_noise.dat","r");
ifp3 = fopen("ph_noise.dat","r");
ofp = fopen("power_pattern.dat","w");
ifpn = fopen("input_thp_phi.dat","w");
fprintf(ifpn,"#theta_P phi_P \n");

```

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<pre> inp1=0; inp2=0; for(th_P[inp1]=th_P0*60-(no_th-1)*d_th_P;th_P[inp1]<th_P0*60+no_th*d_th_P;th_P[inp1]=th_P[inp1-1]+d_th_P) //for generation of input dat //file having thp phip values { for(phi_P[inp2]=phi_P0*60-(no_phi-1)*d_phi_P;phi_P[inp2]<phi_P0*60+no_phi*d_phi_P;phi_P[inp2]=phi_P[inp2-1]+d_phi_P) { fprintf(ifpn,"%f %f\n",th_P[inp1],phi_P[inp2]); inp2++; } inp1++; } fclose(ifpn); fprintf(stderr,"checking...after th phi witting"); fprintf (ofp,"#theta theta_P phi_P phi RA DEC power(th,phi) power(RA,DEC)\n"); k = 2*pi/l; s=0; char spr[300]; ifpn=fopen("input_thp_phipdat","r"); fgets(spr,300,ifpn); for(inp3 = 0;inp3 <2*no_th-1; inp3++) //for obtainin values of th_P,phi_P frm input_thp_phip file { for(inp4 = 0;inp4 <2*no_phi-1; inp4++) { fscanf(ifpn,"%f %f\n",&th_P[inp3],&phi_P[inp4]); for(phi=(phi_0 - r_phi)*60;phi<=(phi_0 + r_phi)*60;phi=(phi + a_phi))//1 degree = 60 arc-minute { for(th=(th_0 - r_th)*60;th<=(th_0 + r_th)*60;th=(th + a_th))//so th and py are in arc-minutes { RA = 4 * phi; // RA is expressed in Second since 1 arc-minute = 4 sec DEC = (90*60) - th; //DEC is expressed in arc-minute RA_0 = 4 * phi_0;//Reference RA direction DEC_0 = (90*60) - th_0;//Reference DEC direction c_polar = 0; s_polar = 0; c_RA_DEC = 0; s_RA_DEC = 0; for(i=0;i<m;i++) { fscanf(ifp2,"%f",&n1); fscanf(ifp3,"%f",&n2); x_comp=x[i]*(sin(min*th)*cos(min*phi)-sin(min*th_P[inp3])*cos(min*phi_P[inp4])); y_comp=y[i]*(sin(min*th)*sin(min*phi)-sin(min*th_P[inp3])*sin(min*phi_P[inp4])); </pre>		

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```

P[inp4]));
    z_comp=z[i]*(cos(min*th)-cos(min*th_P[inp3]));

    c_polar += nl*(cos(k*(x_comp+y_comp+z_comp)) + n2);//the external n for
amp_noise and internal for ph_noise. Here both are taken from an uniform random
distribution. But for more sophisticated version we can add different distribution,
in principle they should be also different for phase and amplitude.

    s_polar += nl*(sin(k*(x_comp+y_comp+z_comp)) + n2);//sin in polar coordinate

    c_RA_DEC += nl*(cos(k*(x[i]*(sin(min*DEC)*cos(min*phi)-sin(min*DEC_0)*
cos(min*phi_0))+y[i]*(sin(min*DEC)*sin(min*phi)-sin(min*DEC_0)*sin(min*phi_0))+z
[i]*(cos(min*DEC)-cos(min*DEC_0)))) + n2);//cosine in RA DEC co-ordinate system

    s_RA_DEC += nl*(sin(k*(x[i]*(sin(min*DEC)*cos(min*phi)-sin(min*DEC_0)*
cos(min*phi_0))+y[i]*(sin(min*DEC)*sin(min*phi)-sin(min*DEC_0)*sin(min*phi_0))+z
[i]*(cos(min*DEC)-cos(min*DEC_0)))) + n2);
}

    p_polar = (c_polar * c_polar)+(s_polar * s_polar);//resultant power pattern
in th-phi
    p_RA_DEC = (c_RA_DEC * c_RA_DEC) +(s_RA_DEC * s_RA_DEC);//resultant power
pattern in RA-DEC

    sprintf(fname,"power_theta_p%d.dat", s);
    ofpnl =fopen(fname,"a");
    fprintf (ofpnl,"%f %f %f %f %f\n",th_P[inp3],phi_P[inp4],th,
phi ,p_polar );
    printf ("IFLOOP%d %f %f %f %f %f %f\n",s,th_P0,th_P[inp3
],phi_P[inp4],th,phi,p_polar);
    //fprintf(ofpnl, "\n");
    fclose(ofpnl);

    fprintf (ofp,"%f %f %f %f %f %f %f %f\n",th,th_P[inp3],phi_P[
inp4],phi,RA,DEC,p_polar,p_RA_DEC);

} //theta
fprintf (ofp, "\n");
ofpnl =fopen(fname,"a");
fprintf(ofpnl, "\n");
fclose(ofpnl); // newline for new value of phi
} //phi
printf("VALUE OF S %d\n",s);
s++;

} //generation of phi_P

} //generation of theta_P

fclose(ifp2);
fclose(ifp3);
fclose(ofp);
fclose(ifpn);

printf ("enter the value of thita in degree u want to fix, fth =");
scanf ("%f",&f_th);

```

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```

FILE *ifp_6;
FILE *ofp_7;
char str[100];
ifp_6 = fopen("power_pattern.dat", "r");
ofp_7 = fopen("power-pattern.fixtheta.dat", "w");
fgets(str, 100, ifp_6);
fprintf ( ofp_7, "#fixed-theta phi RA DEC power(th,phi) power(RA,DEC)\n");

for(th=(th_0-r_th)*60;th<=(th_0+r_th)*60;th=(th+a_th))//so th and py are in arc
-minutes
{
for(phi=(phi_0-r_phi)*60;phi<=(phi_0+r_phi)*60;phi=(phi+a_phi))//1 degree = 60
arc-minute
{
fscanf(ifp_6, "%f %f %f %f %f %f", &th, &phi, &RA, &DEC, &p_polar, &p_RA_DEC);
if(th == (f_th*60))
{
fprintf(ofp_7, "%f %f %f %f %f %f\n", (f_th*60), phi, RA, DEC, p_polar, p_RA_DE
C);
printf("%f %f %f %f %f %f\n", (f_th*60), phi, RA, DEC, p_polar, p_RA_DEC);
}
}
}
fclose(ofp_7);
printf("enter the value of phi in degree u want to fix, fphi =");
scanf("%f", &f_phi);

FILE *ofp_8;

ofp_8 = fopen("power_pattern.fixphi.dat", "w");
rewind(ifp_6);
fgets(str, 100, ifp_6);
fprintf ( ofp_8, "#theta fixed-phi RA DEC power(th,phi) power(RA,DEC)\n");
for(th=(th_0-r_th)*60;th<=(th_0+r_th)*60;th=(th+a_th))//so th and py are in ar
c-minutes
{
for(phi=(phi_0-r_phi)*60;phi<=(phi_0+r_phi)*60;phi=(phi+a_phi))//1 degree = 60
arc-minute
{
fscanf(ifp_6, "%f %f %f %f %f %f", &th, &phi, &RA, &DEC, &p_polar, &p_RA_DEC);
if(phi == (f_phi*60))
{
fprintf(ofp_8, "%f %f %f %f %f %f\n", th, (f_phi*60), RA, DEC, p_polar, p_RA_DE
C);
printf("%f %f %f %f %f %f\n", th, (f_phi*60), RA, DEC, p_polar, p_RA_DEC);
}
}
}
fclose(ifp_6);
fclose(ofp_8);
return(0);
}

/* Generate a set of 30 random number of uniform distribution of some specific m
ean and rms which will be used as noise in the array pattern*/

```

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```

int noise(float rms1,float rms2, float mean1,float mean2)
{
    FILE *ofp1 = NULL;
    FILE *ofp2 = NULL;
    ofp1 = fopen("amp_noise.dat","w");
    ofp2 = fopen("ph_noise.dat","w");
    int i;
    float a,b;

    for(i=0;i<30;i++)
    {
        a = rms1 * rand() + mean1;
        b = rms2 * rand() + mean2;

        fprintf(ofp1,"%f\n",a);
        fprintf(ofp2,"%f\n",b);
        printf("%d %f %f\n",i,a,b);
    }

    fclose(ofp1);
    fclose(ofp2);
}

/* make a program to convert the RA & DEC co-ordinate system in to theta (elivat
ion) and phy (azimath) co-ordinate system */

int RA_DEC_to_th_phi(float RA_hr,float RA_min,float RA_sec,float DEC_deg,float D
EC_min,float DEC_sec)
{
    float th_0, phi_0;
    phi_0 = (RA_sec * 15)/(60*60) + (RA_min * 15)/(60) + (RA_hr * 15);//phi calcula
ted in degree ; 1 hr = 15 deg, 1 min = 15 arc-min, 1 sec = 15 arc-sec
    th_0 = 90 - ((DEC_deg) + (DEC_min/60) + (DEC_sec/(60*60)));// th in Degree

    printf("Theta_0 = %f\n", th_0);
    printf("Phi_0 = %f\n", phi_0);

    return(0);
}

```


DATA FILES

GMRT COORDINATES:

X[I]	Y[I]	Z[I]
6.95	687.88	-20.04
13.24	326.43	-40.35
0.00	0.00	0.00
-51.17	-372.71	133.59
-51.08	-565.94	123.43
79.09	67.82	-246.59
71.23	-31.44	-220.58
130.77	280.67	-400.33
48.56	41.92	-151.65
191.32	-164.88	-587.49
102.42	-603.28	-321.56
209.28	174.85	-635.54
368.58	-639.53	-1117.92
207.30	-473.71	-628.63
-348.04	2814.55	953.67
-707.58	4576.00	1932.46
-1037.11	7780.69	2903.29
-1177.37	10200.00	3343.20
-1571.32	12073.46	4543.13
942.99	633.92	-2805.93
1452.92	-367.07	-4279.16
2184.54	333.03	-6404.96
3072.86	947.68	-8979.50
4592.71	-369.04	-13382.48
-201.32	-1591.90	591.32
-482.67	-3099.41	1419.39
-992.01	-5199.90	2899.11
-1734.55	-7039.03	5067.53
-2706.09	-8103.13	7817.14
-3102.11	-11245.60	8916.26

LINEAR X-Y ARRAY

X[I]	Y[I]	Z[I]
0000.0000	0000.0000	0000.0000
0100.0000	0100.0000	0000.0000
0200.0000	0200.0000	0000.0000
0300.0000	0300.0000	0000.0000
0400.0000	0400.0000	0000.0000
0500.0000	0500.0000	0000.0000
0600.0000	0600.0000	0000.0000
0700.0000	0700.0000	0000.0000
0800.0000	0800.0000	0000.0000
0900.0000	0900.0000	0000.0000
1000.0000	1000.0000	0000.0000
1100.0000	1100.0000	0000.0000
1200.0000	1200.0000	0000.0000
1300.0000	1300.0000	0000.0000
1400.0000	1400.0000	0000.0000
1500.0000	1500.0000	0000.0000
1600.0000	1600.0000	0000.0000
1700.0000	1700.0000	0000.0000
1800.0000	1800.0000	0000.0000
1900.0000	1900.0000	0000.0000
2000.0000	2000.0000	0000.0000
2100.0000	2100.0000	0000.0000
2200.0000	2200.0000	0000.0000
2300.0000	2300.0000	0000.0000
2400.0000	2400.0000	0000.0000
2500.0000	2500.0000	0000.0000
2600.0000	2600.0000	0000.0000
2700.0000	2700.0000	0000.0000
2800.0000	2800.0000	0000.0000
2900.0000	2900.0000	0000.0000

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