

Wideband Radio Frequency Interference in the 233 MHz band

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1 Background

Wideband radio frequency interference (RFI) has been observed in the 233 MHz band as shown in this online display of the visibility amplitudes while observing an off-source position (Fig 1). The top half above the diagonal shows the 130 MHz polarization whereas the bottom half shows the 175 MHz polarisation. The diagonal shows the self counts for both the polarizations. This is an instantaneous display for a given spectral channel. Correlated signal is picked up by the central square antennas and the nearest arm antennas (e.g W01, S01, E02) even when the antennas are pointing at an off-source position where no correlated signal is expected. The correlated signal simulates a source of strength $\sim 20 - 30$ Jy and is present in all the spectral channels unlike narrow-band RFI. Such wide band RFI reduces the dynamic range of the observations by increasing the rms noise in the image and hence reducing the sensitivity of the observations. This is not desirable for observations and it is necessary to determine its origin and eliminate it. We undertook a few experiments, which we describe here alongwith the results, to understand the origin of this RFI and ways to live with it till it is eliminated.

First some history. Such correlated wideband RFI was observed by the observatory about two years ago, but it was certainly absent about six months ago. Some of the remedial measures taken two years ago were:

- Sparking/arcing in power lines were checked for. Apparently a culprit was found near W01. It was not clear if the wideband RFI disappeared.
- Some of the power line insulators were replaced. It was not clear if the wideband RFI disappeared.
- The wideband RFI clearly disappeared after the rains then!

Some of the suspects this time could be:

- sparking/arcing in power lines
- the wideband (8 MHz) TV signals getting intermodulated into the observing band.
- unknown!

We conducted a few experiments in March 2003. The first experiment concentrated on determining the origin of this RFI. The second and third experiments concentrated on finding an appropriate observing method to co-exist with the RFI till it was eliminated. In this memo, we describe the experiments and the results we obtained. In summary, we did not obtain a conclusive result as to the origin of the RFI; however we found that by changing the observing band or introducing high attenuation in the default band, one can reduce the effect of this RFI on observations.

2 Experiment 1

Fig 1 shows how the wideband RFI will be manifested while observing an off-source position. The shorter baselines show finite correlation while the longer baselines where the RFI is uncorrelated show noise. Suspecting the mains power lines to be responsible for the RFI, we decided to power down the central square antennas + arm antennas and run the central square on generator supply. We sequentially turned off the power to the south, west and east arms and noted the amplitude change in the correlated signal. All this work was done towards an off-source position so as to avoid confusion with the source signal. The experiment was done by using default RF settings for the 233 MHz band i.e. the IF band 244-236 MHz. We also made sure that to examine channels which were free of narrow-band RFI. The various steps taken in the experiment with the results are listed below:

1. The power to the south arm was switched off and a drop in the correlated signal was noticed. But the signal did not disappear. On switching off power supply to the east and west arms, no change in the RFI was noticed.
2. Then we switched off the power supply to the analog, electronics and baseband laboratories. We suspect that a few of the spiky narrow-band RFI which we had noticed, but unfortunately had not kept detailed records, had subsided when the analog laboratory was switched off. However, the experiment has to be repeated before any conclusions can be drawn. There was no change in the wideband correlated signal.
3. The power supply to the laboratories was turned on. No change in the correlated signal was noticed.
4. The power supply to the arm antennas was switched on. No change in the correlated signal was noticed. It is not clear whether the drop in the correlated signal noticed in (1) was related to powering down the south arm or was coincidental.
5. The power supply to central square was switched off and the diesel generator was switched on. No change in the correlated signal was noticed.
6. Power was restored on the central square. No change in the correlated signal.

The only change we noted was in the strength of the correlated signal which was 30 counts when we started and dropped to 10-15 counts when we ended. We have no explanation for this. It could either be that the correlated signal was variable during the experiment or else some of the power lines which contribute to the correlated signal are in the south arm and hence the RFI correlation reduced when the south arm antennas were powered down.

3 Experiment 2

In this experiment, we varied the frequency of the first local oscillator and scanned through the 233 MHz RF band to check for a ‘clean’ band which is free from the wideband RFI. One of the possible origins of this wideband RFI is cited as the TV signal which can get intermodulated into our observing band. The Pune TV signal is between 175 and 183 MHz. Junnar TV receives this signal and retransmits between 195 and 203 MHz. If these signals which lie outside the 233 MHz RF band beat with a narrow-band RFI, there is a possibility of them getting translated to within the 233 MHz RF band and we would observe it as a wideband RFI. However, if this is the case then the wideband RFI should have a frequency dependence depending on the translating narrow-band RFI which is beating with the TV signal. This experiment was aimed at investigating this issue. We observed an off-source position for all the frequency settings and checked the magnitude of the wideband correlated signal. Please note that the default settings for 233 MHz observations (244-236 MHz) are: LOI=310 MHz, IFBW=6 MHz, LOIV=66 MHz, BBW=8 MHz. Fig 1 shows the counts registered at an off-source position. Notice the higher counts on the central square antennas which is the signature of the wideband RFI which we are trying to remedy. The LOI frequency was changed whereas the LOIV frequency was held fixed to 66 MHz for all the following cases. The strength of the correlated RFI signal was compared to the strength obtained for the default RF settings.

1. Changed LOI to 300 MHz so that the observing band was 234-226 MHz. The correlated RFI signal was stronger.
2. Changed LOI=295 MHz so observing band was 229-221 MHz. The correlated RFI signal was stronger (See Fig 2). However note that the on-source sensitivity is down by 5 dB.
3. Changed LOI=320 MHz so observing band was 254-246 MHz. The correlated RFI signal was weaker (See Fig 3).
4. Changed LOI=325 MHz, so observing band was 259-251 MHz. The correlated RFI signal was weakest in this case (See Fig 4). The on-source sensitivity was down by 5 dB.

Summarizing, it appears that the wideband RFI is weakest in the 251 to 259 MHz band and strongest in the 221 – 229 MHz band. In other words, the wideband

RFI gets weaker for frequencies from 220 \rightarrow 250 MHz. (Subhashis Roy conducted a similar experiment the next day and confirmed the above results.) Thus observing in the band 251 – 259 MHz seems better for reducing the effect of the wideband RFI on the data. However, our experiment does not give any information on the presence of narrow band RFI in this band. If the narrow band RFI in this band turns out to be severe then it might be a good idea to adhere to the default settings and resort to other measures. We conducted the next experiment to check out if the default observing band at 233 MHz can be used with reduced sensitivity and reduced contamination by the wideband RFI.

4 Experiment 3

Since the system temperature at 233 MHz is high (stronger background non-thermal radiation field) and RFI can be strong which can saturate the system; the normal observing procedure involves introducing a 14 dB solar attenuator in the front-end. At GMRT, a facility is provided for inserting a large attenuator in the front-end. This attenuation can either be 14 dB or 30 dB and are known as solar attenuators. In this experiment, a solar attenuator of 30dB was introduced in the path for the default 233 MHz settings (LOI=310MHz, LOIV=66MHz). 3C147 was observed. The correlation with the default 14dB solar attenuator of 0.07-0.08 (see Fig 5) dropped to 0.02 (see Fig 6) that is a drop of factor ~ 4 ie ~ 6 dB. This experiment was conducted to check if the 30 dB solar attenuator would be sufficient to kill the wideband RFI without losing sensitivity.

The off-source position with a 30 dB attenuation in the front-end (see Fig 7) shows weak signatures of the wideband correlated signal; the cross correlation of the RFI is lower by a factor 4 like the on-source correlation. Thus the results of this experiment and the previous experiment are similar. This experiment demonstrates that the default RF band can be used for observations with a 30 dB attenuator in the front end and this reduces the effect of the wideband RFI. As in the previous experiment, this does not cure the wideband RFI problem.

5 Summary

The origin of the wideband RFI in the 233 MHz band is not clear. Since the weather is fairly dry, it is likely that the cause of the RFI is in some sparking/arcing of power lines. However our control experiment to track this down was not conclusive. We, then, turned our attention to finding ways of observing in the presence of this wideband RFI and arriving at a compromise between reduced sensitivity and reduced wideband RFI. We find that the wideband RFI reduces in strength when scanning in RF from 221 to 259 MHz. The wideband RFI is strongest near 221 MHz and weakest near 259 MHz. There is a drop in sensitivity of ~ 5 dB near 221 and 259 MHz. In another experiment, we find that the wideband RFI can be reduced by introducing a solar attenuator of 30 dB in the front-end. This results in a drop

of 6 dB in the sensitivity. However, in both these cases, we find that the wideband RFI also becomes weaker.

Further investigations are required for pinning down the source of this RFI. Some further experiments that can be conducted are:

- Check for the wideband RFI when power fails and the TV signals are absent.
- Repeat the experiment of running the central square on the diesel generator and the effect on the wideband correlated signal.
- Check if rains cure the problem!

(**P.S.:** Although this memo was prepared soon after the above experiments were conducted, its submission has got delayed beyond the onset of monsoon, giving us an opportunity to examine the effect of rains on the wideband RFI. Unfortunately this time, we find that the wideband RFI persists and the rains have failed to provide the solution. More experiments are being planned and conducted to trace down the source of this RFI.)

Note: The figures are in another file (**figs.ps**). The captions for the figures are given below as also the page number in the figures file.

Figure Captions:

Fig 1: This figure shows the counts registered at an off-source position for default frequency settings in the 240 MHz band. Notice that while most of the arm antennas show a few counts which is noise; the central square antennas and the nearby arm antennas show few tens of counts which is the correlated signal, the source of which we are trying to locate. (page 4 or 5)

Fig 2: Expt 2(b): Off-source counts for LOI=295 MHz; (page 8).

Fig 3: Expt 2(c): Off-source counts for LOI=320 MHz (last page in this file).

Fig 4: Expt 2(d): Off-source counts for LOI=325 MHz; (page 7).

Fig 5: Expt 3: On-source (3C147) counts with standard 14dB solar attenuator; (page 1).

Fig 6: Expt 3: On-source (3C147) counts with 30dB solar attenuator; (page 3).

Fig 7: Expt 3: Off-source counts with 30dB solar attenuator; (page 2).

(254 - 246) MHz

off source

Proj: T130B175		TST Observer: Chan=40		NGK Title: INTERMODULATION TEST		mithuna:AMPMON		Med Har 12 22:47:45 2003		Time=22:47:42.05																								
C00 C01 C02 C03 C04 C05 C06 C08 C09 C10 C11 C12 C13 C14 C15		1000MHz Normalised ccf		Source=3C147		RF=254 L01=320 BLO=66Hz		15 16 17 18 19 20 21 22 23		S01 S02 S03 S04 S06																								
C00	12	25	8	7	27	15	24	4	19	8	10	3	10	27	4	1	3	1	6	5	3	11	3	7	3	2	4	6	1					
C01	7	12	2	15	9	6	14	8	7	6	3	5	8	8	6	1	2	4	4	3	5	4	1	6	3	11	6	3	3					
C02	11	10	1247	9	18	13	10	5	9	11	5	2	5	2	2	3	4	5	5	6	4	4	5	11	3	9	6	3	3					
C03	13	3	16	992	3	5	11	4	13	3	10	9	9	4	5	6	3	2	1	10	6	4	5	2	5	3	7	6	3	2				
C04	9	14	12	28	957	13	4	11	9	13	3	6	11	6	4	4	5	5	2	4	13	4	9	5	8	6	3	11	8	2				
C05	9	2	10	1	11	1496	9	20	3	6	7	4	11	9	6	3	6	4	2	4	3	3	5	3	2	1	5	5	1	4				
C06	18	7	4	15	16	5	185	2	4	4	3	11	4	10	6	2	1	1	4	1	3	1	1	4	6	7	3	2	2					
C08	18	15	11	12	15	7	1197	9	9	5	3	4	8	9	2	4	3	4	7	8	5	8	3	5	8	4	1	3	2					
C09	6	20	7	9	13	8	10	9	1032	7	2	4	12	6	3	4	2	2	2	7	3	3	2	3	4	12	4	3	6	4				
C10	9	3	6	2	5	9	7	5	4	1095	7	5	0	5	5	1	2	4	4	5	5	7	1	6	2	5	7	2	2					
C11	11	9	10	20	18	9	7	15	9	3	525	12	8	3	5	1	3	7	2	2	2	5	2	6	2	1	7	2	1	6	3			
C12	4	7	6	0	2	5	3	16	13	9	8	427	6	5	1	2	0	4	2	5	2	1	5	2	5	3	6	2	5	3	3			
C13	5	4	5	9	7	10	4	5	5	5	7	2	405	2	2	5	3	1	3	4	4	5	2	9	6	8	10	2	3	1				
C14	10	7	4	18	8	2	8	7	4	6	9	9	1	877	4	2	2	7	2	5	5	2	4	7	6	6	2	2	6	2	2			
E02	36	10	8	1	5	4	8	25	8	10	3	8	11	5	543	MNHN	5	131	24	MNHN	8	MNHN	1	4	1	36	1	2	3	2	1	4	5	
E03	3	1	2	3	4	1	3	2	3	6	6	5	2	4	17	MNHN	253	MNHN	99	1	6	1	5	2	2	3	2	5	7	3	13	5	3	
E04	3	1	2	4	3	2	6	2	1	3	3	2	6	1	4	1	36	1	2	3	2	1	2	1	1	2	1	1	1	4	5	3		
E05	5	8	6	2	3	4	2	2	4	2	3	3	2	6	6	1	2	429	3	2	4	3	2	3	4	1	0	2	1	4	2	1	4	
E06	0	4	3	2	4	4	3	1	1	0	3	4	4	4	7	2	4	3	26	4	3	2	0	4	1	3	2	1	1	1	4	4		
S01	1	1	3	7	5	2	5	6	3	2	1	5	1	3	3	1	2	3	16	MNHN	16	MNHN	3	2	6	17	1	2	4	2	2	3		
S02	3	2	7	1	13	3	9	8	4	8	3	5	10	2	3	5	2	2	1	2	824	3	2	7	5	2	2	2	4	2	4	2	3	
S03	1	3	4	3	3	2	4	4	5	4	1	2	1	2	3	4	7	1	2	4	3	2	0	4	1	3	2	1	1	1	4	4		
S04	4	3	5	3	4	6	2	6	6	8	6	5	3	8	3	3	0	4	5	15	7	3	3	2	5	1	2	5	3	2	4	2		
S06	9	6	4	4	5	3	3	5	4	1	4	3	3	3	4	2	3	3	908	MNHN	908	MNHN	5	4	7	3	2	4	2	4	2	2		
M01	8	5	3	9	8	8	5	15	6	1	3	5	7	7	1	1	2	3	3	3	2	3	3	2	3	2	2	4	3	4	2	1		
M02	3	2	4	6	4	6	4	4	2	5	3	7	8	10	7	3	4	3	4	4	5	2	3	2	4	7	5	4	4	6	4	6		
M03	2	6	4	9	5	6	5	2	7	4	2	3	2	1	4	3	6	2	2	3	5	3	2	2	4	5	1	887	MNHN	1382	MNHN	5	4	4
M04	3	1	1	2	2	4	2	2	1	5	4	3	2	4	1	4	3	3	402	MNHN	402	MNHN	3	4	6	7	12	837	MNHN	402	MNHN	7	3	
M05	2	2	3	4	3	7	1	2	3	3	6	3	5	3	3	1	2	2	6	2	3	6	3	5	7	5	7	1046	MNHN	1235	MNHN	6		
M06	2	5	4	4	4	2	3	2	5	1	5	3	1	1	1	2	5	3	4	3	3	2	5	4	6	0	3	1	3	1	3	5	24	

OFF-SOURCE with 30 dB solar attenuator

Proj: 1308175		TST Observer:		HGR Title: INTERCOMPARISON TEST		Mod Mar 12 23:48:08 2003		RF -244 101-510 BFD-66MHz		Time: 23:48:04.66																							
Chan: 30	1000	Normalised	cf	Source: G147																													
C00	0	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15	16	17	18	19	20	21	22	23	24	25	26	27	28	29			
C01	3	3	3	3	3	3	3	3	3	3	3	3	3	3	3	3	3	3	3	3	3	3	3	3	3	3	3	3	3	3	3	3	
C02	10	6	1464	10	17	9	2	8	3	6	9	1	7	3	6	2	2	4	4	5	4	4	7	2	7	3	2	4	1	2	4		
C03	9	7	11	1266	20	11	8	3	5	6	6	2	8	4	2	2	2	2	3	6	3	3	6	10	3	2	3	3	2	3	2		
C04	5	4	20	10	1480	5	4	6	5	5	1	4	3	8	3	4	8	3	4	2	5	0	2	3	7	5	5	3	3	3	3		
C05	3	5	5	6	1691	10	5	6	10	2	0	7	1	4	3	1	3	1	2	6	4	1	3	3	4	4	4	5	3	1	1		
C06	7	9	9	8	12	5	370	6	7	6	2	1	6	3	3	2	4	1	8	4	5	1	1	8	2	4	4	1	6	9	9		
C07	0	2	8	3	6	3	4	1376	8	10	3	6	4	7	2	2	1	2	7	2	3	2	2	1	3	6	8	7	2	2	2		
C08	1	1	2	6	1	3	2	5	2	2	2	3	3	3	3	2	3	2	1	1	4	3	5	1	3	3	5	2	2	2	2		
C09	4	1	4	2	5	1	2	2	6	1367	2	0	6	1	2	3	2	0	2	2	2	2	1	4	2	8	5	4	8	1	4		
C10	4	1	4	2	5	1	2	2	6	1189	2	0	6	1	2	3	2	0	2	2	2	2	1	4	2	8	5	4	8	1	4		
C11	4	9	0	4	10	5	5	1	1	4	838	2	5	3	3	3	2	3	2	2	3	4	3	5	5	2	3	6	3	8	8		
C12	4	4	4	2	4	6	2	8	1	5	1	3	948	4	1	4	2	8	5	4	6	3	2	3	7	5	2	4	8	3	1		
C13	3	1	3	2	5	5	2	8	1	5	1	3	1145	4	1	4	2	8	5	4	6	3	2	3	7	5	2	4	3	3	1		
C14	2	2	7	3	8	1	1	8	8	5	7	4	2	1086	8	1	2	4	3	2	4	3	1	4	8	4	3	4	3	1	4		
E02	1	9	3	3	2	6	3	2	0	3	5	1	1	4	201	9	1	4	1	6	0	3	1	2	2	4	2	2	2	2	1	1	
E03	4	3	5	9	8	2	3	3	1	3	4	3	3	1	7	207	3	8	1	1	6	3	4	3	6	2	2	4	2	4	2	4	
E04	2	2	3	5	2	5	5	4	4	3	5	5	2	1	4	2	37	2	2	5	4	2	3	5	5	1	8	1	4	4	4		
E05	5	2	1	4	4	1	2	2	1	6	4	4	4	4	1	2	794	2	5	4	2	3	4	4	4	4	4	2	3	1	1	1	
E06	2	2	1	4	2	2	2	7	6	3	2	3	3	4	2	3	1	1093	2	5	4	2	3	5	3	5	3	5	3	1	1	1	
S01	1	4	7	0	3	2	6	6	1	1	1	5	2	3	2	6	4	0	2	1269	3	4	21	2	2	1	3	3	4	1	4	2	
S02	1	4	9	5	2	2	4	4	7	5	3	2	6	1	4	3	1	4	6	3	828	3	3	4	3	3	3	1	1	1	1	4	
S03	1	6	5	5	2	2	3	3	6	2	8	5	4	5	2	1	5	3	2	1419	5	1	1	3	2	5	2	5	2	3	3	4	
S04	7	2	4	2	5	3	1	2	1	1	7	4	2	4	1	2	3	3	2	344	4	2	2	2	2	3	3	3	3	3	4	4	
S05	5	2	2	7	3	6	2	4	5	3	4	2	2	5	6	2	3	3	6	1	2	4	3	1090	3	0	2	4	1	4	1	4	
M01	3	3	4	5	7	4	4	3	3	7	2	1	2	6	1	2	7	3	2	2	2	2	2	9	3	1396	4	4	3	1	2	2	
M02	2	7	3	1	5	4	2	1	3	4	2	3	7	4	4	3	7	4	3	7	2	3	1	3	4	5	2	1142	1	2	4	2	4
M03	5	3	2	4	5	2	0	8	2	5	2	4	2	5	4	1	3	6	1	4	0	4	2	6	4	4	942	4	4	2	4	2	
M04	3	3	2	5	3	4	4	5	7	3	2	5	1	1	3	5	1	1	4	3	2	2	1	1	4	3	3	3	3	3	3	3	
M05	2	4	3	1	6	4	2	4	4	8	1	4	8	2	4	3	4	4	1	1	1	1	1	1	1	2	1	7	8	1	2	1	2
M06	3	5	4	6	4	2	2	4	2	4	4	1	2	3	2	3	1	6	3	1	6	3	4	4	4	4	6	5	5	2	4	4	4

244-236 mkrz

noisyly ↓

ON-SOURCE with 30 dB solar attenuator

IF data ↓
#11
8 dB

Proj:	TEST Observer:	NGK Title:	INTERMODULATION TEST	Mod Har	12	23	41	21	2003																							
0	1308176	Charyo20	Source=3C147	RF=244	LO=310	EBL=66MHz	Time=23:41:14.63																									
000	001	002	003	004	005	006	007	008	009																							
000	001	002	003	004	005	006	007	008	009																							
C00	002	33	24	6	14	16	23	24	12	13	1	1.6	2	10	19	9	17	15	0	10	18	14	10	3								
C01	27	319	27	21	27	27	25	24	21	18	26	19	15	4	21	2	10	23	13	20	24	14	20	26	29	18	3					
C02	19	26	1432	18	22	21	19	26	13	23	18	12	20	22	13	5	20	3	16	19	12	23	24	13	20	19	22	4				
C03	16	28	19	1271	26	23	25	19	21	22	21	13	22	20	25	14	2	20	2	14	20	10	20	19	13	22	21	24	13	3		
C04	15	33	20	23	1420	20	24	21	27	18	16	18	20	19	13	2	23	2	14	21	14	27	24	10	22	24	21	16	1			
C05	15	20	17	48	18	2620	16	27	20	21	18	14	24	23	20	10	2	18	4	15	18	11	26	23	18	22	22	10	2			
C06	22	31	26	27	28	22	372	24	18	22	23	16	19	24	17	7	2	19	2	21	21	9	20	20	11	22	21	17	16	2		
C08	11	10	15	15	16	18	16	1425	12	25	22	9	19	23	15	12	2	21	2	14	17	12	19	18	12	19	21	22	16	2		
C09	18	25	20	17	24	12	17	14	1245	15	20	16	20	18	17	11	4	15	1	7	16	10	19	21	12	20	20	19	12	2		
C10	21	27	24	17	19	16	19	15	18	1421	19	14	20	20	16	15	1	19	4	17	20	10	21	18	13	16	20	19	18	3		
C11	21	30	25	21	26	21	30	15	22	26	830	12	23	20	17	6	3	17	4	15	17	9	16	17	16	20	19	18	14	3		
C12	18	17	16	17	22	19	21	18	14	16	15	1183	10	15	14	0	2	11	2	4	12	6	15	13	10	13	15	15	10	5		
C13	24	33	25	25	25	25	32	22	19	21	23	19	24	14	21	15	210	8	4	18	2	12	19	10	14	19	14	18	22	19	17	3
C14	15	23	24	19	20	19	21	18	18	22	22	16	25	1100	14	10	3	17	3	16	17	15	23	20	18	21	19	17	15	2		
E02	21	17	29	19	17	12	23	18	11	19	24	14	21	15	210	8	4	18	2	12	19	10	14	19	14	18	22	19	17	3		
E03	12	17	13	11	14	11	12	10	15	8	14	9	13	10	16	211	3	14	3	7	7	0	11	9	5	12	9	12	7	3		
E04	2	2	1	4	3	5	3	3	2	0	5	2	2	4	2	4	37	2	3	2	3	4	1	2	5	2	2	4	5	2	2	
E05	18	27	23	23	24	13	24	17	17	20	19	16	22	20	20	10	6	897	3	13	19	6	22	20	10	20	22	20	15	2		
E06	1	3	1	3	1	6	4	3	4	2	2	5	2	4	4	4	2	1	20	5	4	1	2	4	2	2	0	5	3	2		
S01	6	10	5	12	10	5	11	8	7	7	11	9	8	6	8	4	2	9	4	1278	10	9	24	8	6	19	13	15	8	2		
S02	19	22	20	18	17	8	18	13	18	20	24	16	21	14	19	7	5	19	2	1546	3	1390	11	21	21	16	25	18	24	12	2	
S03	7	17	13	10	14	12	15	12	9	13	11	14	13	10	10	8	3	6	10	1307	12	11	4	8	5	12	10	6	5	0		
S04	15	25	24	23	19	21	21	14	23	21	24	14	25	20	22	13	3	18	2	11	25	14	819	23	45	22	19	24	13	5		
S06	16	24	22	19	20	25	24	19	21	19	28	17	25	21	22	15	1	21	3	7	20	10	22	1029	11	24	22	20	15	2		
M01	14	19	13	16	19	15	18	12	11	14	19	12	16	15	18	6	2	17	3	11	16	4	12	18	1360	10	11	17	6	0		
M02	22	30	27	27	22	22	31	13	22	22	20	22	18	18	12	4	22	4	11	22	11	22	26	22	1119	23	24	13	4			
M03	12	23	21	22	27	17	26	13	19	13	22	19	26	22	17	9	7	20	4	9	16	11	20	22	15	20	201	27	18	6		
M04	15	30	22	16	26	18	22	18	22	20	26	18	23	22	21	15	5	19	4	7	18	9	22	28	13	18	24	1471	12	2		
M05	7	16	19	16	11	11	11	20	9	13	11	16	9	15	8	10	9	3	12	0	2	15	11	17	10	4	14	13	1377	1228	4	
M06	1	3	5	2	1	3	4	3	1	4	2	4	4	4	2	5	1	5	3	4	2	0	2	1	2	2	2	4	2	2	5	24

244-13 K MHz

[259-251]

GN-source

Proj: T1326176		TST Observer: 1		MCK Title: INTERMODULATION TEST		Med Mar 12 22:09:36 2003	
Chan	Coef	Coef	Coef	Coef	Coef	Coef	Coef
0	0	0	0	0	0	0	0
C00	145	C00	145	C00	145	C00	145
C01	60	C01	60	C01	60	C01	60
C02	34	C02	34	C02	34	C02	34
C03	33	C03	33	C03	33	C03	33
C04	41	C04	41	C04	41	C04	41
C05	31	C05	31	C05	31	C05	31
C06	29	C06	29	C06	29	C06	29
C07	39	C07	39	C07	39	C07	39
C08	41	C08	41	C08	41	C08	41
C09	41	C09	41	C09	41	C09	41
C10	40	C10	40	C10	40	C10	40
C11	39	C11	39	C11	39	C11	39
C12	29	C12	29	C12	29	C12	29
C13	41	C13	41	C13	41	C13	41
C14	30	C14	30	C14	30	C14	30
E02	21	E02	21	E02	21	E02	21
E03	15	E03	15	E03	15	E03	15
E04	1	E04	1	E04	1	E04	1
E05	39	E05	39	E05	39	E05	39
E06	1	E06	1	E06	1	E06	1
S01	17	S01	17	S01	17	S01	17
S02	32	S02	32	S02	32	S02	32
S03	4	S03	4	S03	4	S03	4
S04	31	S04	31	S04	31	S04	31
S05	29	S05	29	S05	29	S05	29
M01	34	M01	34	M01	34	M01	34
M02	40	M02	40	M02	40	M02	40
M03	41	M03	41	M03	41	M03	41
M04	43	M04	43	M04	43	M04	43
M05	40	M05	40	M05	40	M05	40
M06	5	M06	5	M06	5	M06	5

sevoir
~3dB

MITHRAS AMP/CON

Chan	Coef	Coef	Coef	Coef	Coef	Coef	Coef
0	0	0	0	0	0	0	0
C00	145	C00	145	C00	145	C00	145
C01	60	C01	60	C01	60	C01	60
C02	34	C02	34	C02	34	C02	34
C03	33	C03	33	C03	33	C03	33
C04	41	C04	41	C04	41	C04	41
C05	31	C05	31	C05	31	C05	31
C06	29	C06	29	C06	29	C06	29
C07	39	C07	39	C07	39	C07	39
C08	41	C08	41	C08	41	C08	41
C09	41	C09	41	C09	41	C09	41
C10	40	C10	40	C10	40	C10	40
C11	39	C11	39	C11	39	C11	39
C12	29	C12	29	C12	29	C12	29
C13	41	C13	41	C13	41	C13	41
C14	30	C14	30	C14	30	C14	30
E02	21	E02	21	E02	21	E02	21
E03	15	E03	15	E03	15	E03	15
E04	1	E04	1	E04	1	E04	1
E05	39	E05	39	E05	39	E05	39
E06	1	E06	1	E06	1	E06	1
S01	17	S01	17	S01	17	S01	17
S02	32	S02	32	S02	32	S02	32
S03	4	S03	4	S03	4	S03	4
S04	31	S04	31	S04	31	S04	31
S05	29	S05	29	S05	29	S05	29
M01	34	M01	34	M01	34	M01	34
M02	40	M02	40	M02	40	M02	40
M03	41	M03	41	M03	41	M03	41
M04	43	M04	43	M04	43	M04	43
M05	40	M05	40	M05	40	M05	40
M06	5	M06	5	M06	5	M06	5

LOI = 295 MHz

229-221 MHz

off source

Proj:	TST	Observer:	NGX	Tilt	Hz	Modulation	TEST	band	Time	21	22	23	24	25	26	27	28	29															
0	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15	16	17	18	19														
000	001	002	003	004	005	006	007	008	009	010	011	012	013	014	015	016	017	018	019														
C00	402	109	32	69	26	39	37	70	74	22	26	81	24	18	68	6	5	3	2	72	34	2	12	6	14	3	1	2	5	3			
C01	14	299	47	75	24	41	79	71	401	52	31	55	43	24	66	10	3	7	2	83	53	2	23	0	10	5	7	4	5	2			
C02	8	1290	38	19	26	22	33	54	31	9	43	18	6	15	5	1	1	8	41	23	5	10	4	17	2	4	3	2	1				
C03	16	12	31	378	51	34	60	63	95	36	31	76	26	13	20	15	3	2	1	82	50	5	21	12	18	5	5	4	5	2			
C04	6	3	4	28	1072	25	26	22	32	5	3	28	21	3	8	7	5	4	1	35	14	2	8	5	3	6	4	6	3	4			
C05	10	7	11	19	17	1784	21	19	45	17	8	82	17	8	20	7	3	4	3	39	22	2	6	3	10	4	3	7	3	2			
C06	15	20	5	18	11	9	325	33	65	35	26	48	21	6	26	6	3	3	3	55	28	3	8	4	11	3	6	2	4	2			
C08	45	51	18	12	11	27	66	1176	52	19	24	87	27	8	37	1	4	3	2	50	30	3	17	2	12	3	3	1	2	2			
C09	1	9	11	33	16	22	21	15	1513	20	34	66	39	4	38	3	1	4	3	94	48	1	22	5	25	4	5	2	1	5			
C10	5	17	5	3	11	9	18	18	6	1118	12	35	11	6	13	12	2	6	2	34	24	3	13	2	10	5	6	2	1	3			
C11	14	6	2	17	10	5	4	17	9	11	778	11	12	13	2	5	2	2	2	14	13	4	5	3	7	4	4	4	5	1			
C12	9	14	10	31	14	25	20	49	14	6	13	426	28	8	38	9	3	5	2	82	42	3	15	4	20	5	4	4	5	7			
C13	2	14	4	11	7	12	14	33	9	3	6	14	397	4	7	9	3	2	5	26	27	1	8	4	10	5	8	7	1	3			
C14	1	15	7	12	9	13	11	22	3	11	13	3	11	1225	3	2	2	2	2	15	3	1	3	5	4	3	2	3	4				
E02	28	23	1	9	13	22	27	53	18	7	9	22	9	1	248	94	6	10	5	35	5	2	8	7	3	2	4	1	6	3			
E03	1	9	5	6	6	9	16	21	6	3	5	7	4	4	170	101	4	12	1	8	14	2	9	3	7	5	2	8	2	2			
E04	3	3	3	2	4	5	1	6	3	2	1	3	2	5	4	4	38	1	1	3	1	2	4	2	5	2	3	10	3	3			
E05	3	4	3	2	3	1	5	7	3	4	6	4	2	6	4	7	5	206	3	3	1	2	3	3	2	5	4	9	6	1			
E06	2	2	3	6	2	2	2	3	2	3	5	5	3	4	1	4	1	5	25	2	4	1	0	3	2	1	6	2	1	3			
E08	9	12	1	18	14	11	23	45	13	13	4	20	11	2	22	6	1	2	2	497	44	3	13	4	14	6	1	3	3	2			
S02	11	11	6	17	20	16	31	54	15	9	3	21	16	3	6	25	0	2	1	24	736	1	26	11	16	1	5	4	1	2			
S03	3	2	3	4	2	6	3	2	3	1	2	2	4	1	1	3	4	2	1	3	36	3	1	2	1	4	6	5	3				
S04	1	4	1	9	5	4	8	12	12	5	1	4	4	1	0	13	2	3	3	8	14	2	709	8	9	1	6	3	2	7			
S06	3	4	2	1	3	4	4	4	4	4	2	0	4	4	3	2	3	2	4	2	3	0	5	0	5	2	7	4	6	2	4		
M01	5	5	5	6	6	4	8	7	5	7	5	4	5	5	0	4	3	2	2	4	4	3	3	3	3	3	4	6	5	3	1		
M02	3	5	5	4	6	1	0	3	2	3	2	3	6	6	8	3	4	4	4	4	5	2	5	1	5	12	2	6	2	6	2		
M03	7	5	6	11	11	1	3	2	3	5	10	4	3	2	2	2	4	3	4	4	4	4	4	2	2	2	20	1202	6	5	1		
M04	2	2	5	3	1	0	2	2	2	2	3	4	3	3	3	7	9	5	3	3	4	4	2	11	5	2	8	1547	24	4			
M05	3	1	4	3	4	3	2	6	4	3	0	6	2	0	4	3	1	2	3	2	4	2	4	5	4	6	7	65	1001	2			
M06	5	2	2	4	4	5	3	6	2	4	5	2	2	0	2	1	5	4	2	1	3	2	2	0	4	4	4	2	1	2	5		



