# Wideband Radio Frequency Interference in the 233 MHz band 30 July 2003

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# 1 Background

Wideband radio frequency interference (RFI) has been observed in the 233 MHz band as shown in this online display of the visibility amplitudes while observing an off-source position (Fig 1). The top half above the diagonal shows the 130 MHz polarization whereas the bottom half shows the 175 MHz polarisation. The diagonal shows the self counts for both the polarizations. This is an instantaneous display for a given spectral channel. Correlated signal is picked up by the central square antennas and the nearest arm antennas (e.g W01, S01, E02) even when the antennas are pointing at an off-source position where no correlated signal is expected. The correlated signal simulates a source of strength  $\sim 20-30$  Jy and is present in all the spectral channels unlike narrow-band RFI. Such wide band RFI reduces the dynamic range of the observations by increasing the rms noise in the image and hence reducing the sensitivity of the observations. This is not desirable for observations and it is necessary to determine its origin and eliminate it. We undertook a few experiments, which we describe here alongwith the results, to understand the origin of this RFI and ways to live with it till it is eliminated.

First some history. Such correlated wideband RFI was observed by the observatory about two years ago, but it was certainly absent about six months ago. Some of the remedial measures taken two years ago were:

- Sparking/arcing in power lines were checked for. Apparently a culprit was found near W01. It was not clear if the wideband RFI disappeared.
- Some of the power line insulators were replaced. It was not clear if the wideband RFI disappeared.
- The wideband RFI clearly disappeared after the rains then!

Some of the suspects this time could be:

- sparking/arcing in power lines
- the wideband (8 MHz) TV signals getting intermodulated into the observing band.
- unknown!

We conducted a few experiments in March 2003. The first experiment concentrated on determining the origin of this RFI. The second and third experiments concentrated on finding an appropriate observing method to co-exist with the RFI till it was eliminated. In this memo, we describe the experiments and the results we obtained. In summary, we did not obtain a conclusive result as to the origin of the RFI; however we found that by changing the observing band or introducing high attenuation in the default band, one can reduce the effect of this RFI on observations.

## 2 Experiment 1

Fig 1 shows how the wideband RFI will be manifested while observing an off-source position. The shorter baselines show finite correlation while the longer baselines where the RFI is uncorrelated show noise. Suspecting the mains power lines to be responsible for the RFI, we decided to power down the central square antennas + arm antennas and run the central square on generator supply. We sequentially turned off the power to the south, west and east arms and noted the amplitude change in the correlated signal. All this work was done towards an off-source position so as to avoid confusion with the source signal. The experiment was done by using default RF settings for the 233 MHz band i.e. the IF band 244-236 MHz. We also made sure that to examine channels which were free of narrow-band RFI. The various steps taken in the experiment with the results are listed below:

- 1. The power to the south arm was switched off and a drop in the correlated signal was noticed. But the signal did not disappear. On switching off power supply to the east and west arms, no change in the RFI was noticed.
- 2. Then we switched off the power supply to the analog, electronics and baseband laboratories. We suspect that a few of the spiky narrow-band RFI which we had noticed, but unfortunately had not kept detailed records, had subsided when the analog laboratory was switched off. However, the experiment has to be repeated before any conclusions can be drawn. There was no change in the wideband correlated signal.
- 3. The power supply to the laboratories was turned on. No change in the correlated signal was noticed.
- 4. The power supply to the arm antennas was switched on. No change in the correlated signal was noticed. It is not clear whether the drop in the correlated signal noticed in (1) was related to powering down the south arm or was coincidental.
- 5. The power supply to central square was switched off and the diesel generator was switched on. No change in the correlated signal was noticed.
- 6. Power was restored on the central square. No change in the correlated signal.

The only change we noted was in the strength of the correlated signal which was 30 counts when we started and dropped to 10-15 counts when we ended. We have no explanation for this. It could either be that the correlated signal was variable during the experiment or else some of the power lines which contribute to the correlated signal are in the south arm and hence the RFI correlation reduced when the south arm antennas were powered down.

# 3 Experiment 2

In this experiment, we varied the frequency of the first local oscillator and scanned through the 233 MHz RF band to check for a 'clean' band which is free from the wideband RFI. One of the possible origins of this wideband RFI is cited as the TV signal which can get get intermodulated into our observing band. The Pune TV signal is between 175 and 183 MHz. Junnar TV receives this signal and retransmits between 195 and 203 MHz. If these signals which lie outside the 233 MHz RF band beat with a narrow-band RFI, there is a possibility of them getting translated to within the 233 MHz RF band and we would observe it as a wideband RFI. However, if this is the case then the wideband RFI should have a frequency dependence depending on the translating narrow-band RFI which is beating with the TV signal. This experiment was aimed at investigating this issue. We observed an off-source position for all the frequency settings and checked the magnitude of the wideband correlated signal. Please note that the default settings for 233 MHz observations (244-236 MHz) are: LOI=310 MHz, IFBW=6 MHz, LOIV=66 MHz, BBW=8 MHz. Fig 1 shows the counts registered at an off-source position. Notice the higher counts on the central square antennas which is the signature of the wideband RFI which we are trying to remedy. The LOI frequency was changed whereas the LOIV frequency was held fixed to 66 MHz for all the following cases. The strength of the correlated RFI signal was compared to the strength obtained for the default RF settings.

- 1. Changed LOI to 300 MHz so that the observing band was 234-226 MHz. The correlated RFI signal was stronger.
- 2. Changed LOI=295 MHz so observing band was 229-221 MHz. The correlated RFI signal was stronger (See Fig 2). However note that the on-source sensitivity is down by 5 dB.
- Changed LOI=320 MHz so observing band was 254-246 MHz. The correlated RFI signal was weaker (See Fig 3).
- 4. Changed LOI=325 MHz, so observing band was 259-251 MHz. The correlated RFI signal was weakest in this case (See Fig 4). The on-source sensitivity was down by 5 dB.

Summarizing, it appears that the wideband RFI is weakest in the 251 to 259 MHz band and strongest in the 221 - 229 MHz band. In other words, the wideband

RFI gets weaker for frequencies from  $220 \rightarrow 250$  MHz. (Subhashis Roy conducted a similar experiment the next day and confirmed the above results.) Thus observing in the band 251 - 259 MHz seems better for reducing the effect of the wideband RFI on the data. However, our experiment does not give any information on the presence of narrow band RFI in this band. If the narrow band RFI in this band turns out to be severe then it might be a good idea to adhere to the default settings and resort to other measures. We conducted the next experiment to check out if the default observing band at 233 MHz can be used with reduced sensitivity and reduced contamination by the wideband RFI.

## 4 Experiment 3

Since the system temperature at 233 MHz is high (stronger background non-thermal radiation field) and RFI can be strong which can saturate the system; the normal observing procedure involves introducing a 14 dB solar attenuator in the front-end. At GMRT, a facility is provided for inserting a large attenuator in the front-end. This attenuation can either be 14 dB or 30 dB and are known as solar attenuators. In this experiment, a solar attenuator of 30dB was introduced in the path for the default 233 MHz settings (LOI=310MHz, LOIV=66MHz). 3C147 was observed. The correlation with the default 14dB solar attenuator of 0.07-0.08 (see Fig 5) dropped to 0.02 (see Fig 6) that is a drop of factor  $\sim 4$  ie  $\sim 6$  dB. This experiment was conducted to check if the 30 dB solar attenuator would be sufficient to kill the wideband RFI without losing sensitivity.

The off-source position with a 30 dB attenuation in the front-end (see Fig 7) shows weak signatures of the wideband correlated signal; the cross correlation of the RFI is lower by a factor 4 like the on-source correlation. Thus the results of this experiment and the previous experiment are similar. This experiment demonstrates that the default RF band can be used for observations with a 30 dB attenuator in the front end and this reduces the effect of the wideband RFI. As in the previous experiment, this does not cure the wideband RFI problem.

### 5 Summary

The origin of the wideband RFI in the 233 MHz band is not clear. Since the weather is fairly dry, it is likely that the cause of the RFI is in some sparking/arcing of power lines. However our control experiment to track this down was not conclusive. We, then, turned our attention to finding ways of observing in the presence of this wideband RFI and arriving at a compromise between reduced sensitivity and reduced wideband RFI. We find that the wideband RFI reduces in strength when scanning in RF from 221 to 259 MHz. The wideband RFI is strongest near 221 MHz and weakest near 259 MHz. There is a drop in sensitivity of 5dB near 221 and 259 MHz. In another experiment, we find that the wideband RFI can be reduced by introducing a solar attenuator of 30 dB in the front-end. This results in a drop of 6 dB in the sensitivity. However, in both these cases, we find that the wideband RFI also becomes weaker.

Further investigations are required for pinning down the source of this RFI. Some further experiments that can be conducted are:

- Check for the wideband RFI when power fails and the TV signals are absent.
- Repeat the experiment of running the central square on the diesel generator and the effect on the wideband correlated signal.
- Check if rains cure the problem!

(**P.S.**: Although this memo was prepared soon after the above experiments were conducted, its submission has got delayed beyond the onset of monsoon, giving us an opportunity to examine the effect of rains on the wideband RFI. Unfortunately this time, we find that the wideband RFI persists and the rains have failed to provide the solution. More experiments are being planned and conducted to trace down the source of this RFI.)

Note: The figures are in another file (figs.ps). The captions for the figures are given below as also the page number in the figures file.

#### Figure Captions:

Fig 1: This figure shows the counts registered at an off-source position for default frequency settings in the 240 MHz band. Notice that while most of the arm antennas show a few counts which is noise; the central square antennas and the nearby arm antennas show few tens of counts which is the correlated signal, the source of which we are trying to locate. (page 4 or 5)

Fig 2: Expt 2(b): Off-source counts for LOI=295 MHz; (page 8).

Fig 3: Expt 2(c): Off-source counts for LOI=320 MHz (last page in this file).

Fig 4: Expt 2(d): Off-source counts for LOI=325 MHz; (page 7).

**Fig 5:** Expt 3: On-source (3C147) counts with standard 14dB solar attenuator; (page 1).

Fig 6: Expt 3: On-source (3C147) counts with 30dB solar attenuator; (page 3).

Fig 7: Expt 3: Off-source counts with 30dB solar attenuator; (page 2).

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